Multi-frequency VLBI studies of the OVV quasar NRAO 530

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Introduction



A well know OVV:

- z = 0.902, 1 mas \rightarrow 7.8 pc
- erratic and strong broad-band variability
- $\beta_{app} \approx 10-40 \text{ c}$

<u>Morphology:</u> <u>kpc scale:</u>

a core jet structure in P.A. -50°, double lobes in the E-W direction

pc scale:

oscillating jet consists of a number of emission components north to the core

Data collection

• Main calibrator in 2007 Sgr A* observing campaign:



See my poster for other results from this campaign

• Mojave 15 GHz (1999-2009, 17 epochs)

component identification: an example



Spectra and spectral evolution





Core separation

- Core identification: the compact component at the south end
- $B \approx 76 \cdot \delta$ mG based on SSA
- systematic change of α along the jet

v-dependent positions of components



Core location r_{core} varies with v: $r_{core} \propto v^{-1/kr}$ (K_r is related to the electron energy distribution, B-field, and the electron number density)

Synchrotron self-absorption, $K_r = 1$

v-dependent positions of components



Inter-day Variability: flux density



- The probability for variability for most components is low
- m<4 % for the core, and m< 20% for the jet

Jet kinematics at 15 GHz



Core separation vs. time

Jet kinematics at 15 GHz



P.A. swing for comp. *d*, *e*, *h*, and *i*

€

P.A. keeps nearly constant for comp. *f*, *g*, *j*

P.A. vs. time

Morphology evolution: jet wobbling



- Visible only in the innermost regions
- Similar to many others, like NRAO150, BL Lac etc.





One-sided core-jet structure with spectral evolution

2-D position shifts along the jet, some of which are probably due to the core-shift

Variations of flux density and structure on daily time-scales: flux density: m < 4 % for the core and m < 20 % for the jet; jet speeds < 170 c</p>



evolving jet ridge line, a consequence of 3-D motion

Thank you !