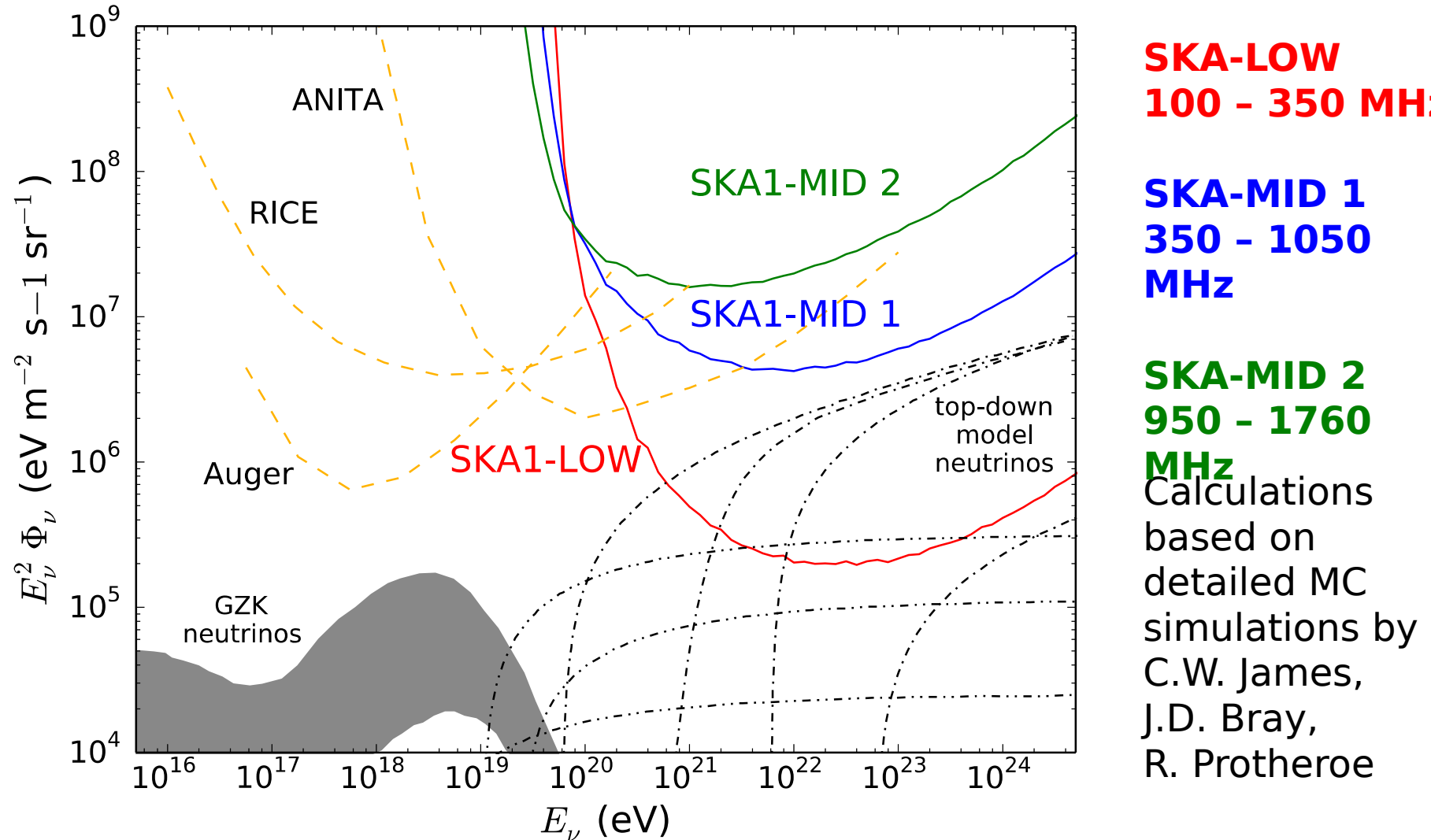


The Lunar **ASKA**ryan project

Neutrino interactions & the Landau-Pomeranchuk-Migdal (LPM) effect

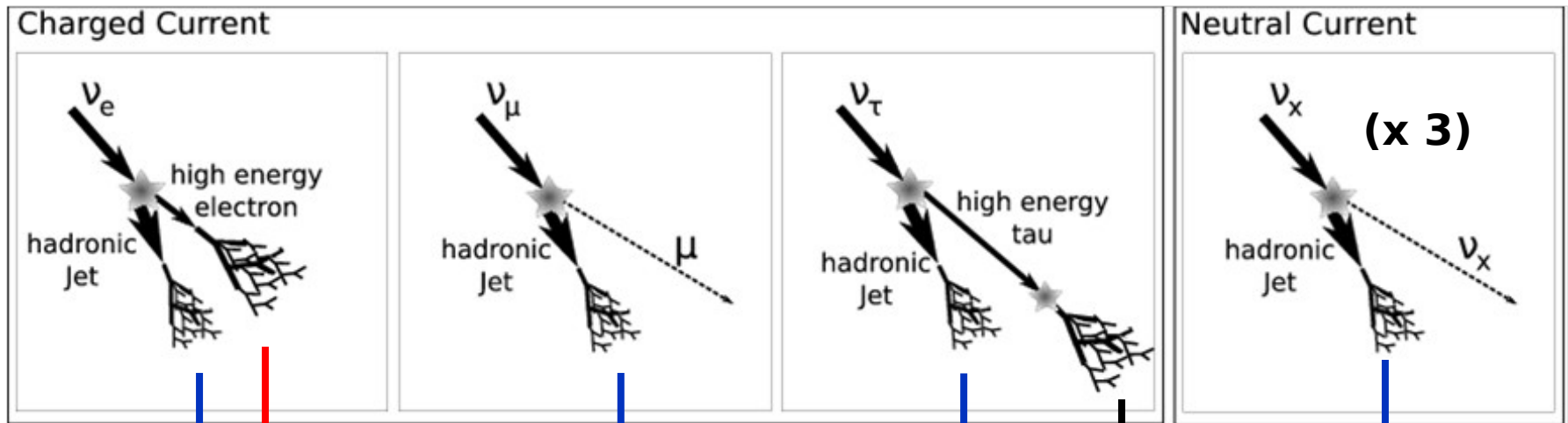
Jaime Alvarez-Muñiz
Dept. Física de Partículas
Universidade Santiago de Compostela
Spain

SKA sensitivity to Extremely High-Energy neutrinos



Neutrino – induced showers

Neutrino – nucleon Deep Inelastic Scattering (DIS)



“Mixed” showers

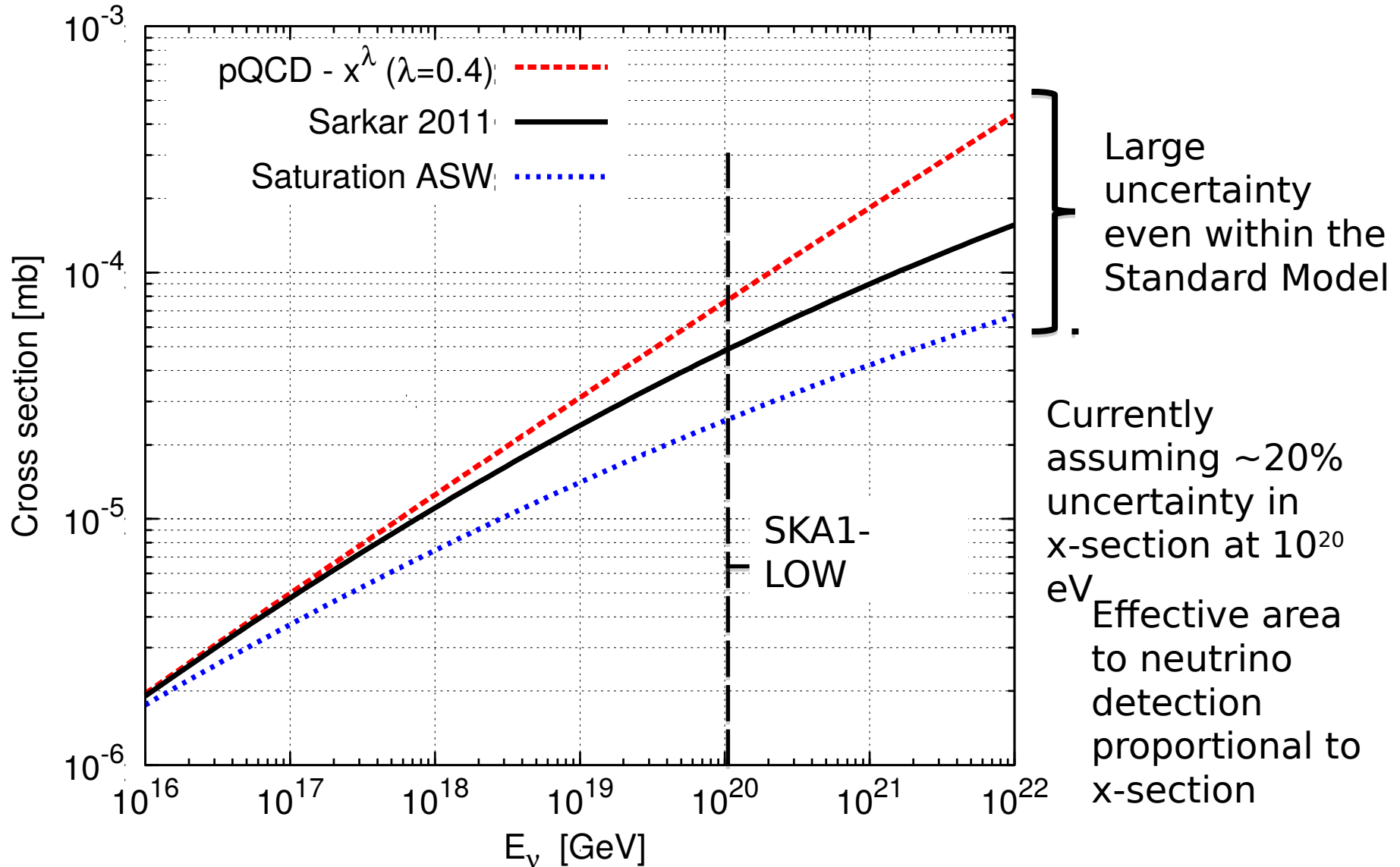
$E_{\text{electromagnetic}} \approx 80\% E_\nu @ \text{EeV}$

$E_{\text{hadronic}} \approx 20\% E_\nu @ \text{EeV}$

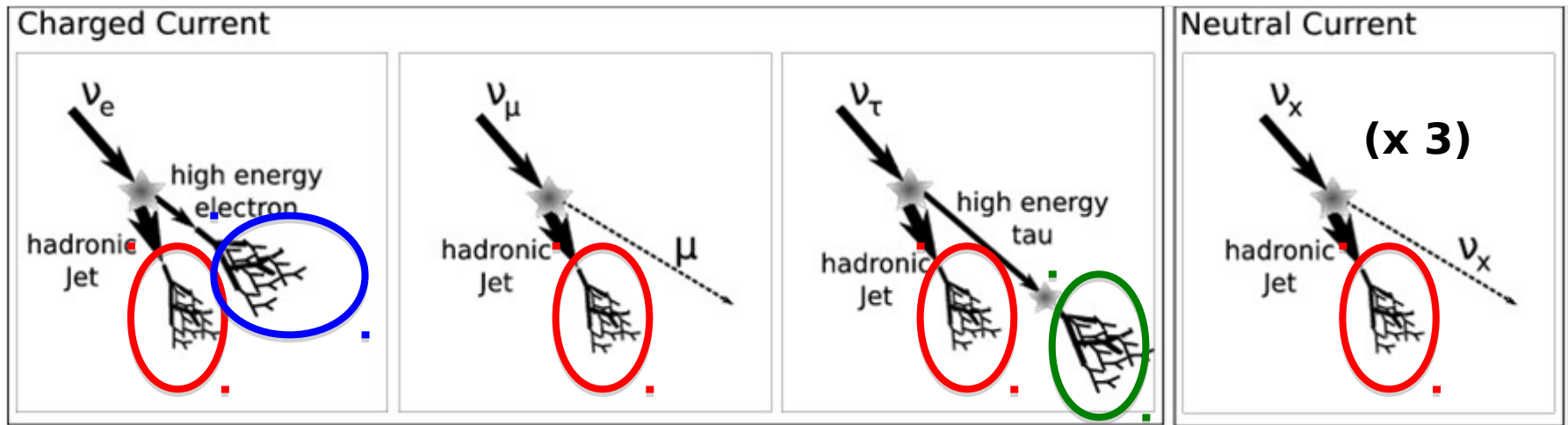
Hadronic showers
 $E \approx 20\% E_\nu @ \text{EeV}$

Electromagnetic
or
Hadronic

ν - nucleon CC cross-section



Relative contribution to aperture in SKA-LOW at $E_\nu = 10^{21}$ eV

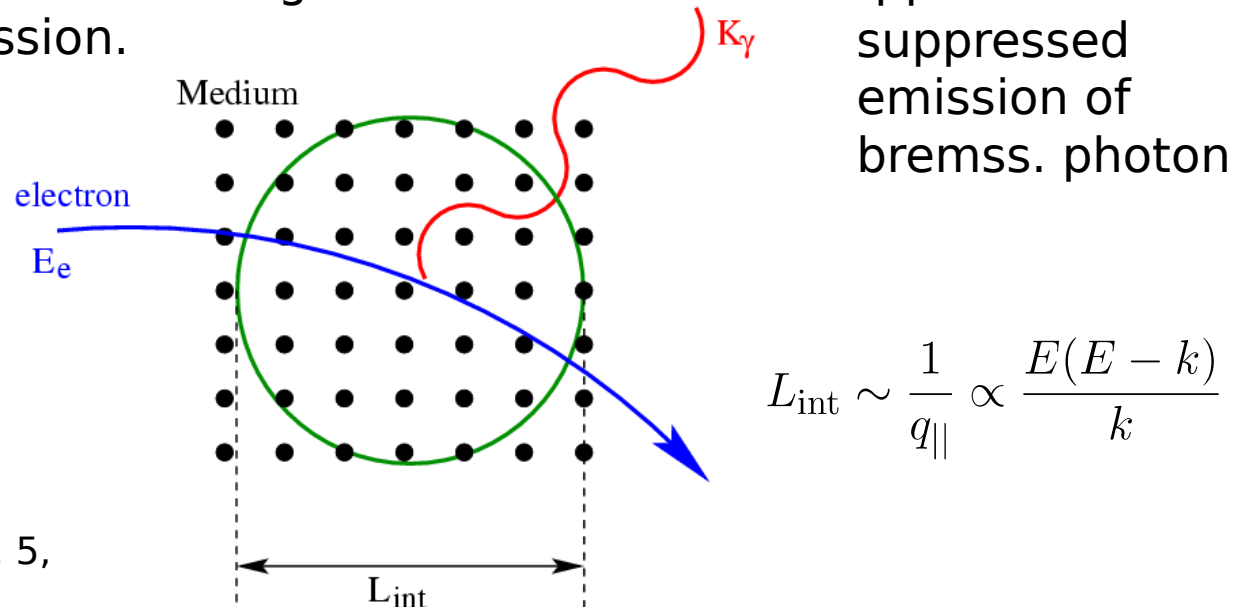


Type of shower/interaction	Relative contribution
Primary hadronic ν_μ & ν_τ CC & all ν NC	$\sim 93\%$ 
Primary electromagnetic ν_e CC	$\sim 1\%$
Tau photonuclear ν_τ CC	$\sim 5\%$
Other (muon, tau brems, etc.)	$\sim 1\%$

The Landau-Pomeranchuk-Migdal (LPM) effect

- Quantum mechanical effect that suppresses:
 - bremsstrahlung by electrons or positrons
 - pair production by photons
- Interactions do not occur at a point, but along a distance (formation length) $L_{\text{int}} \sim 1/q_{\parallel}$ ($q_{\parallel}$ = parallel momentum transfer) that increases with Energy
- Typically if $E > E_{\text{LPM}} \sim 10^{15}$ eV then $L_{\text{int}} \gg \text{interatomic distance} \Rightarrow$ interactions along formation length scatter electron & suppress bremsstrahlung emission.

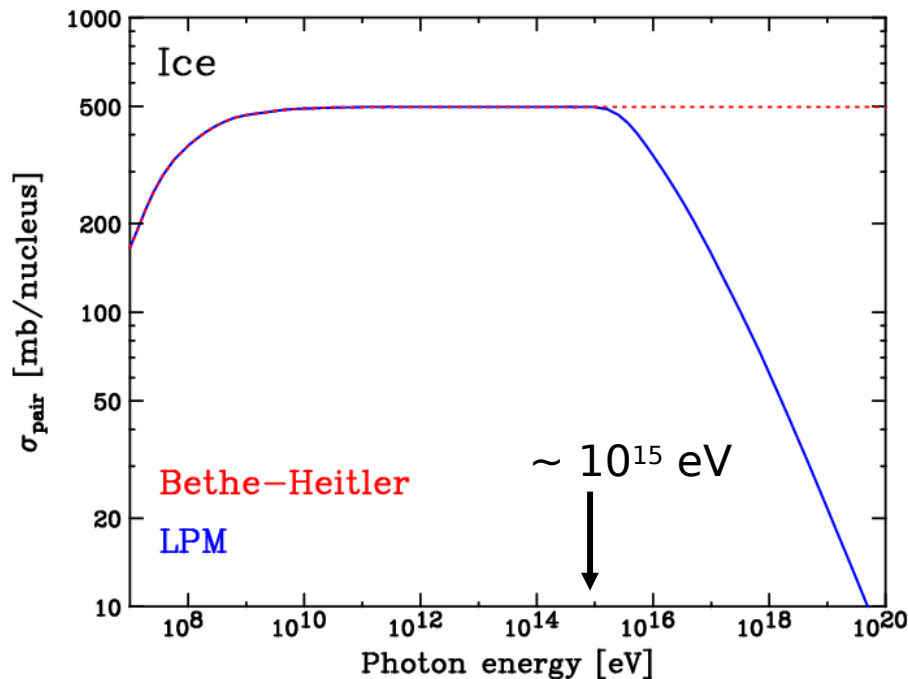
L. Landau and I. Pomeranchuk, Dokl. Akad. Nauk SSSR **92**, 535, 1953
A.B. Migdal, Phys. Rev. **103**, 1811, 1956



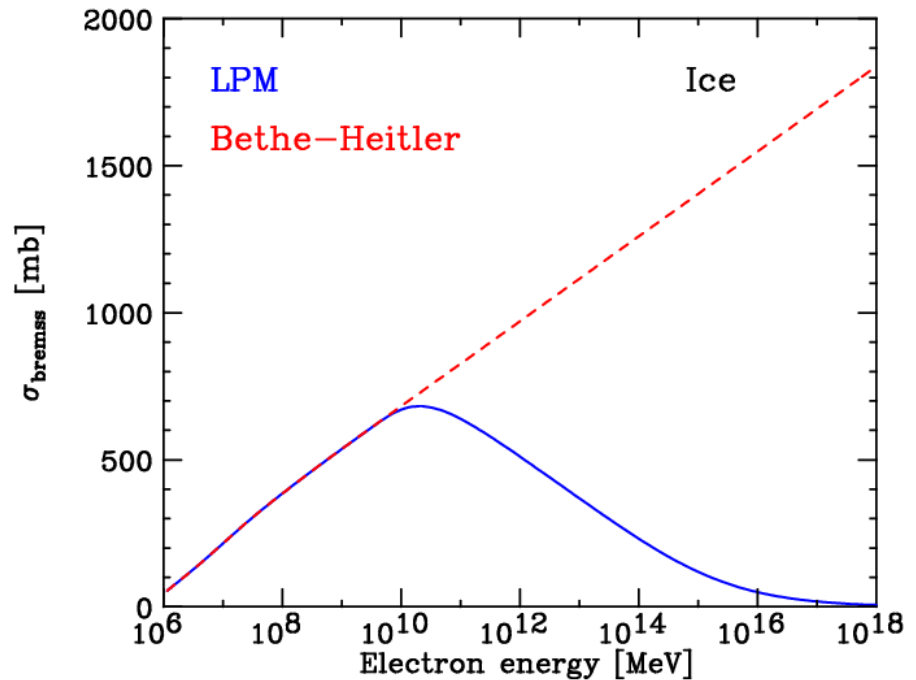
$$L_{\text{int}} \sim \frac{1}{q_{\parallel}} \propto \frac{E(E - k)}{k}$$

LPM effect: Cross sections

Pair production by photons



Bremsstrahlung by electrons



- Significant reduction of pair-production and bremsstrahlung cross sections with respect to Bethe-Heitler predictions
- Also suppression of “soft” bremsstrahlung

LPM effect & showers

Electromagnetic showers

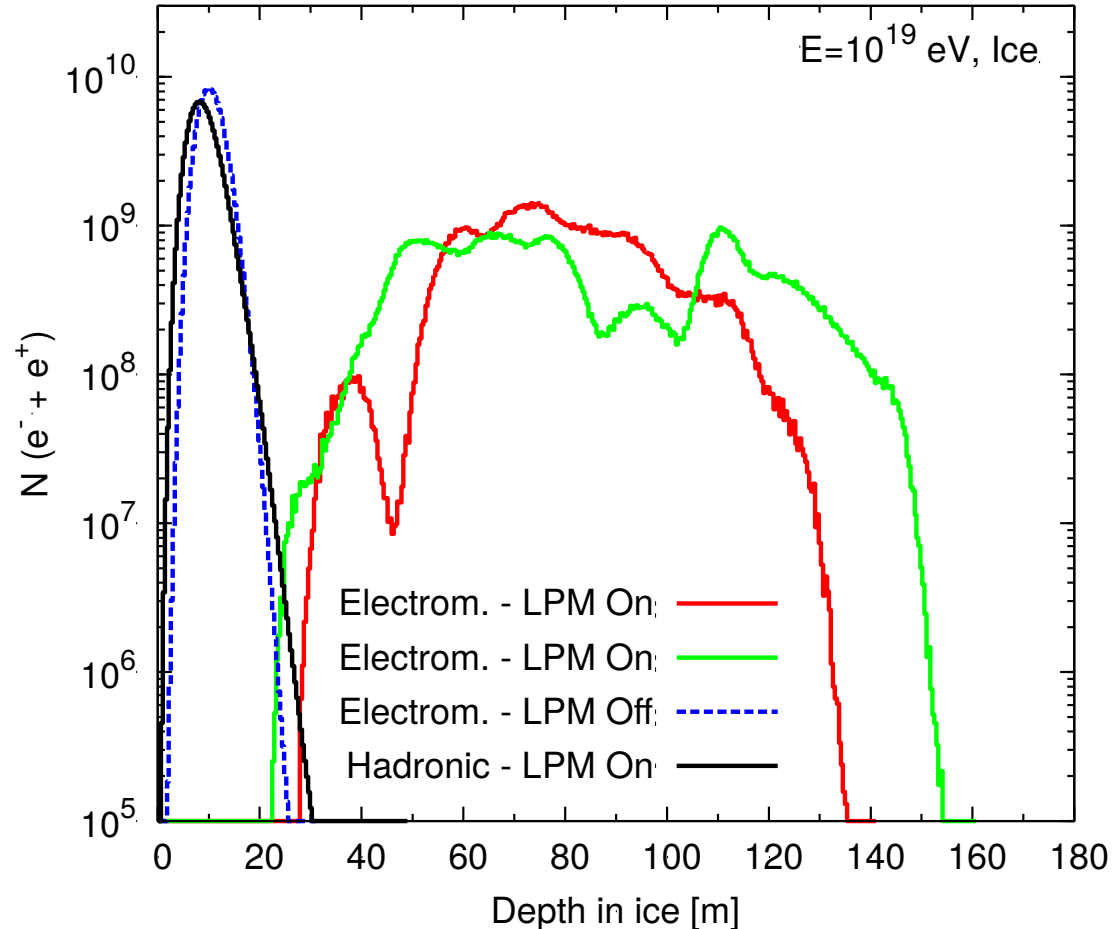
having $E > E_{\text{LPM}}$

($\sim 10^{15}$ eV regolith):

- strongly affected by LPM effect
- more penetrating than those with $E < E_{\text{LPM}}$.
- Width of long. profile (shower length L) increases faster than $\sim \log E$, typically as $E^{1/3}$
- Energy deposition is clumpy

Hadronic showers

- much less affected by LPM effect

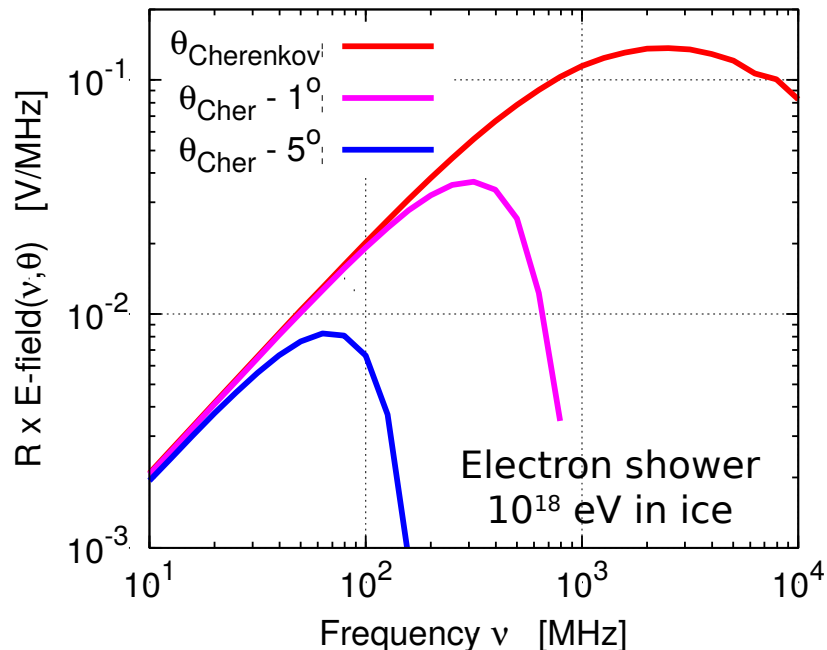
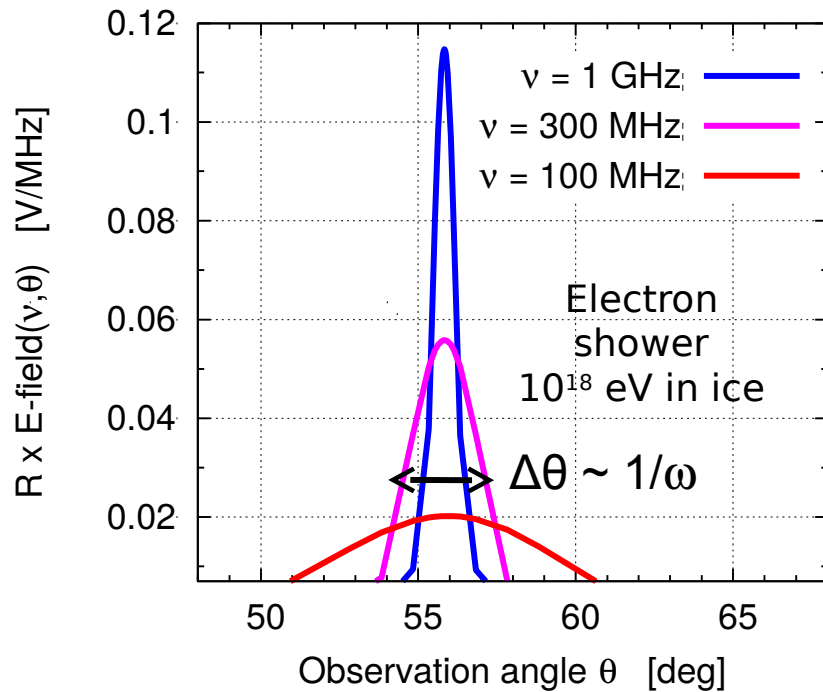


J.A-M & E.Zas, Phys. Lett. B **411** (1997) 218

J.A-M & E.Zas, Phys. Lett. B **434** (1998) 396

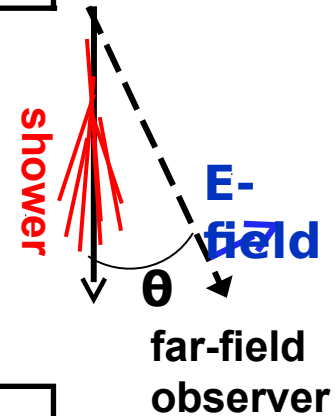
J.A-M, R.A. Vazquez & E.Zas, Phys. Rev. D **61** (1999)

Properties of signal (freq.-domain)



Angular distribution

- Radiation emitted in a cone around Cherenkov angle.
- Width proportional to inverse of frequency.



Frequency spectrum

- Broad-band frequency spectrum.
- Field increases with frequency up to a cut-off depending on observation angle.

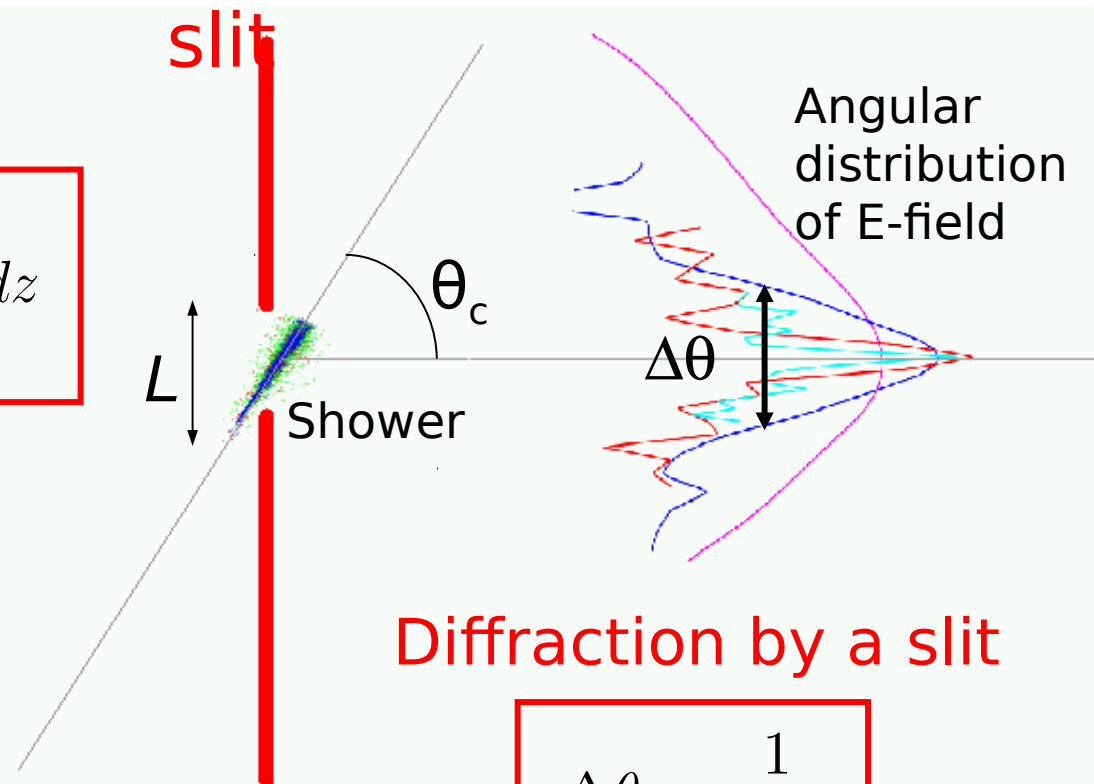
E. Zas, F. Halzen, T. Stanev, PRD 45, 162 (1992),
 J. A-M & E. Zas, PLB 411, 218 (1997)
 J.A-M, C.W. James, R.J. Protheroe, E. Zas, APP 32, 100 (2009)

Optics analogy to understand radio emission: diffraction by a single slit

- The shower plays the role of the **slit** in elementary optics.
- Electric-field can be obtained **Fourier-transforming** the longitudinal profile $Q(z)$ of excess charge (Fraunhofer diffraction)

$$E(\omega, \theta) \propto \omega \int Q(z) e^{ipz} dz$$

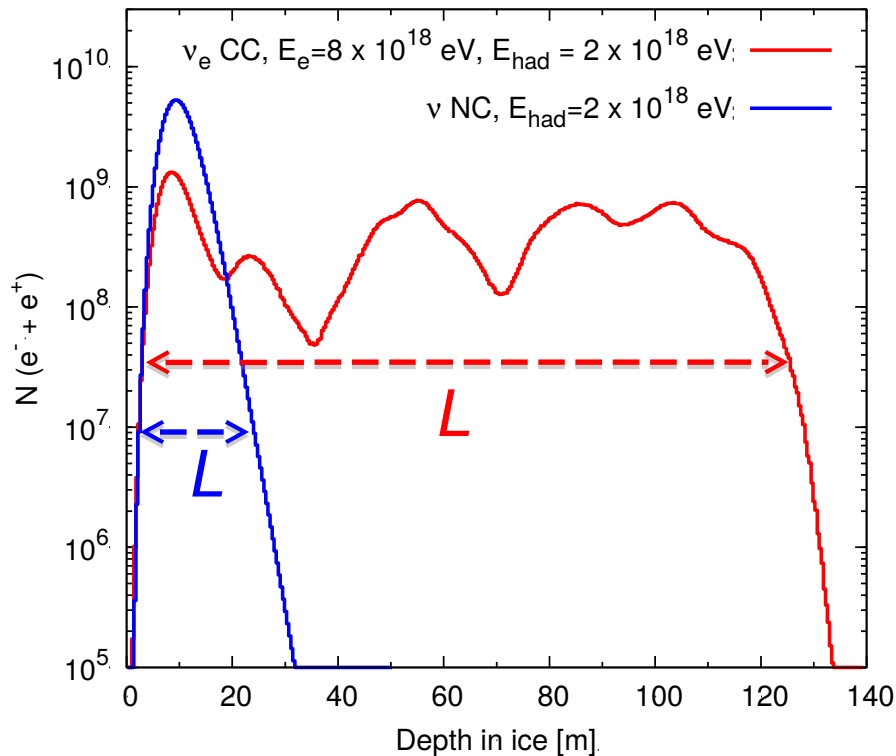
$$p(\omega, \theta) = \omega (1 - n \cos \theta)/c$$



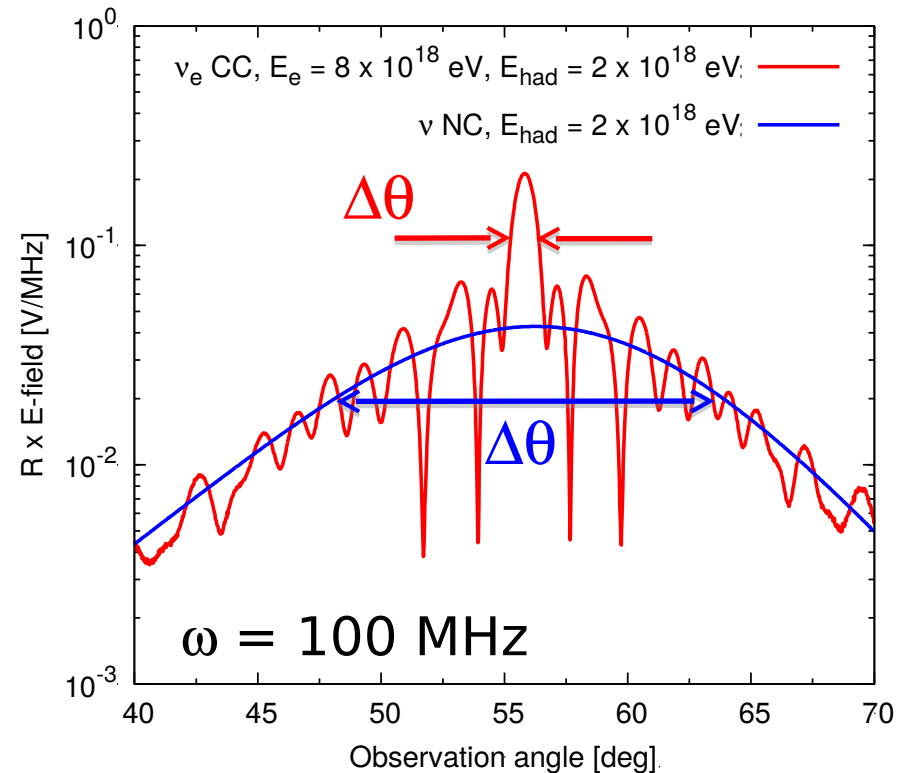
$$\Delta\theta \sim \frac{1}{\omega L}$$

Radioemission & LPM effect in neutrino-induced showers

Longitudinal profile of shower



θ distr. of E-field

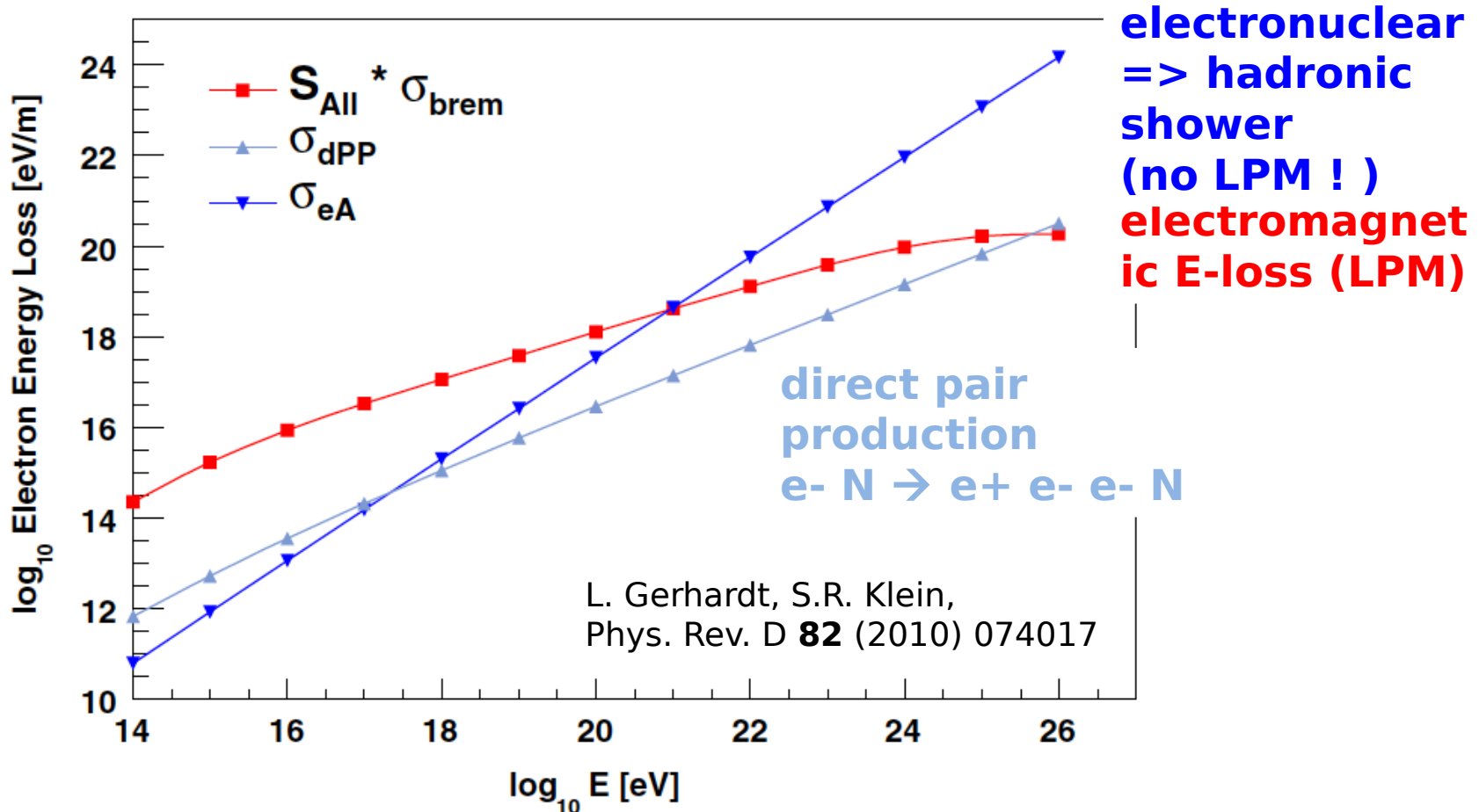


- LPM effect “stretches” electron-induced showers in ne CC => Cherenkov cone “shrink
- Reduced solid angle: $\Delta\theta \sim 1^\circ$ @ 10 EeV & 100 MHz

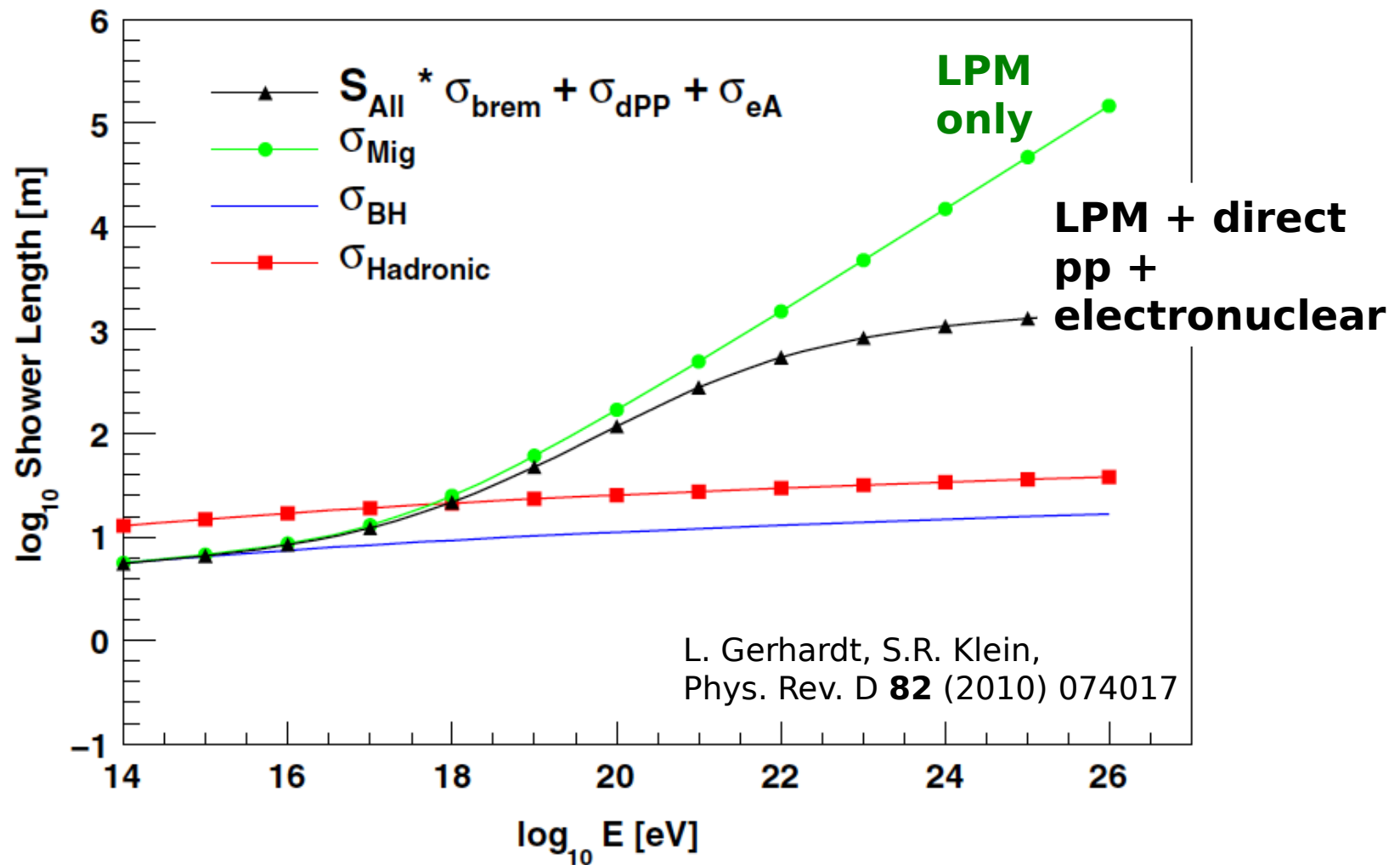
$$\Delta\theta \sim \frac{1}{\omega L}$$

LPM effect vs electronuclear interactions

- LPM effect suppresses electron bremsstrahlung above 10^{15} eV
- Suppression of energy loss to electromagnetic showers is so large above $\sim 10^{20}$ eV that electron starts to lose energy to hadronic showers (little affected by LPM !)



Shower length L vs energy

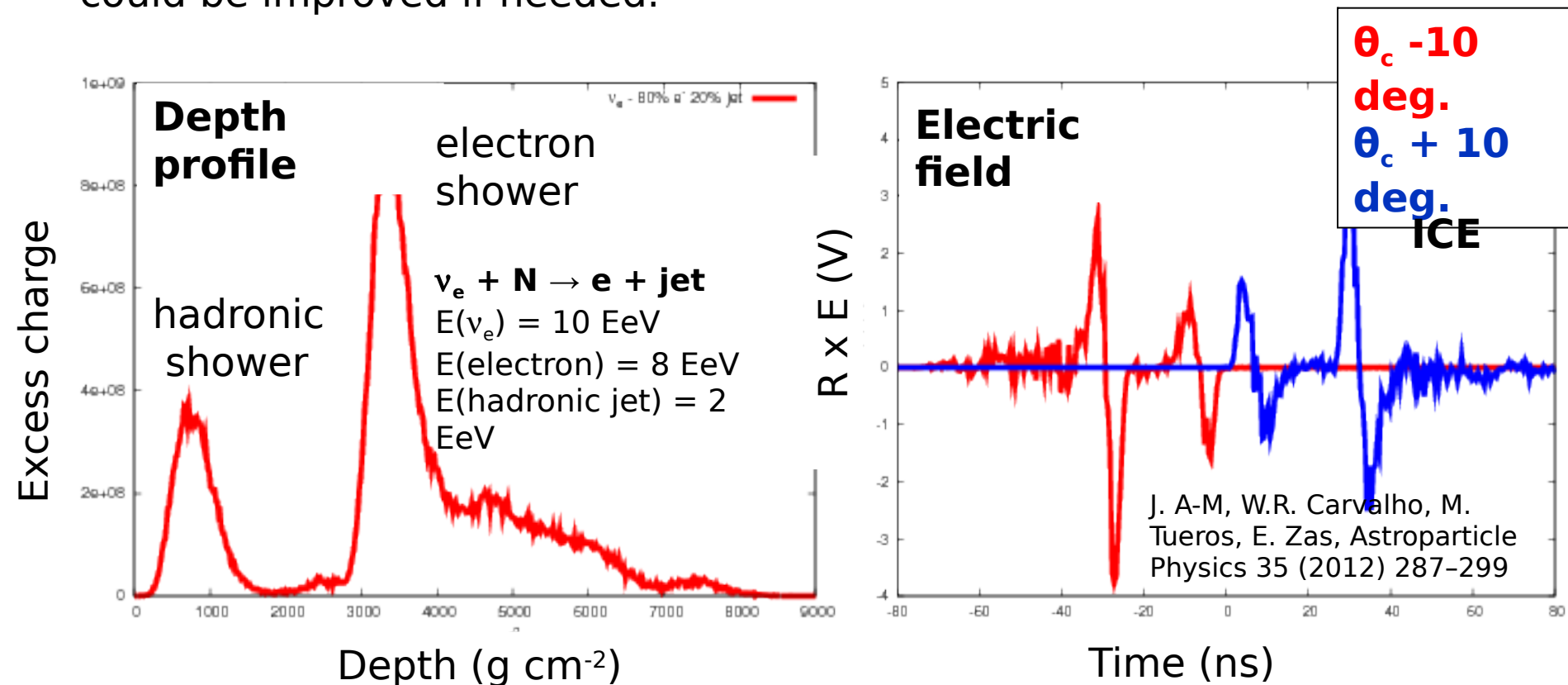


- Moderate rise in electron neutrino-induced shower length, ameliorating the dramatic shower stretching due to LPM effect.
- Should increase relative contribution to exposure from electron neutrinos

Full simulations of ν -induced showers in regolith/megaregolith/etc...

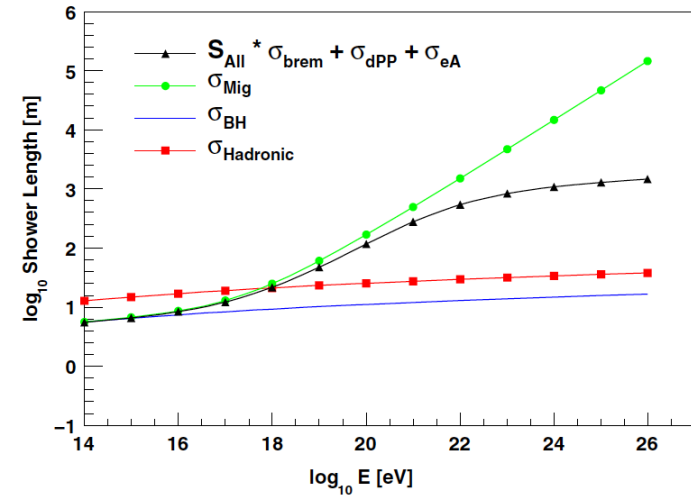
Full simulations of ν -induced showers in (almost any) dense dielectric media now possible with ZHAireS: hadronic, mixed, tau decay-shower,...

Scaling of radiopulse from ice to regolith probably accurate enough but could be improved if needed.

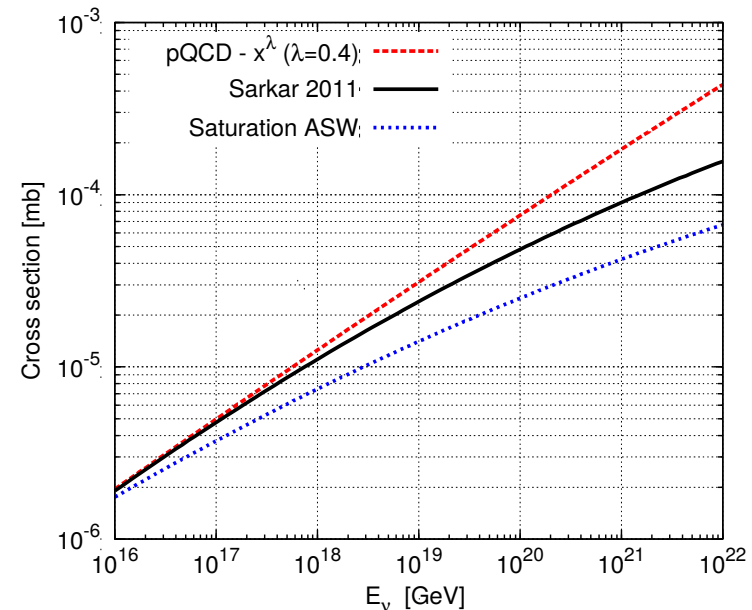


Improvements in modeling

1. Take into account, in calculations of exposure to EHE neutrinos, the hadronic behaviour of electron above 10^{20} eV
=> increase in exposure



1. Include larger uncertainty in the neutrino cross-section

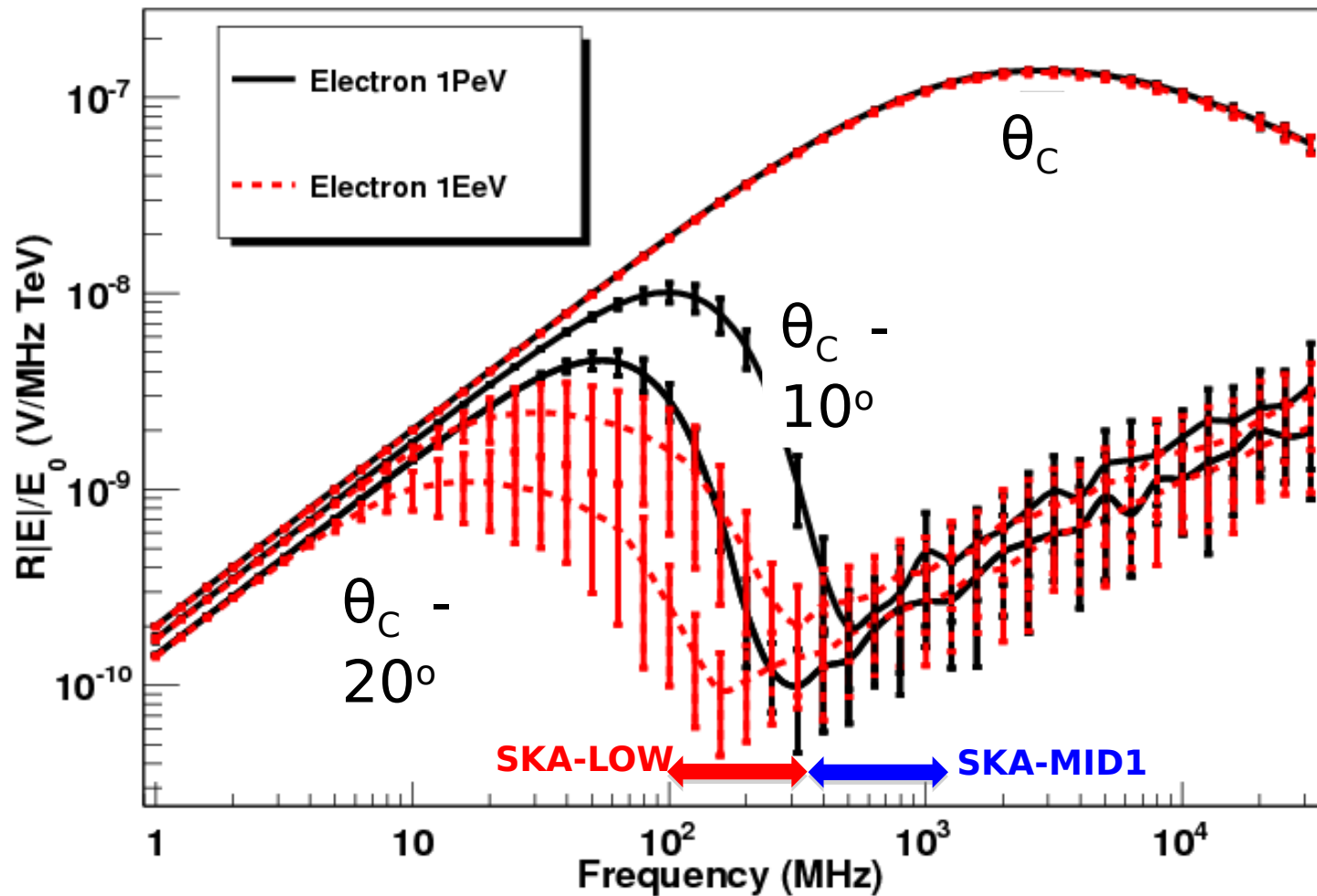


1. New parameterization of showers in regolith/megaregolith ?

Backup

Frequency spectrum in LPM showers

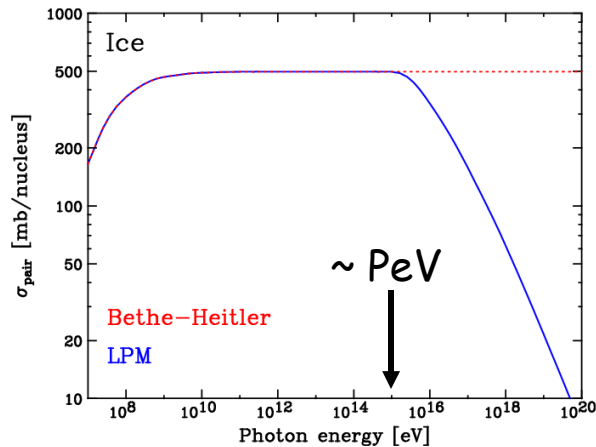
Field normalized to primary energy



- Elongated profiles at EeV induce smaller cut-off frequencies at $\theta \neq \theta_c$
- Cut-off frequency at Cherenkov angle unaffected.

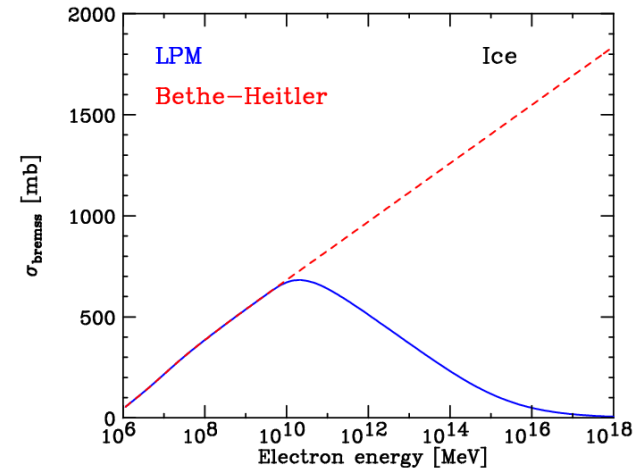
LPM effect: Cross sections

Pair production cross sections

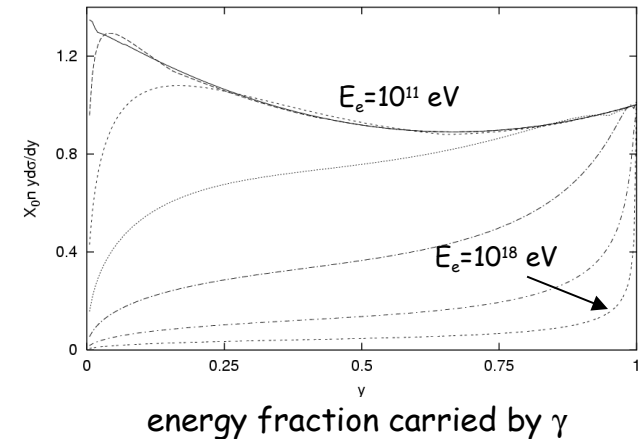
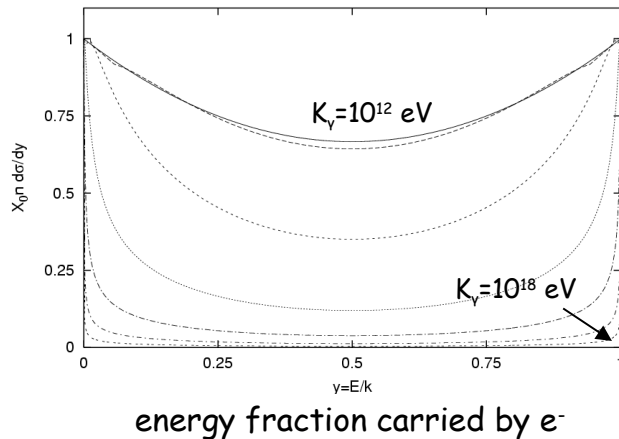


Total
(Ice)

Bremsstrahlung cross sections



Differential
(Lead)



e^-e^+ pair asymmetric in energy

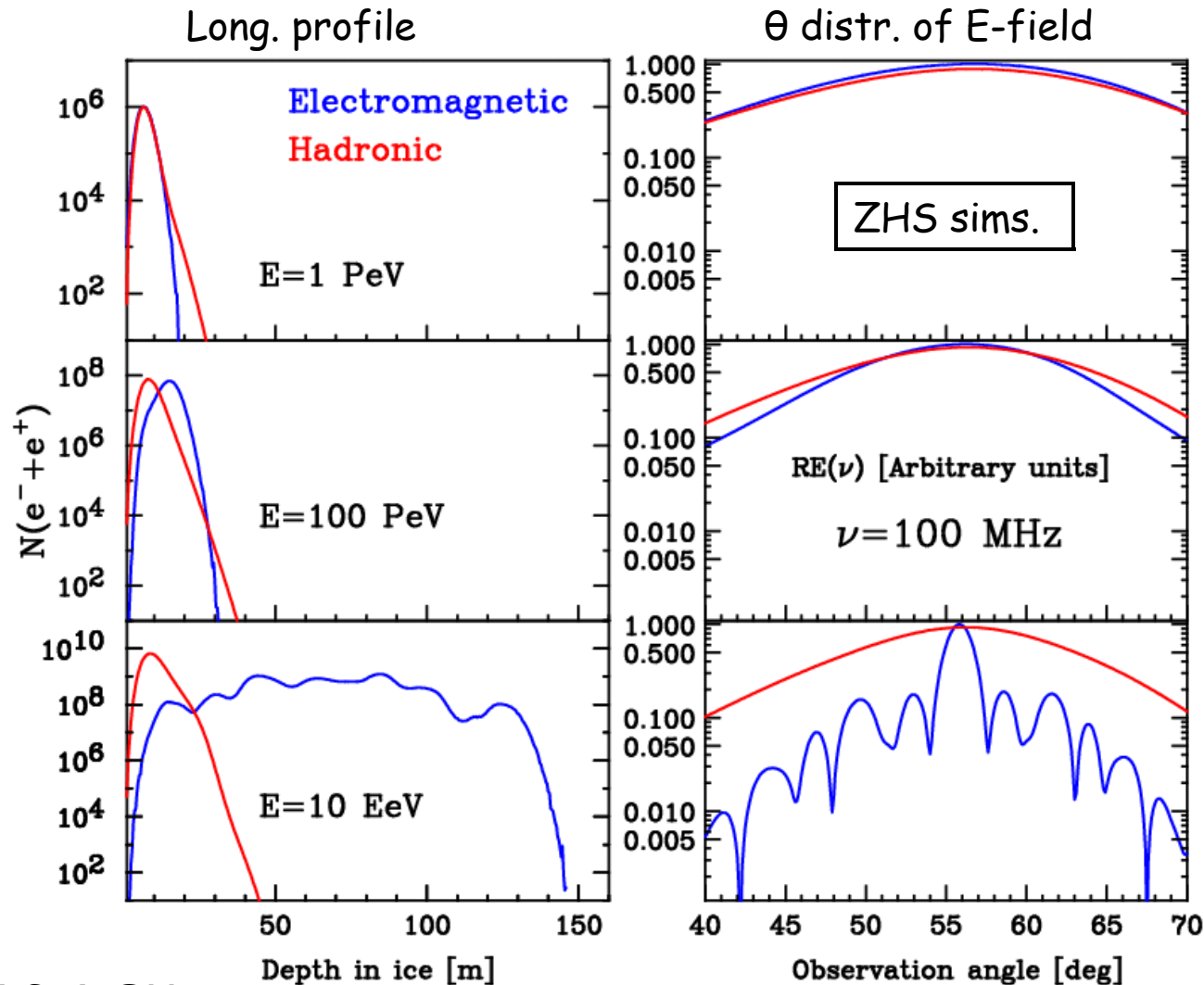
Suppression of emission of low K_γ photons

Radioemission & LPM effect

$\Delta\theta \sim 1/L$ (L = width of the longitudinal shower development).
(Power @ θ_c still $\sim E_{\text{shower}}$)

Cone shrinks

Peak field at Cherenkov angle unchanged



$\Delta\theta \sim 2$ mrad @ 20 EeV & 1 GHz

update value using new params. J.A-M, R.A. Vázquez & E.Zas PRD 61 (1999)