

Binary pulsar searching with GPUs

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Outline

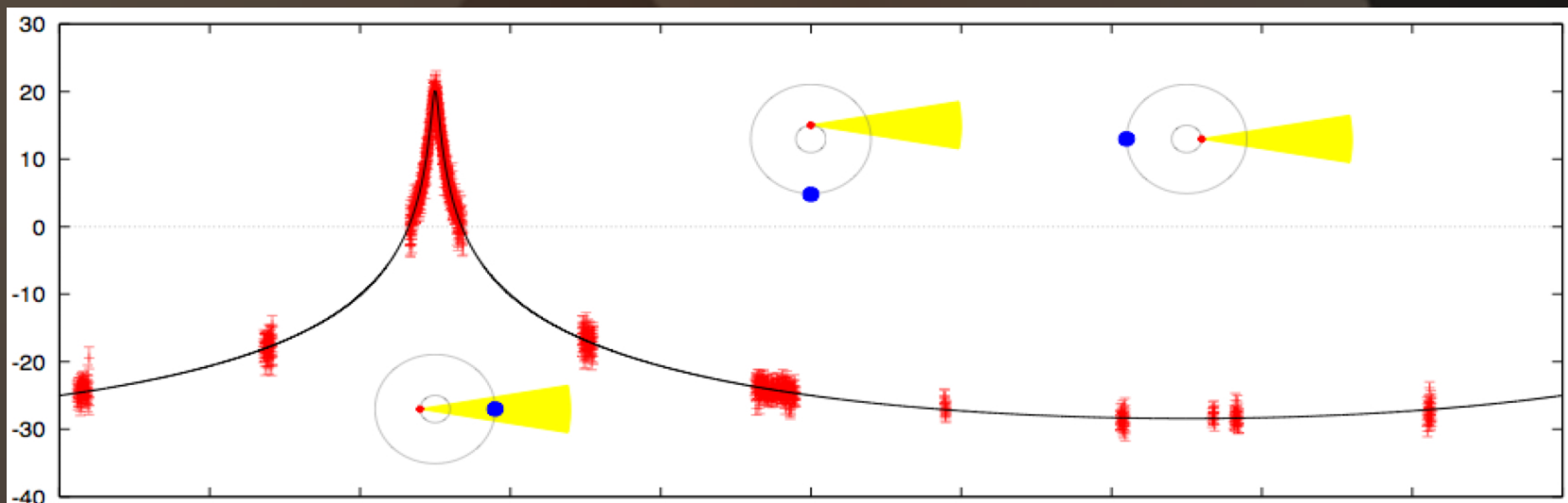
- Motivation
- Pulsar searching
 - Fourier analysis
 - Binary motion modifications
 - Revealing the pulsar
- Applications of GPUs
 - What are GPUs?
 - GPUs in searching

Motivation

- To find new pulsars!
- Particularly millisecond pulsars (MSPs)
- Mass-accretion phase to spin-up to millisecond periods
- Slowing increasing very slowly — long-lived
- GW detection-possibility improvement
- Often in binary systems (PSR-WD, PSR-NS, or PSR-PSR!)
- Timing binaries to test theories of gravity

Pulsars in binary systems

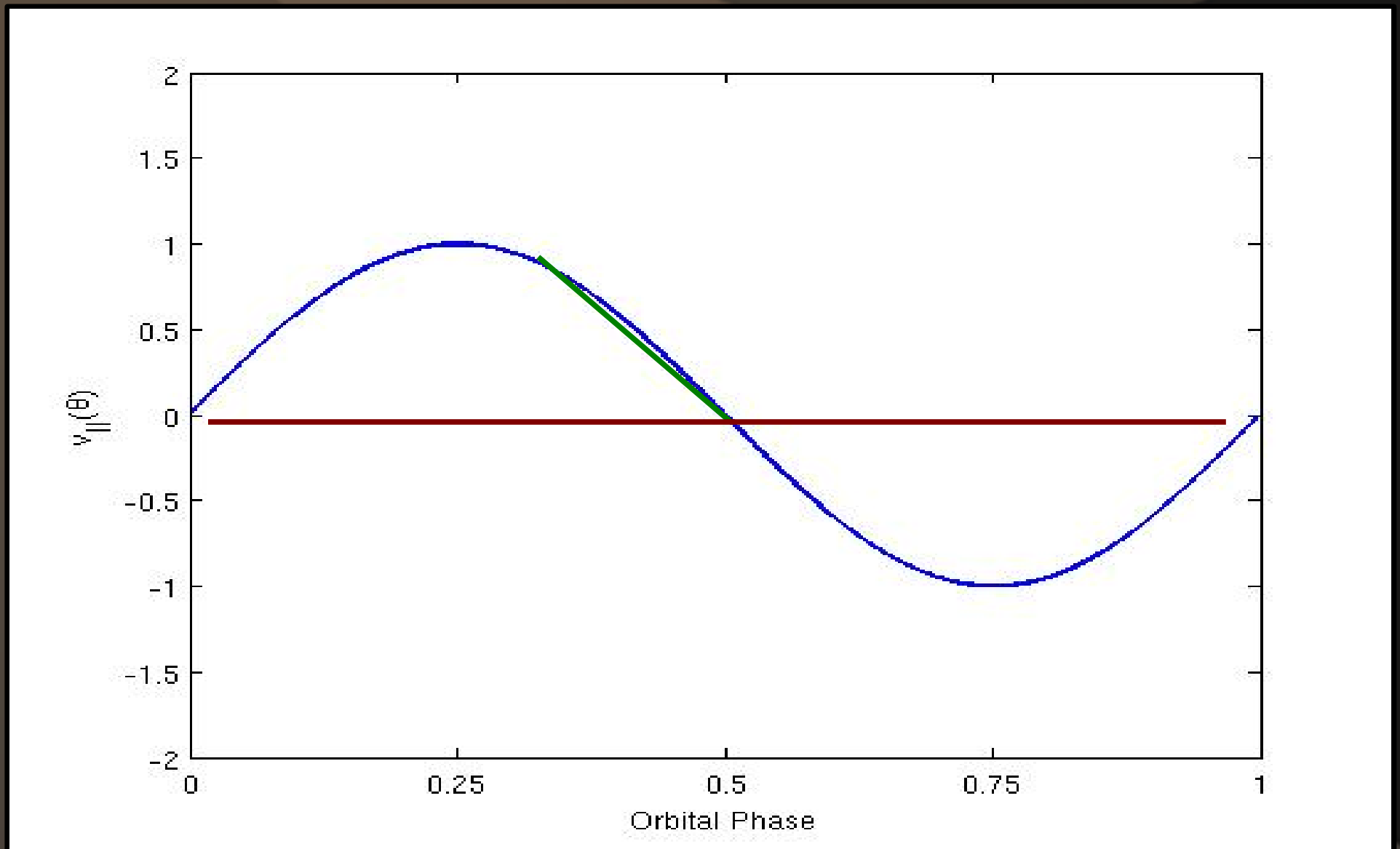
- Why are they interesting?
 - Gravitational effects (e.g Shapiro Delay)
 - Binary stellar evolution models
- Doppler effects make pulsars in binary systems difficult to detect



Fourier Analysis

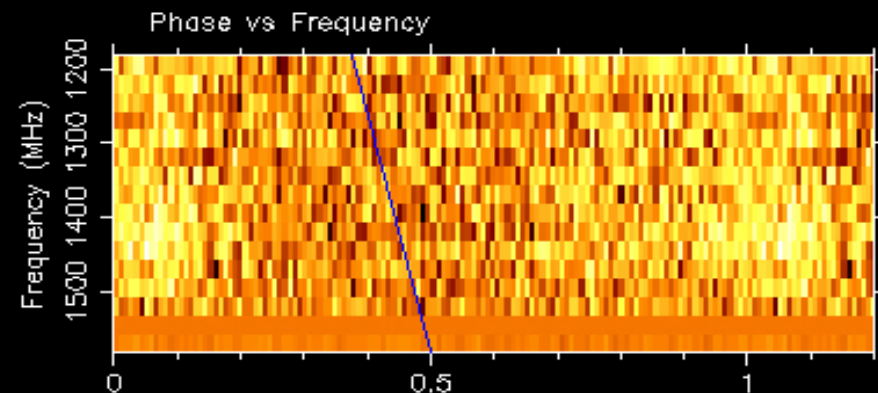
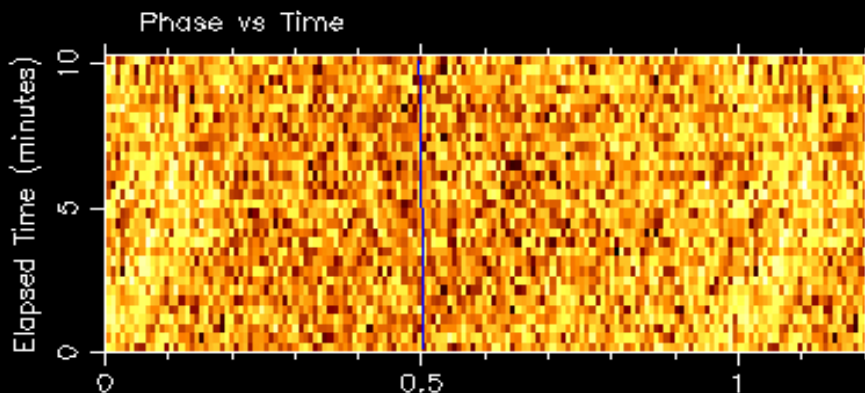
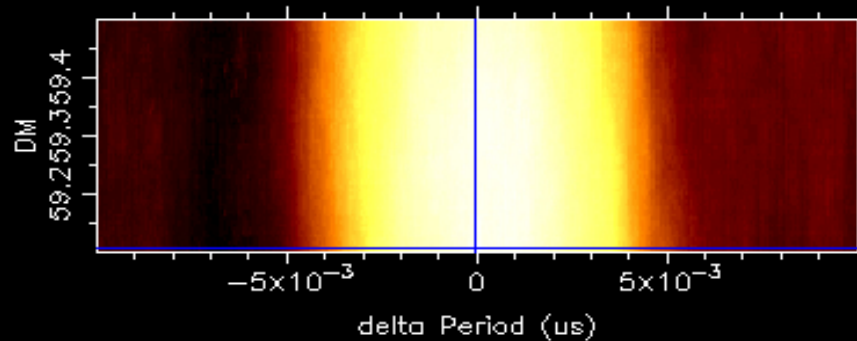
- Pulsars are usually detected by searching for a pulsed signal at an unknown period and DM
- Dedisperse at a test DM value and `fft(time_series)`
- Harmonic folding
- Pick out significant peaks
- In a binary system the power in the Fourier spectrum is spread around the true pulse frequency by Doppler effect
- Need methods to restore the spread power to a single Fourier bin

Doppler effect on observed period

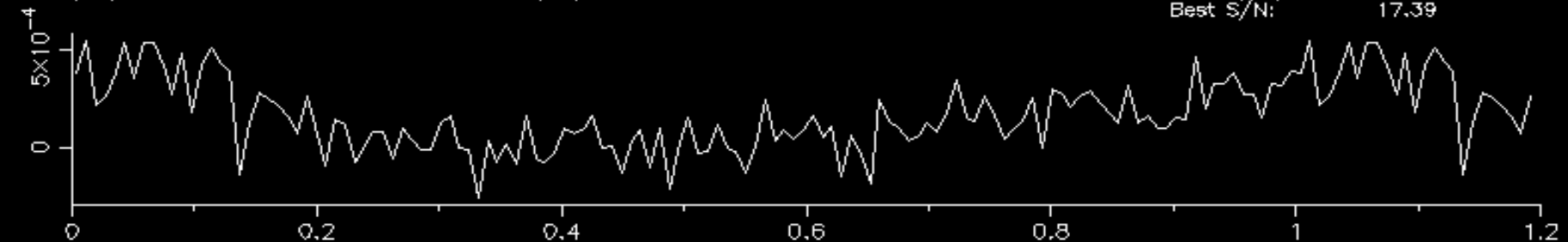


A brand new MSP

BC P(ms)= 2.011794606 TC P(ms)= 2.011940033 DM= 59.300 RAJ= [redacted] DecJ= [redacted]
 BC MJD = 55756.295298 Centre freq(MHz) = 1382.000 Bandwidth(MHz) = -400 l = [redacted] b = [redacted]
 NBin = 128 NChan = 16 NSub = 31 TBin(ms) = 0.016 TSub(s) = 20.000 TSpan(s) = 616.611
 P(us): offset = 0.00000, step = 0.00005, range = 0.01000 DM: offset = 0.000, step = 0.006, range = 0.200



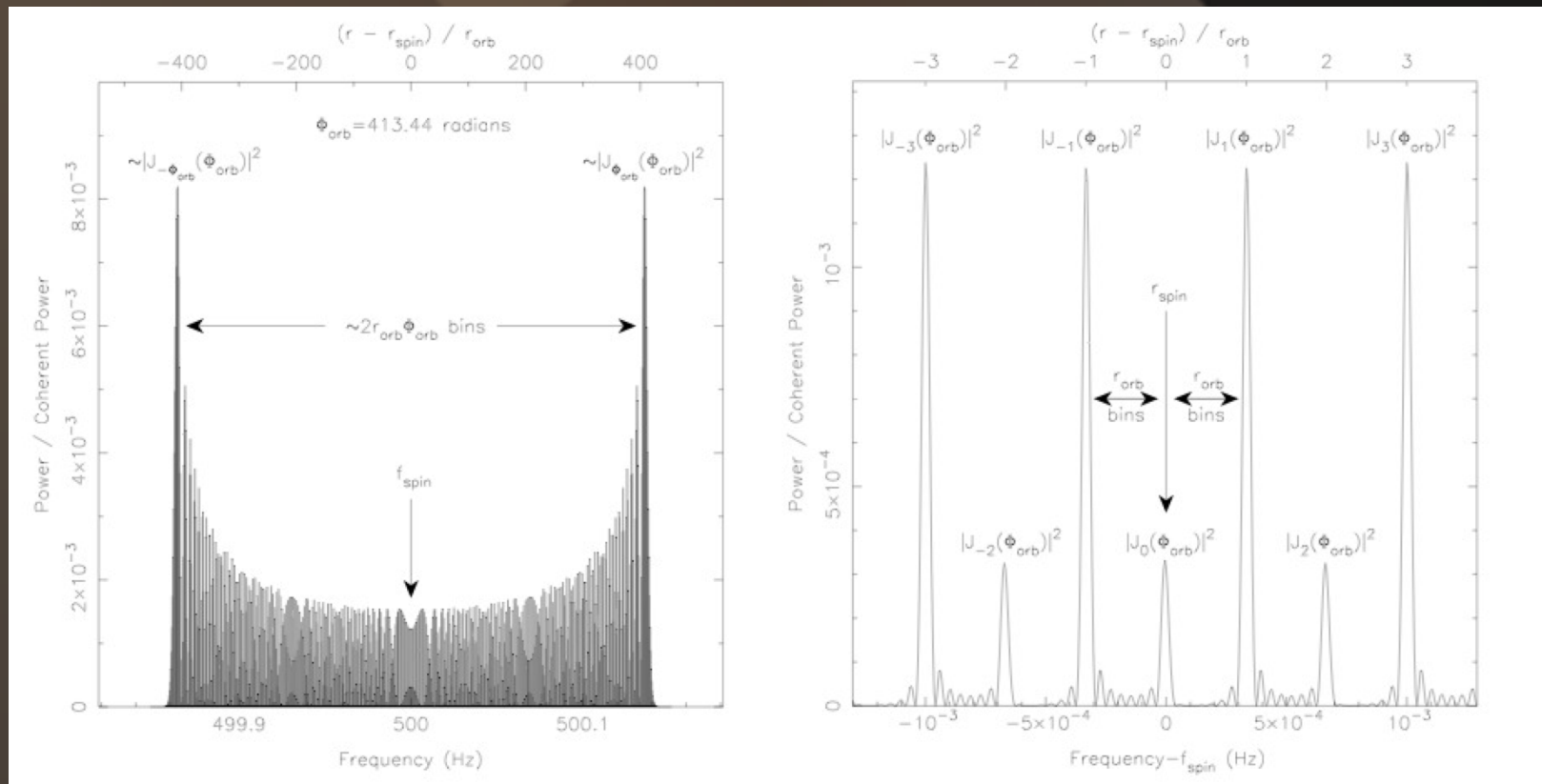
BC prd (ms):	2.011794555	TC prd (ms):	2.011939982	DM:	59.106	BC freq (Hz):	497.068648251
Corrn (ms):	-0.000000051	Corrn (ms):	-0.000000051	Corrn:	-0.194	Freq err. (Hz):	0.000065072
Error (ms):	0.000000263	Error (ms):	0.000000263	Error:	0.064	Width (ms):	0.440
						Best S/N:	17.39



Orbital Motion Corrections

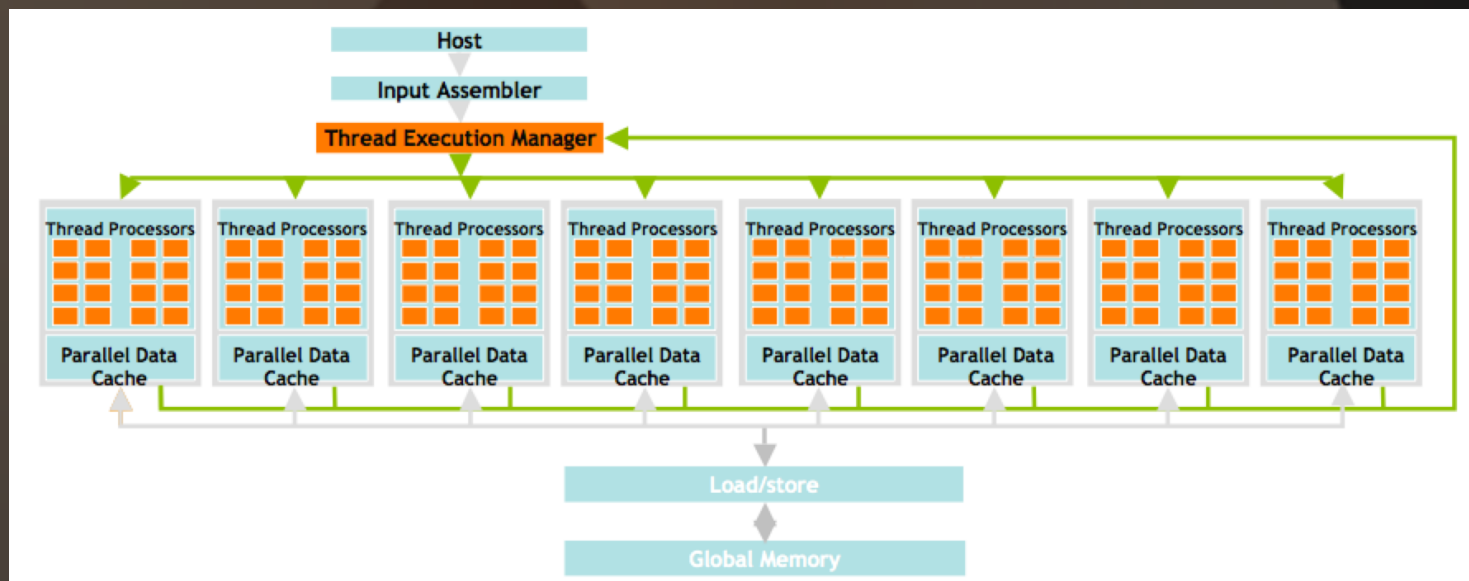
- Stretch and resample the time-series pre-FFT to apply a constant acceleration
- The signal appears to originate from a static pulsar
- Normal Fourier analysis continues on the modified time-series
- Good for long orbits relative to the observation time
- Reasonable acceleration range would take ~ 4 days / beam on a HTRU med-lat pointing!
- Need a faster way for the time-consuming steps
- For longer pointings, there is another method

- In this case the power spectrum is affected in a known way about the true pulse frequency
- By convolving the power spectrum with an inverse of the transfer filter power is restored to a single bin
- Fourier transforming small sections of the power spectrum



What are GPUs?

- Massively parallel processing
- Can have hundreds of processing cores on a single GPU
- $\text{£}/\text{flop}$ is lower than traditional supercomputing
- Problem must lend itself to parallelisation



Acceleration on GPUs

- Current pipeline is sequential
- N acceleration test values increases time by factor N
- Already at ~ 4 hours for a med-lat HTRU pointing
- Dedispersion already applied on GPUs (Barsdell, 2010)
 - From 80 minutes to 5 minutes with identical results
- The time consuming steps of acceleration searching could be strongly parallel with high memory locality
 - Both creating the stretched time-series, or searching the power spectrum for modulation

Conclusion

- Pulsars in binary systems are an important tool in gravitational astronomy
- Techniques exist to find these Doppler shifted systems but are simply too time-consuming to be practical
- GPUs have already proven effective in de-dispersion, and acceleration searching also lends itself to massively parallel processing

Thank you