

# RFI | Characterisation and Excision

DARA 2021

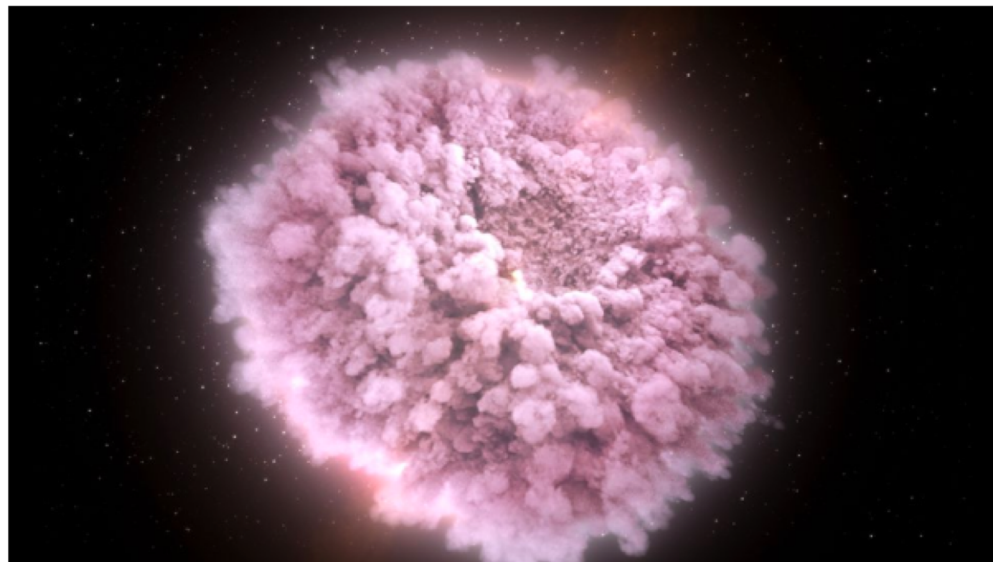
**By**  
Nadeem Oozeer  
Data Science  
RARG



# Mysterious Gravitational Wave Sparks Days-Long Hunt — But It Was Just a Glitch

By [Rafi Letzter - Staff Writer](#) 3 days ago [Space](#)

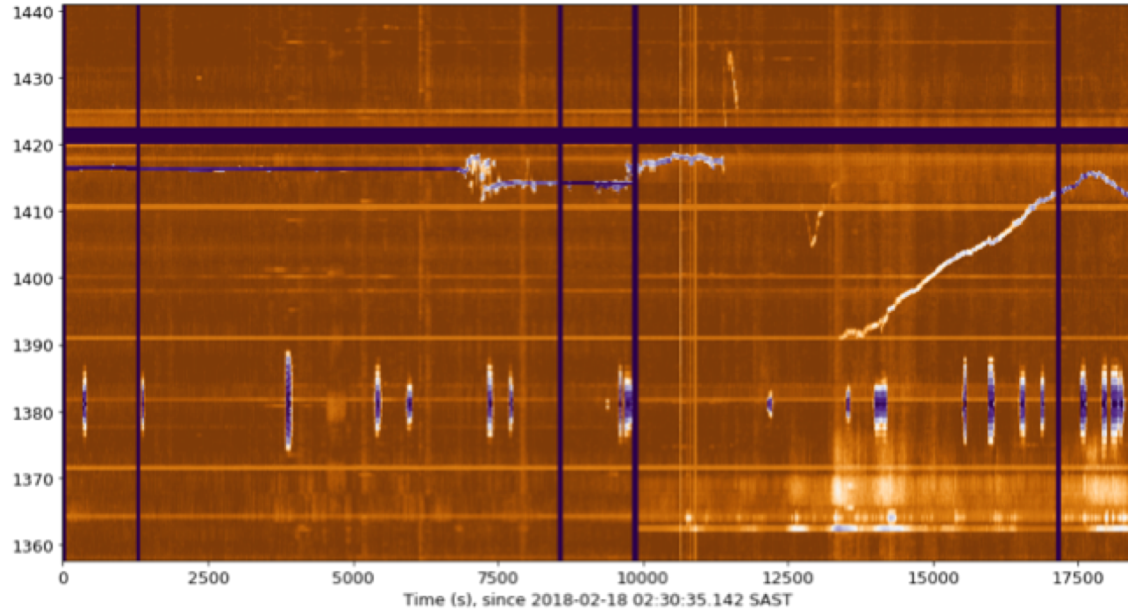
Astronomers and physicists all over the world were disappointed.



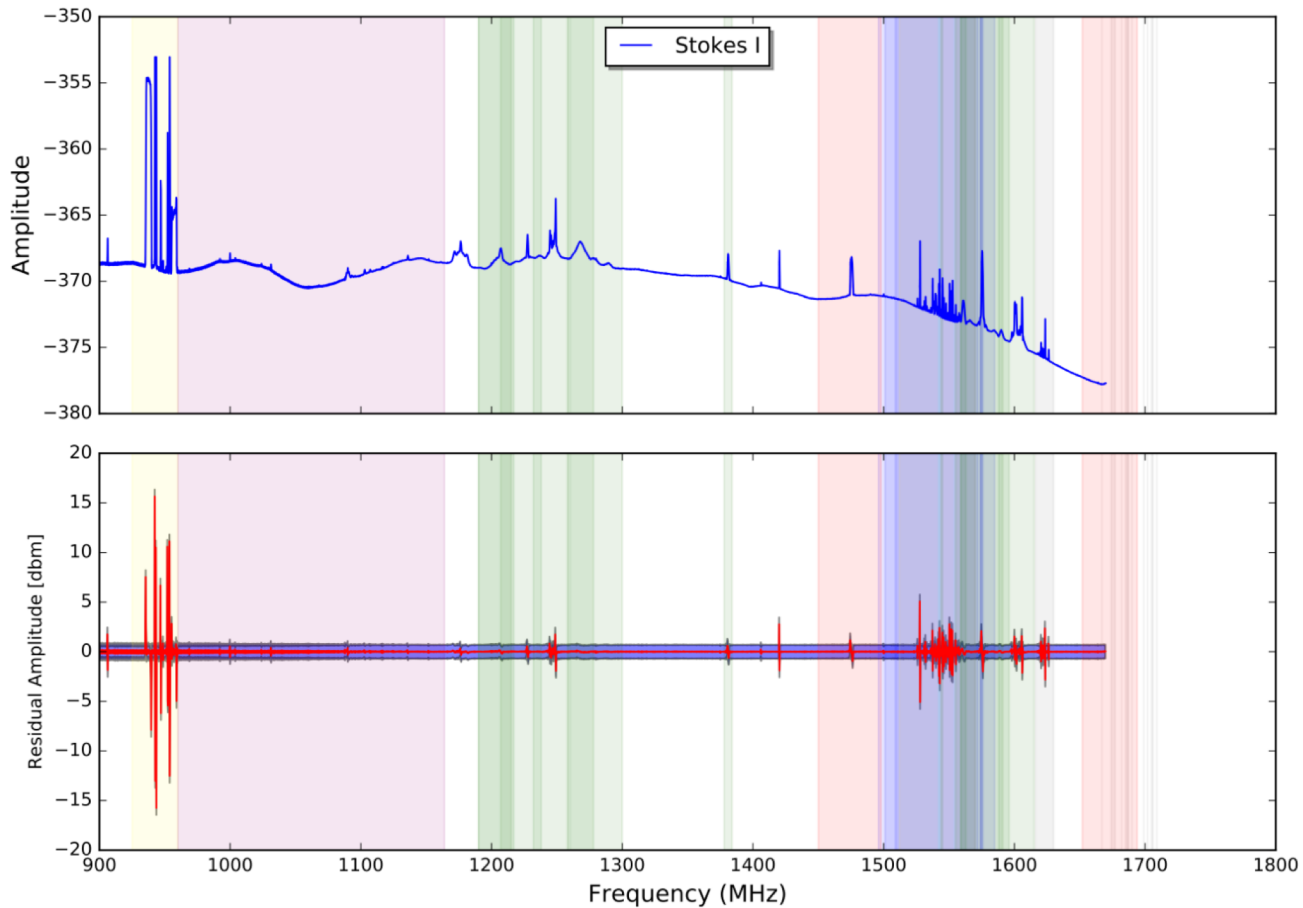
An illustration shows a supernova, one of the events raised as a possible source for the wave.  
(Image: © NASA Goddard Space Flight Center)

# Background

- RFI comes in various forms.
- RFI can for example be caused by television stations, aeroplanes and radar, while broadband RFI caused by phenomena such as lightning, high-voltage power cables and sparking electrical fences are often local in time.
- Strong RFI that is problematic is often either local in frequency or in time.
- A different class of RFI is caused by weakly transmitting but stationary -- and therefore systematic -- devices on site.



# L-band RFI



Observer: RTS

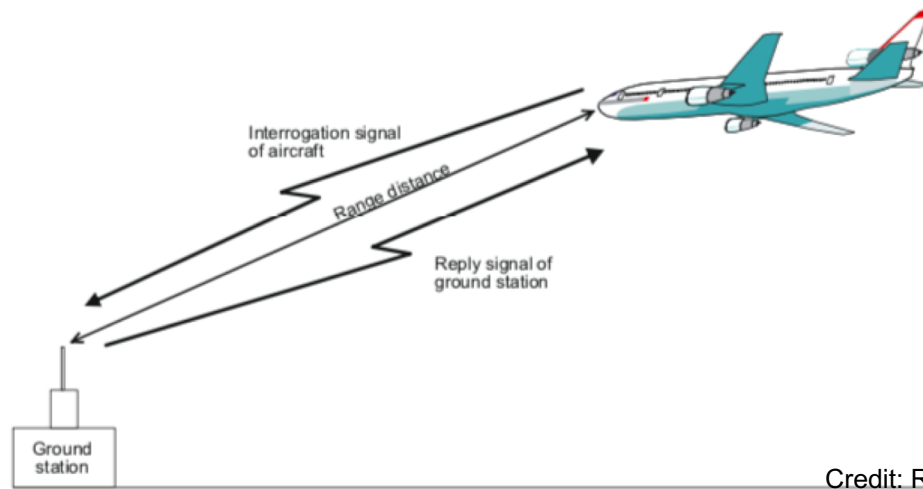
Product Type: 2.9 Self\_Generated\_RFI  
Start time: 2015-08-04 16:25:14.874

Observation data:  
/var/kat/archive/data/RTS/telescope\_products/2015/08/04  
/1438705513.h5

Antenna used: m063

Number of tracks:  
111

Colour band used  
Yellow - GSM  
Purple - Astro Nav  
Red - Afristar  
Green - GPS sat



Credit: Rohde & Schwarz

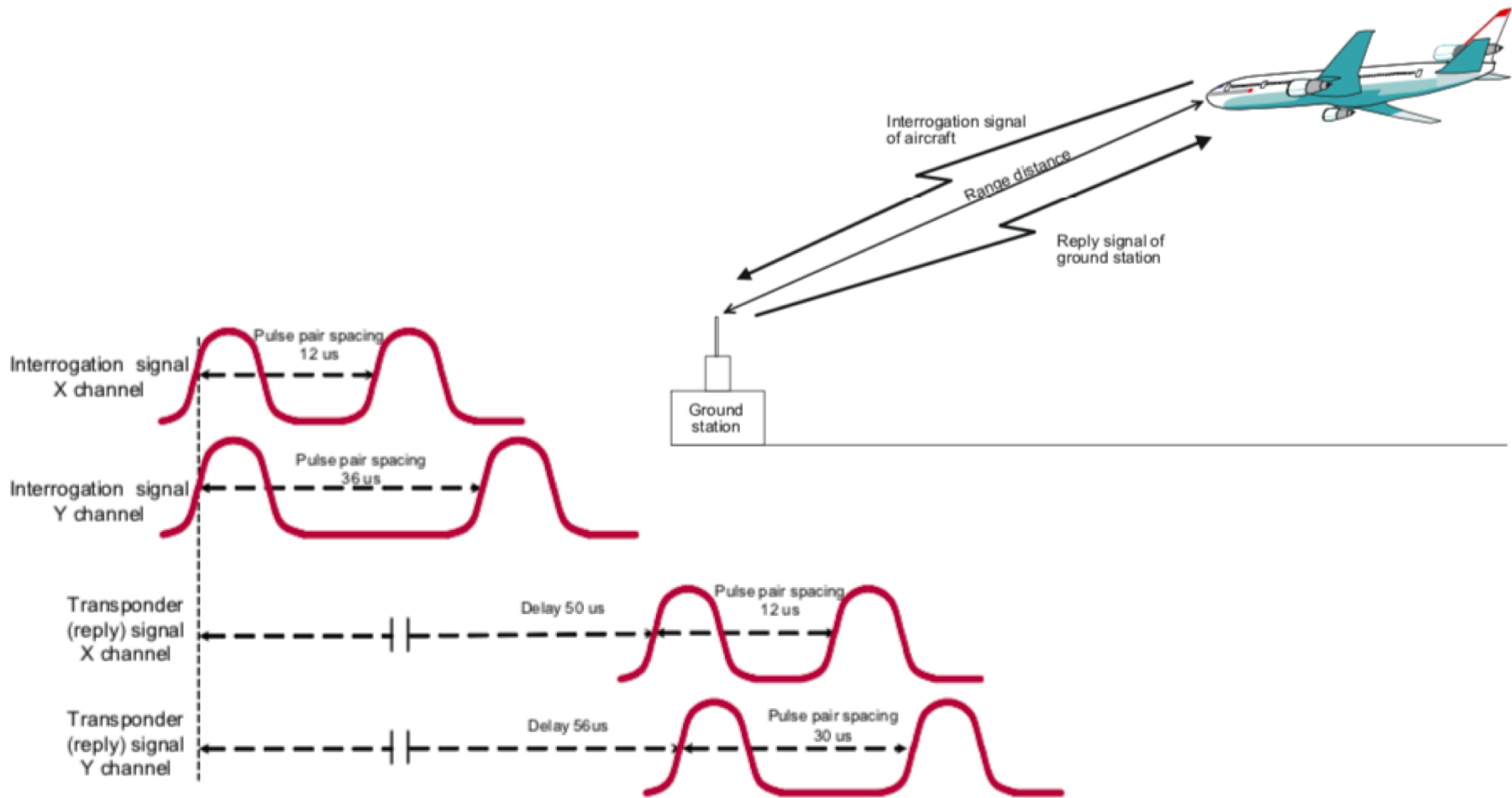
# DME RFI

Nadeem Oozer, SARAO Data Science Team

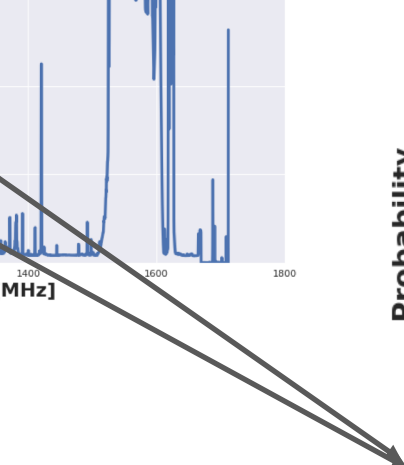
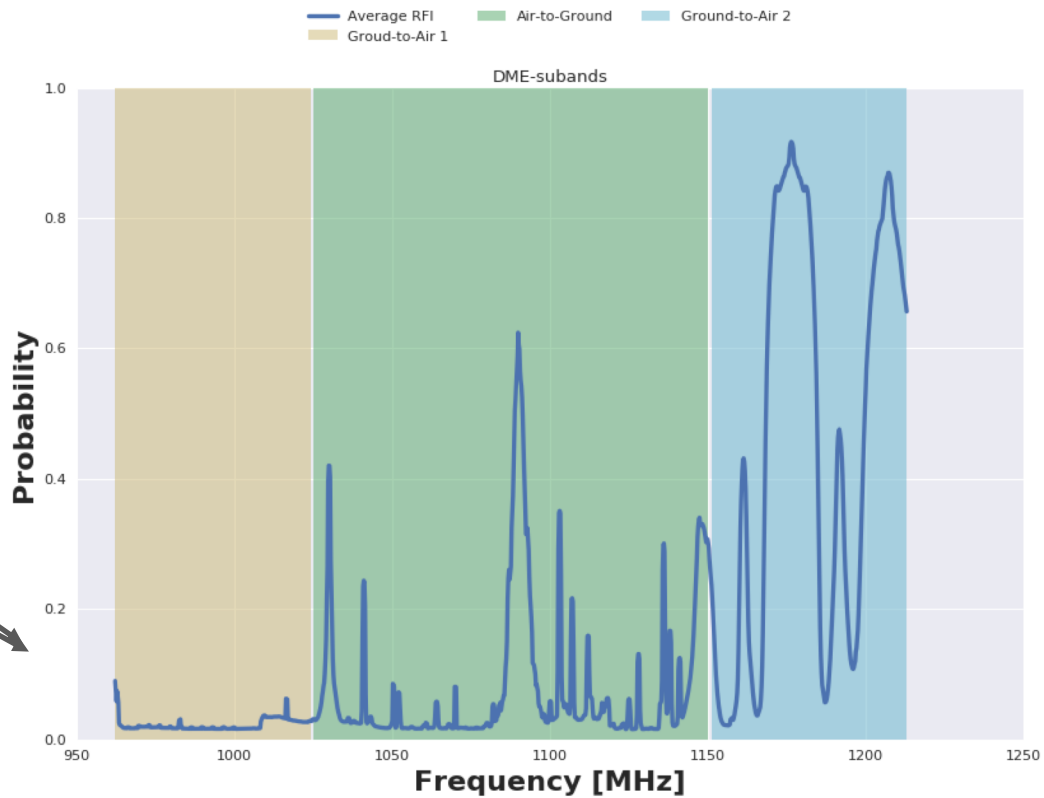
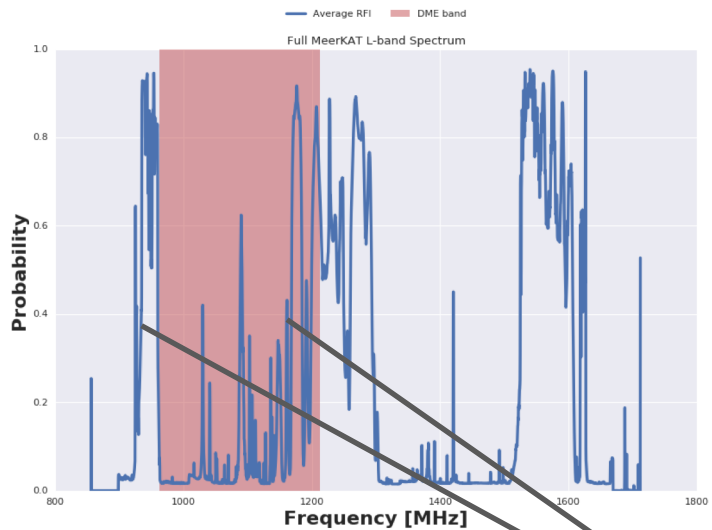
Email : [nadeem@ska.ac.za](mailto:nadeem@ska.ac.za)

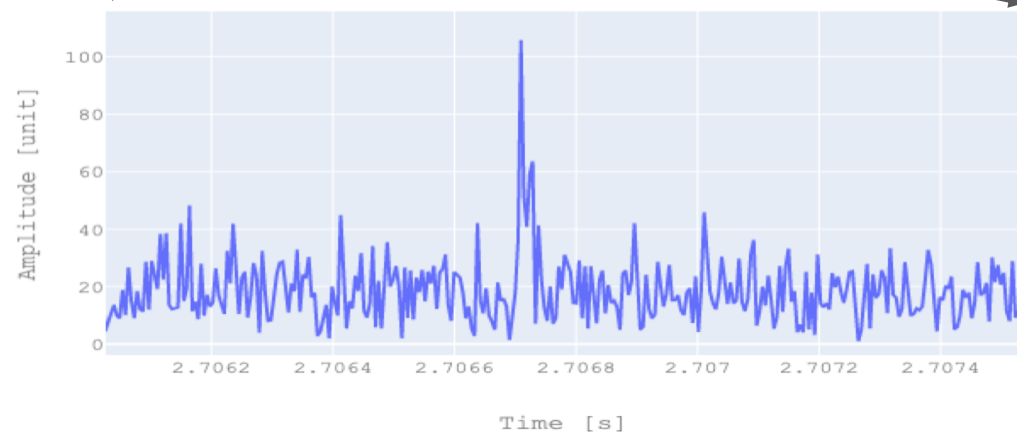
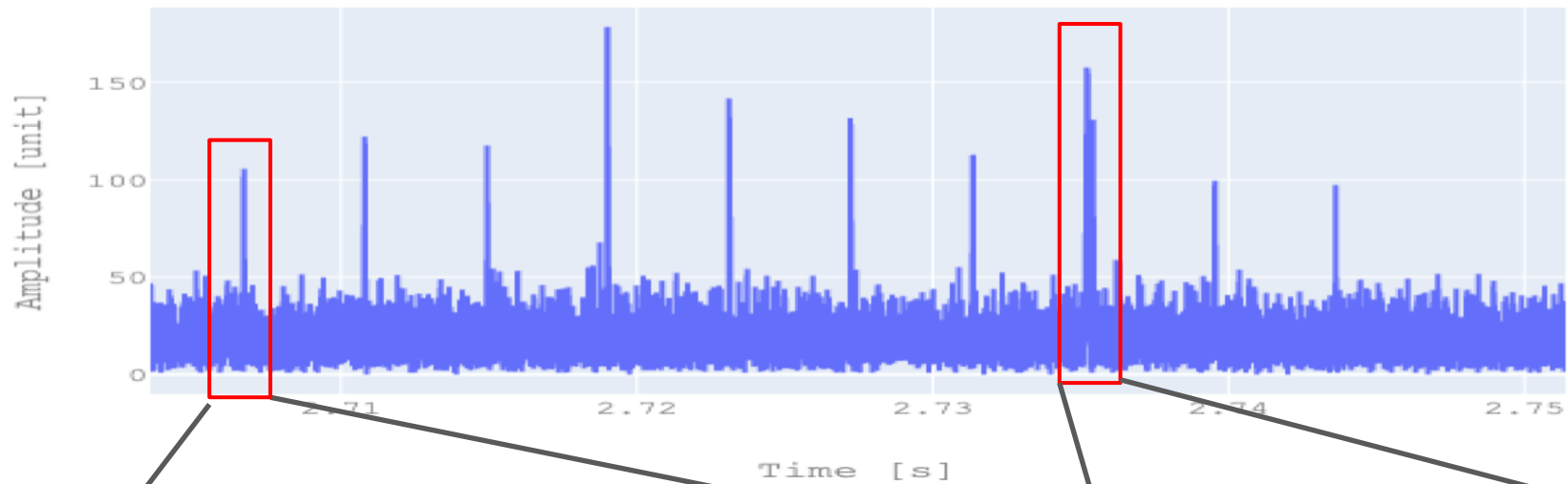
B. Bassett, C. Finlay, I. Sihlangu

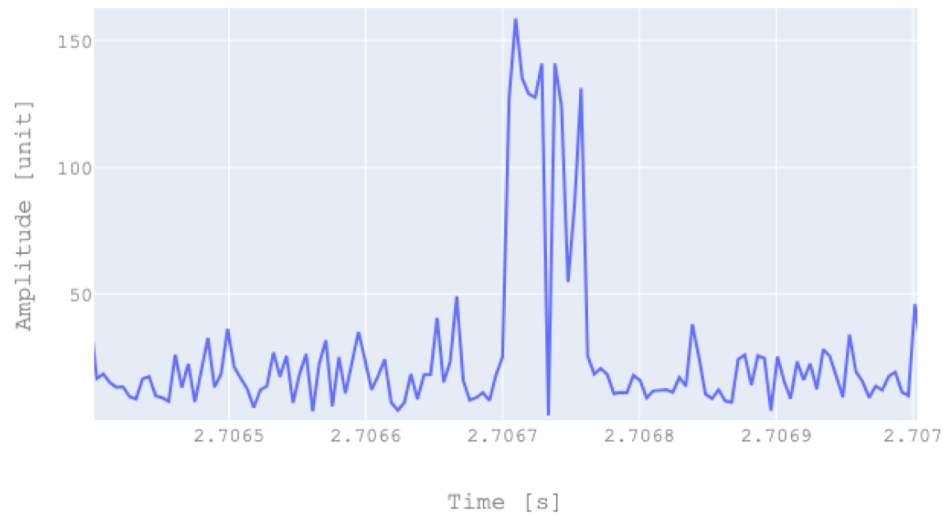
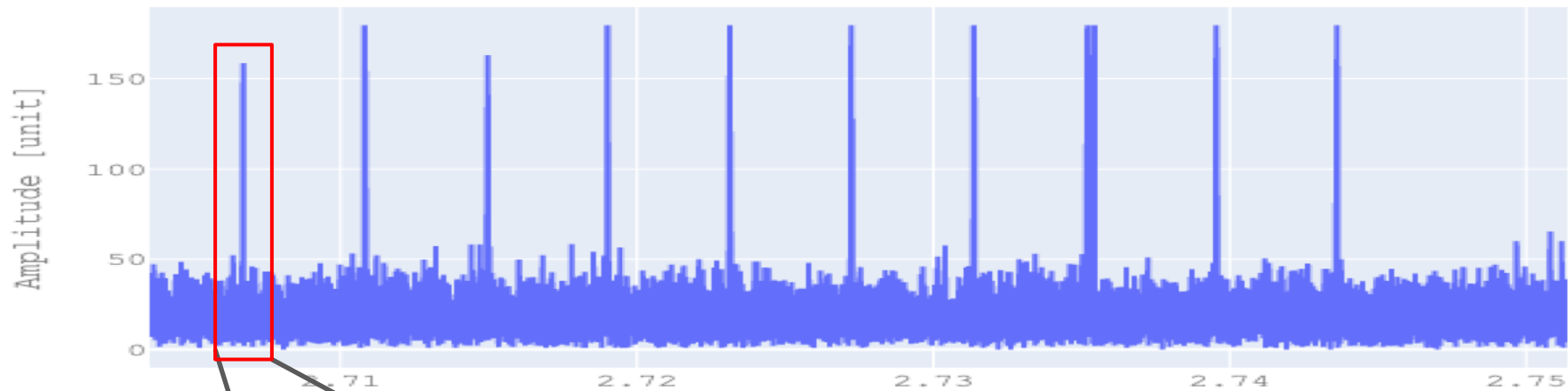




**Figure 2: Time characteristic of DME signal envelope for X and Y channel.**







# MeerKAT RFI

## Historical Probabilities (HP)

Isaac Sihlangu, SRAO Data Science Team

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R. Armstrong, B. Bassett, C. Finlay, N.Oozeer



Karoo

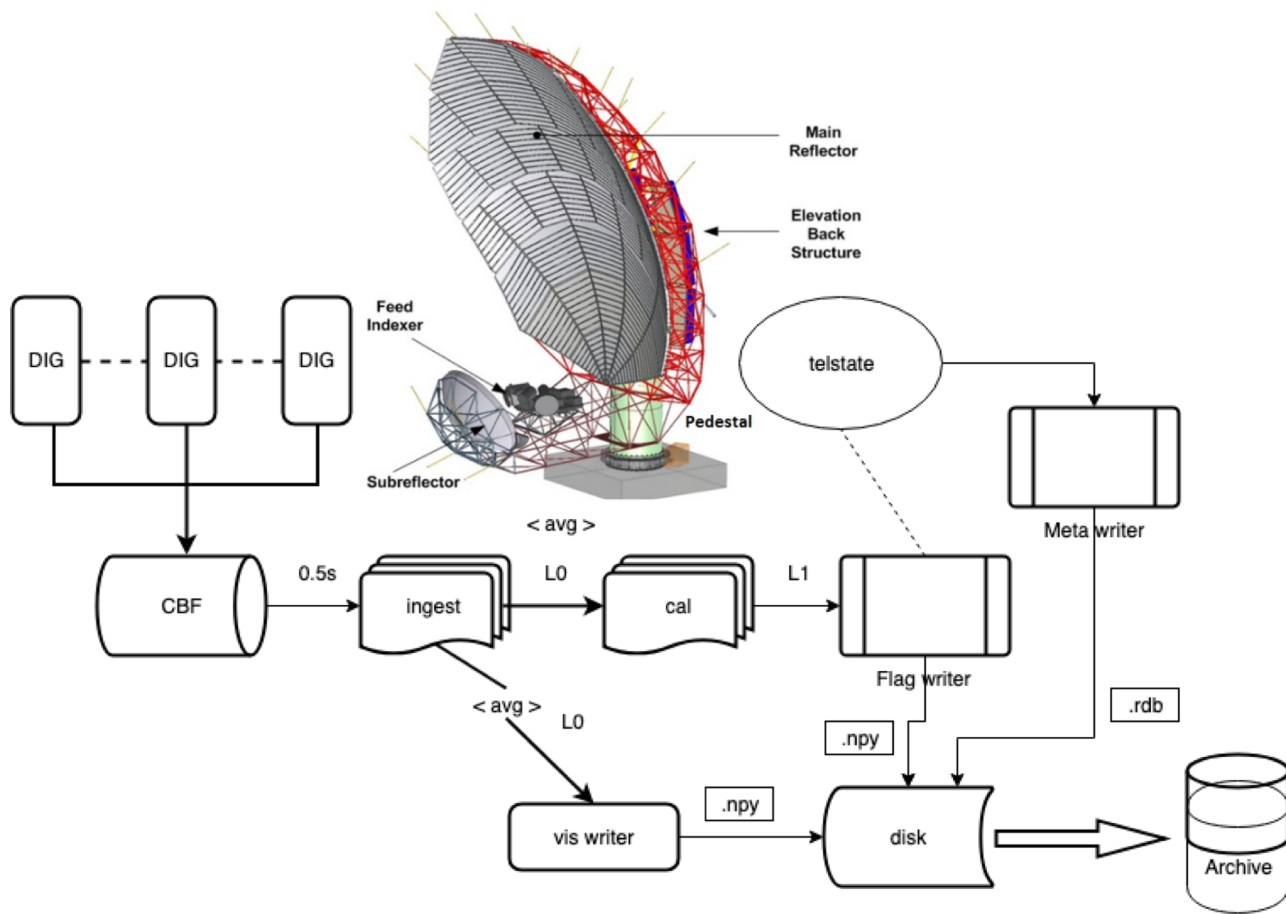
Array

Telescope

Historical

Probability

RFI



Isaac Sihlangu et al.

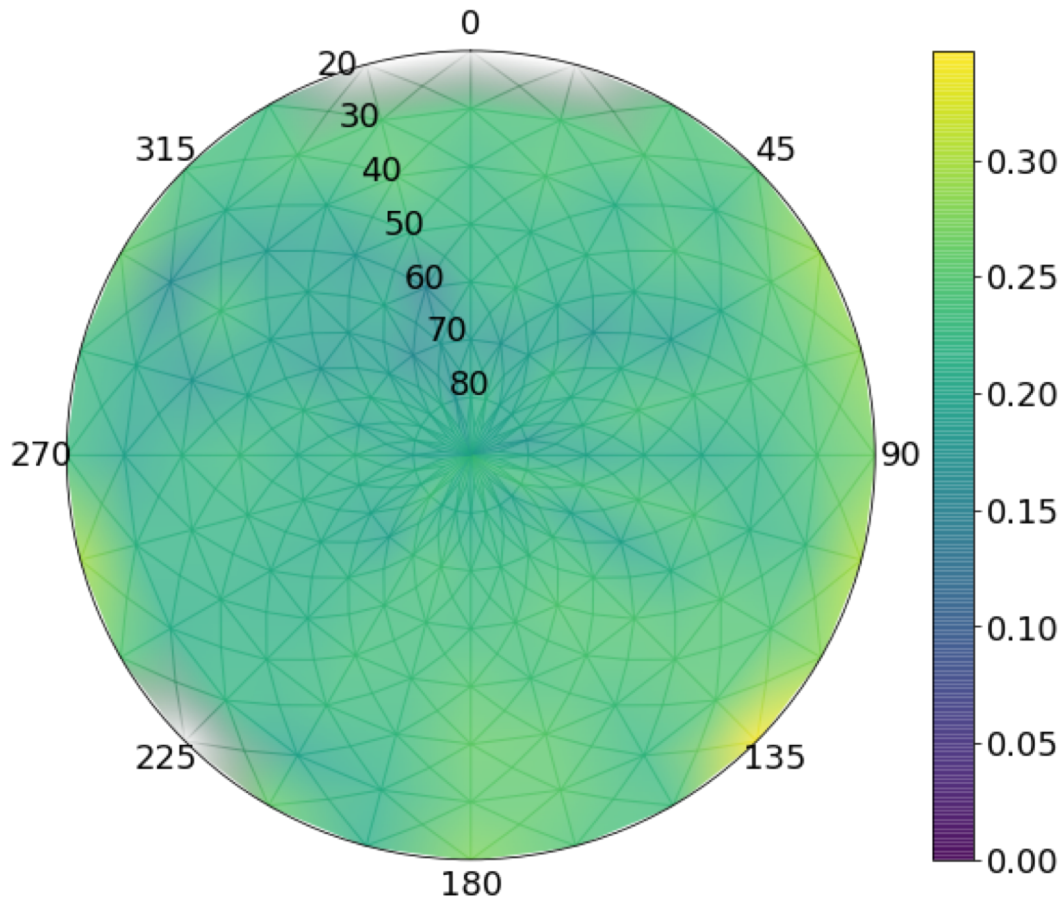
<https://arxiv.org/pdf/2008.08877>

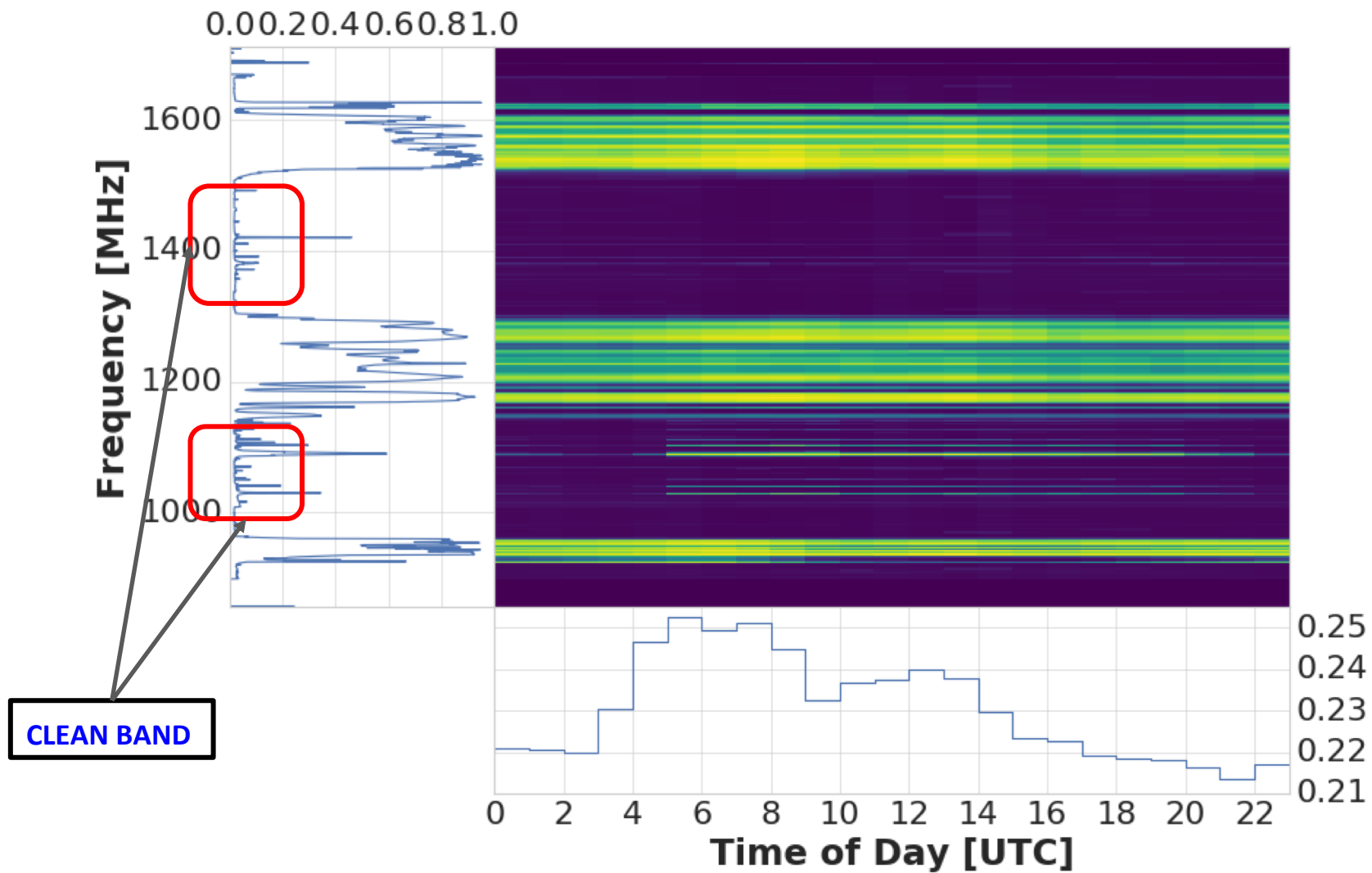
# Use cases

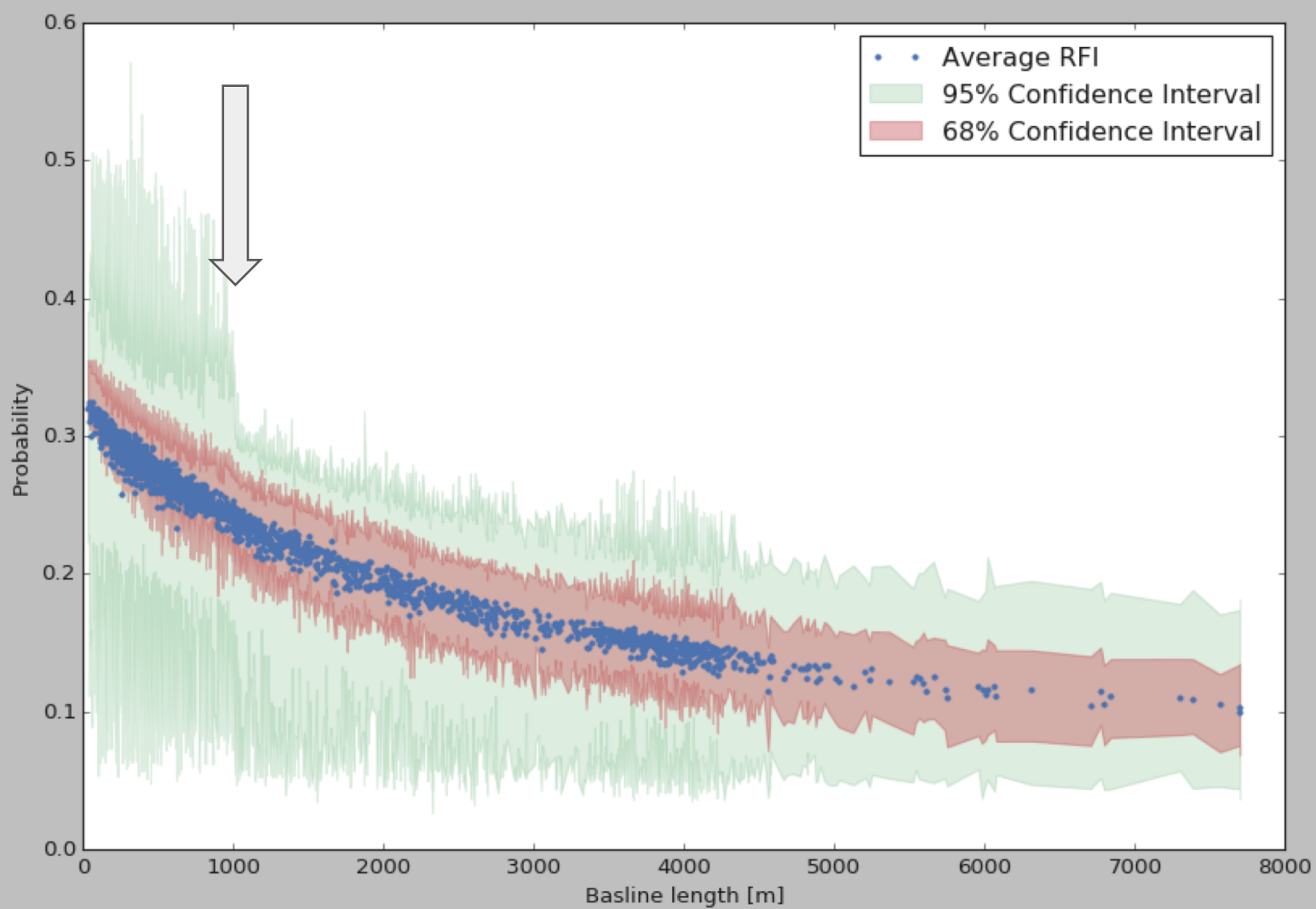
- **Operations:**
  - Schedule desired critical observations
- **RFI working group**
  - Provide an alert system if level of RFI at certain frequency goes beyond a critical threshold
- **Astronomers**
  - Optimise flagging
- **Policy makers**
  - Spectrum and telecoms team to monitor occupancy of dedicated frequency bands

# Full sky - RFI probability

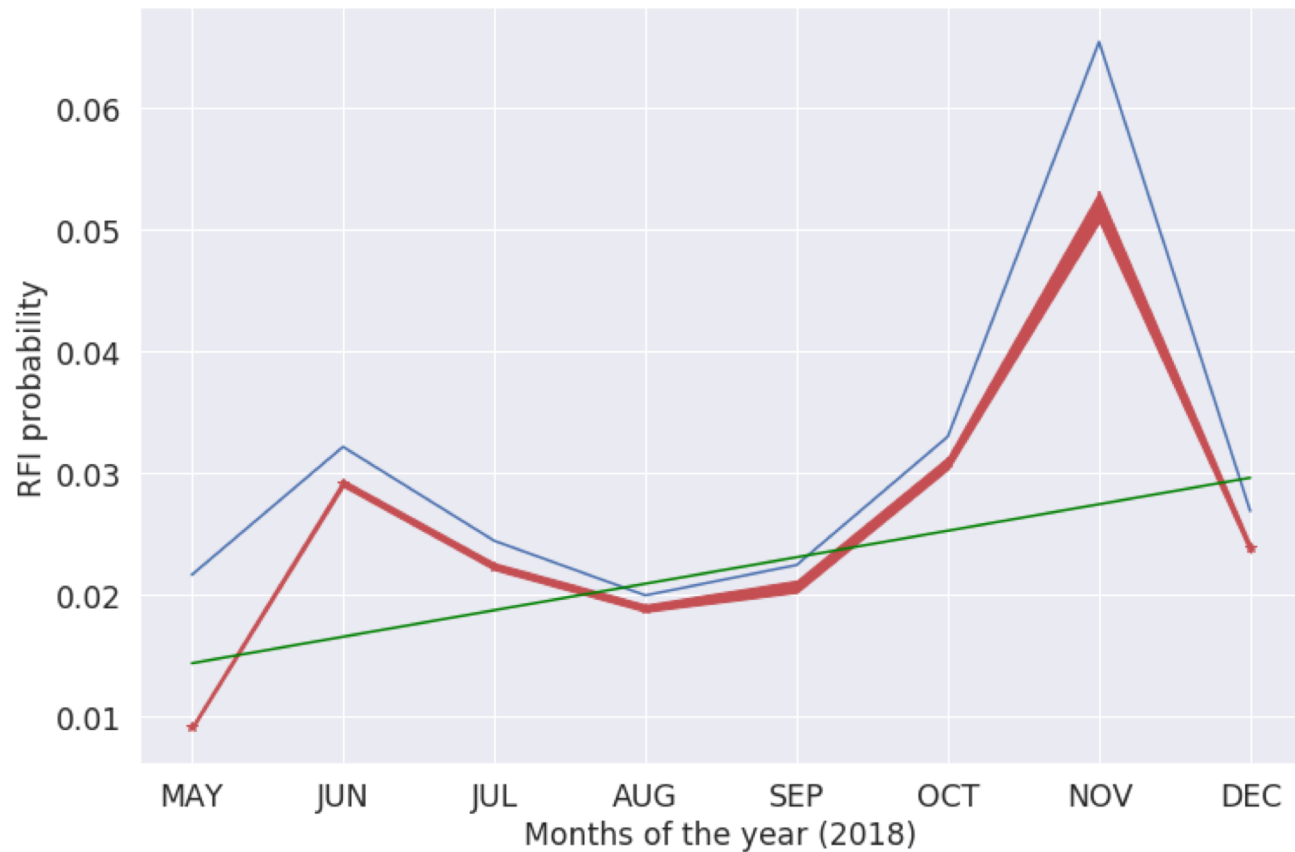
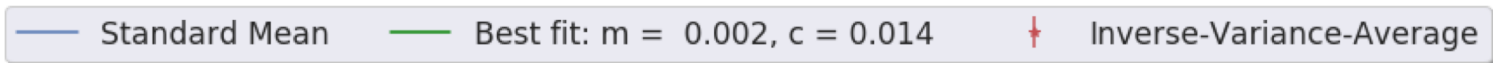
- May - Dec 2018
- 1500 hrs obs
- 400 TB imaging data
- T, F, B, Az, El



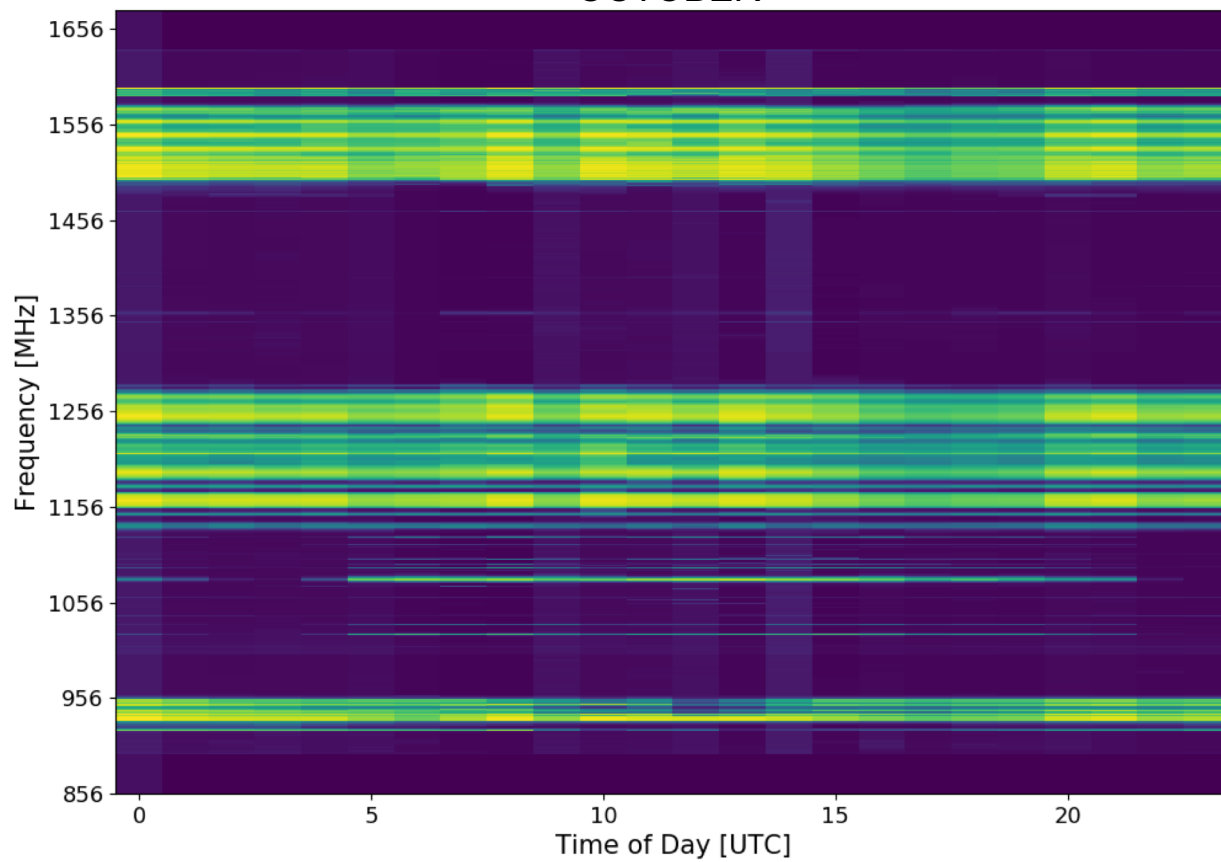




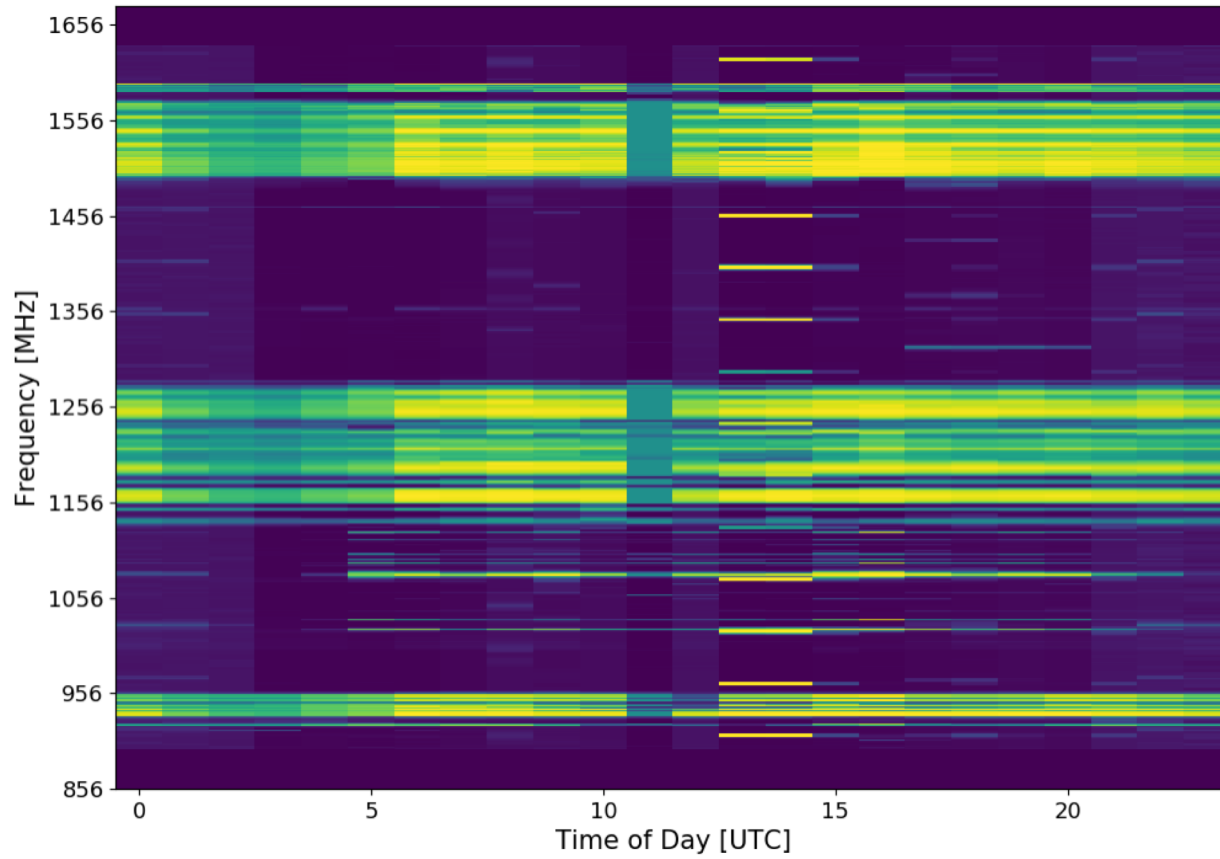
# Time Evolution - RFI probability



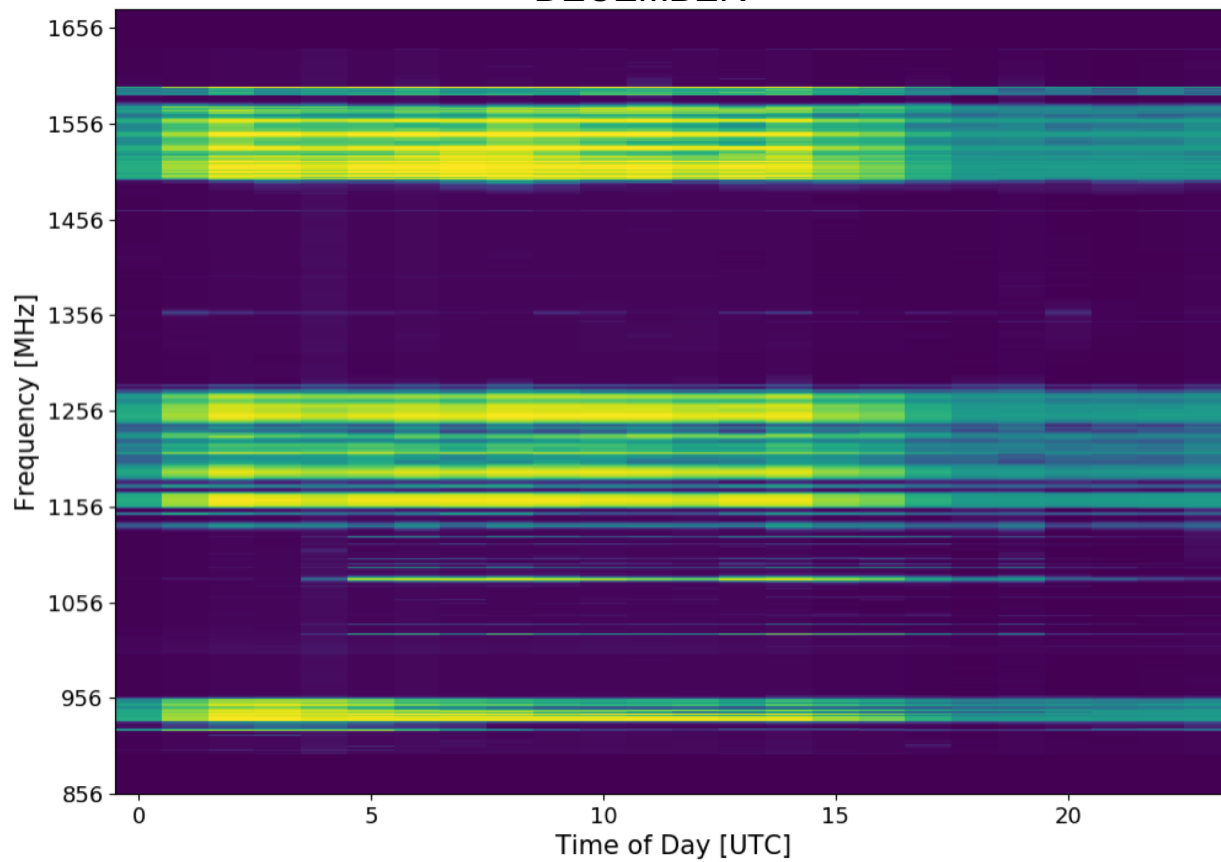
# OCTOBER



# NOVEMBER



# DECEMBER



# Editing RFI

- Bad antennas -> remove them
- Correlator malfunctions -> remove time steps
- Initial pointing delay -> quack first timesteps
- Bandpass issue -> remove channels
- Low elevation -> elevation cut-off
- Interference -> remove antennas, timestep, frequencies or baselines..

Can NOT calibrate in the presence of issues in the data

# Excising RFI

- Manual detection
  - Subjective
  - Enormous data volumes, computationally expensive...
- Automatic Thresholding
  - AoFlagger
  - RFIClean
  - Casa tools

# RFI strategies

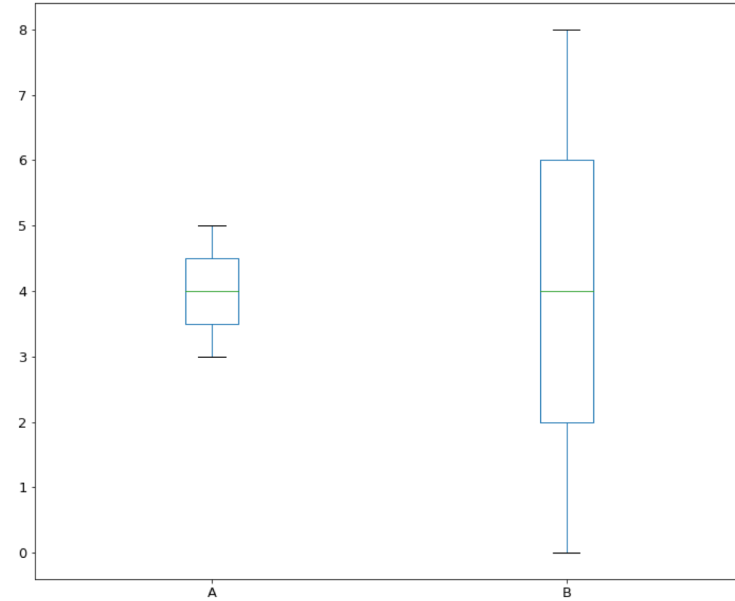
- Online pre-/post-correlation mitigation
  - Computationally expensive
  - Require high time resolution
- Offline
  - Post-optimized, data can be re-ordered and manipulated

# Automated excision

- Detection
  - Find and throw away affected data
  - Filtering or subtracting: estimate RFI contribution and restore affected data
- Detection (“flagging”)
  - Flag regions of bad data and do not use them in calibration and try correction afterwards
- Filtering
  - Hard, resulting data quality is not well understood, computationally expensive.

# Mathematical tools

- Measures of spread
  - Mean, median, skewness and kurtosis
    - What should we use?
    - $A = [3, 4, 5]$
    - $B = [0, 4, 8]$
    - What are the mean?
    - What can be the issue with the Mean?



- Kurtosis

$$\beta_2 = \frac{\mu_4}{(\sigma^2)^2}$$

$\beta_2$  : Kurtosis.

$\mu_4$  : Fourth moment of the probability distribution function.

$\sigma^2$  : Variance of the probability distribution function.

- Median Absolute deviation (MAD)

$$M = \text{Median}(X)$$

$$MAD = \text{Median}|X - M|$$

- Filtering:

$$\sigma_r = 1.4826 * MAD$$

- MAD filter

- sgolay...

# MeerKAT

## RFI **sim**

## RFI **Removal**

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R. Armstrong, B. Bassett, C. Finlay, N.Oozeer



# RFIsim



## What do we want?

- A simulator that produces realistic visibilities corrupted by external RFI sources.
- To know the exact contribution of RFI in each visibility voxel.

## Why do we want it?

- To compare different RFI detection and excision algorithms. Astronomers cannot provide what we need.

# Where are we now?

A MeerKAT simulator that produces realistic RFI-corrupted visibilities.	
A dataset that is being used to train a neural network for RFI detection.	

## **BONUS**

A way to predict satellite positions which has been used to optimize a MeerKAT early science observation proposal.

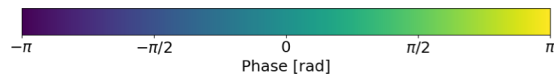
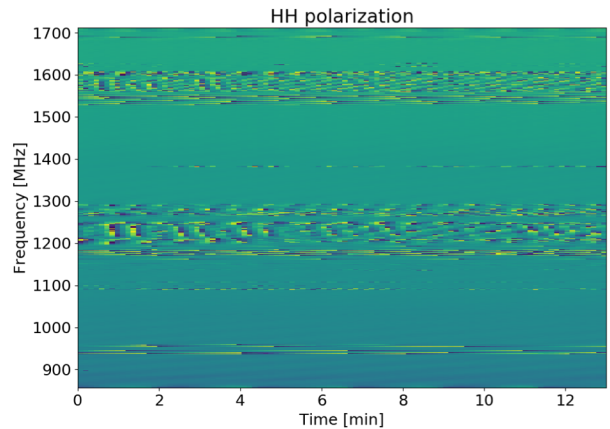
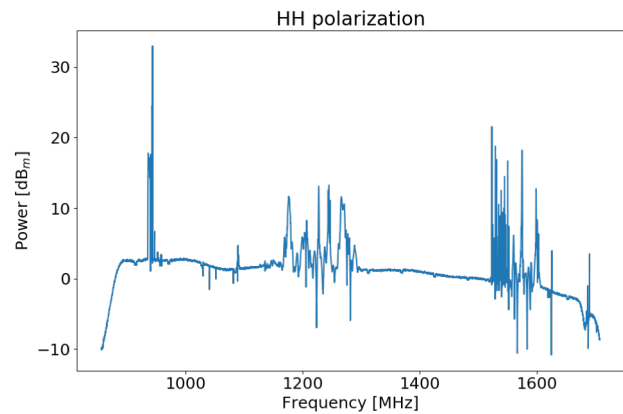
# What effects are included?

Real satellite paths	✓
Ground-based RFI	✓
Realistic RFI characteristics (amplitude and frequency)	✓
Accurate astronomical sky model	✓
MeerKAT (u,v) coordinates	✓
Realistic Bandpass	✓
Realistic primary beam	✓
Time variation (gains, RFI amplitude)	✓

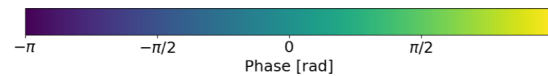
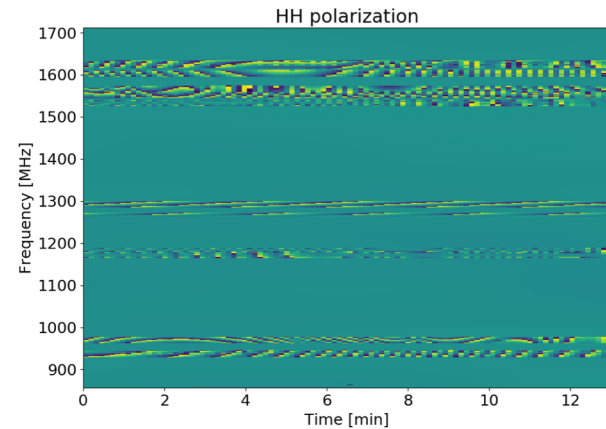
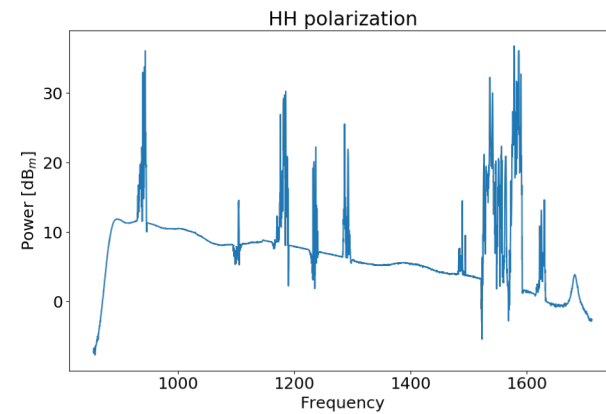


# Results

## Real



## Simulated



# MeerKAT

## RFI detection using

# Deep Learning

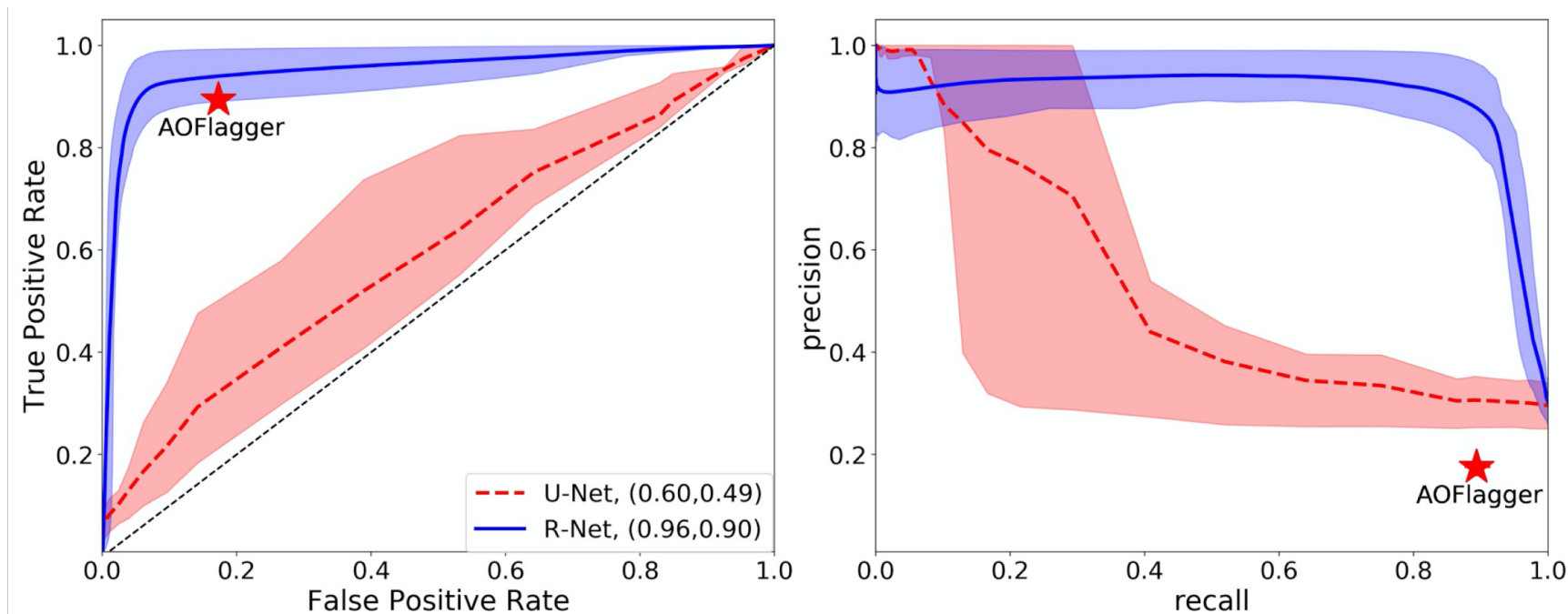
Alireza Vafaei Sadr, SARAQ Data Science Team

B. Bassett, (Y. Fantaye; C. Finlay), N.Oozeer, (R. Ramphul)



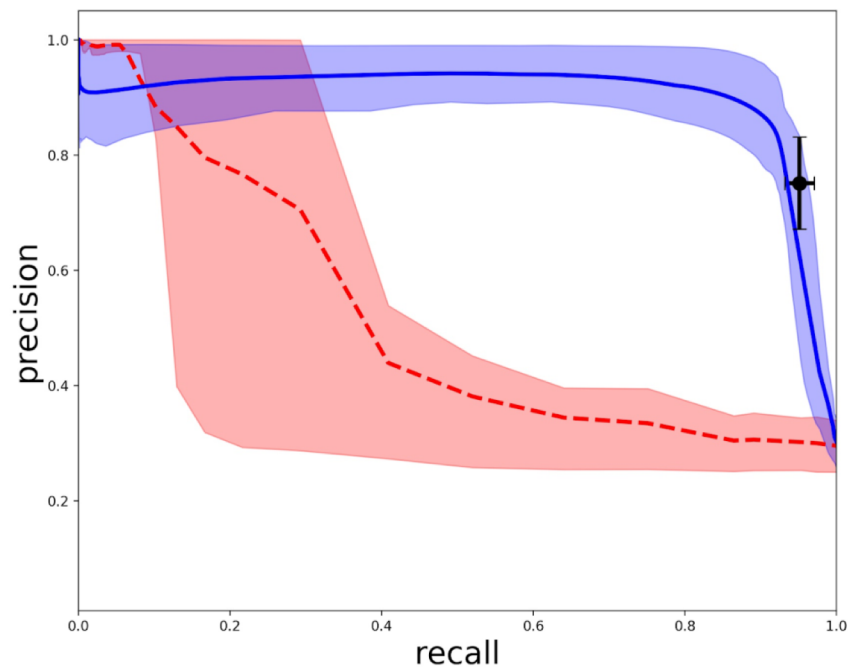
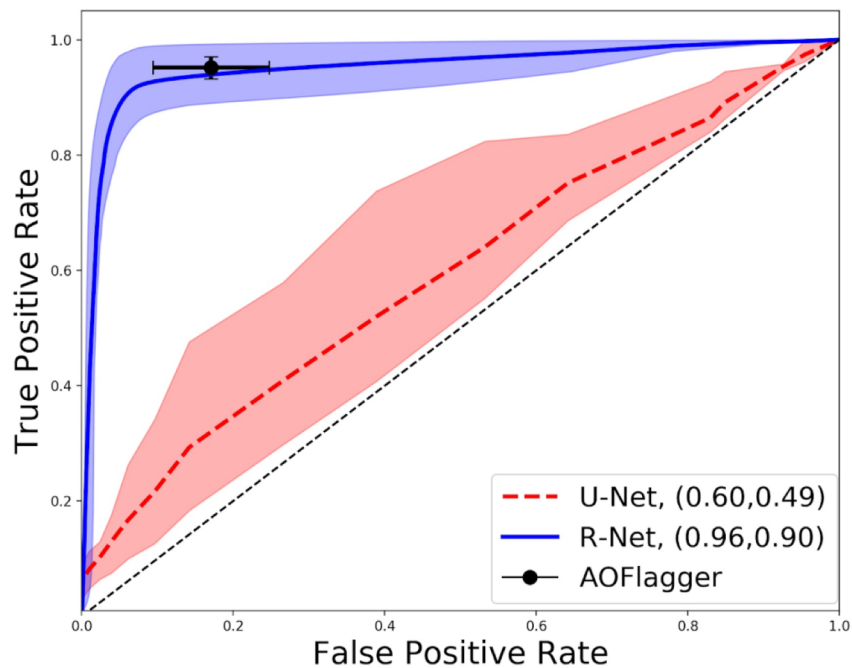
**SARAQ**  
South African Radio  
Astronomy Observatory

# Comparison with Default AOflagger

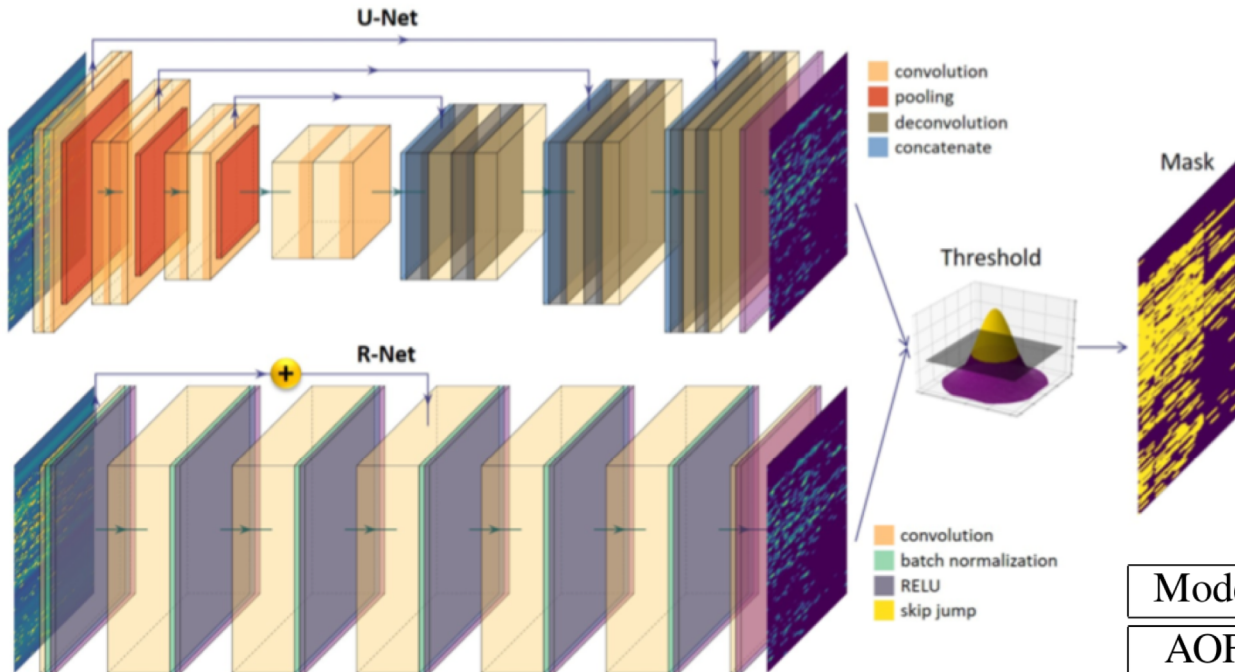


Deep learning is significantly better than AOFlagger unless AOFlagger is extensively tuned (which no one does). Our algorithm tunes automatically, and is flexible (not just a **single** False Positive Rate).

# Comparison with Optimised AOflogger

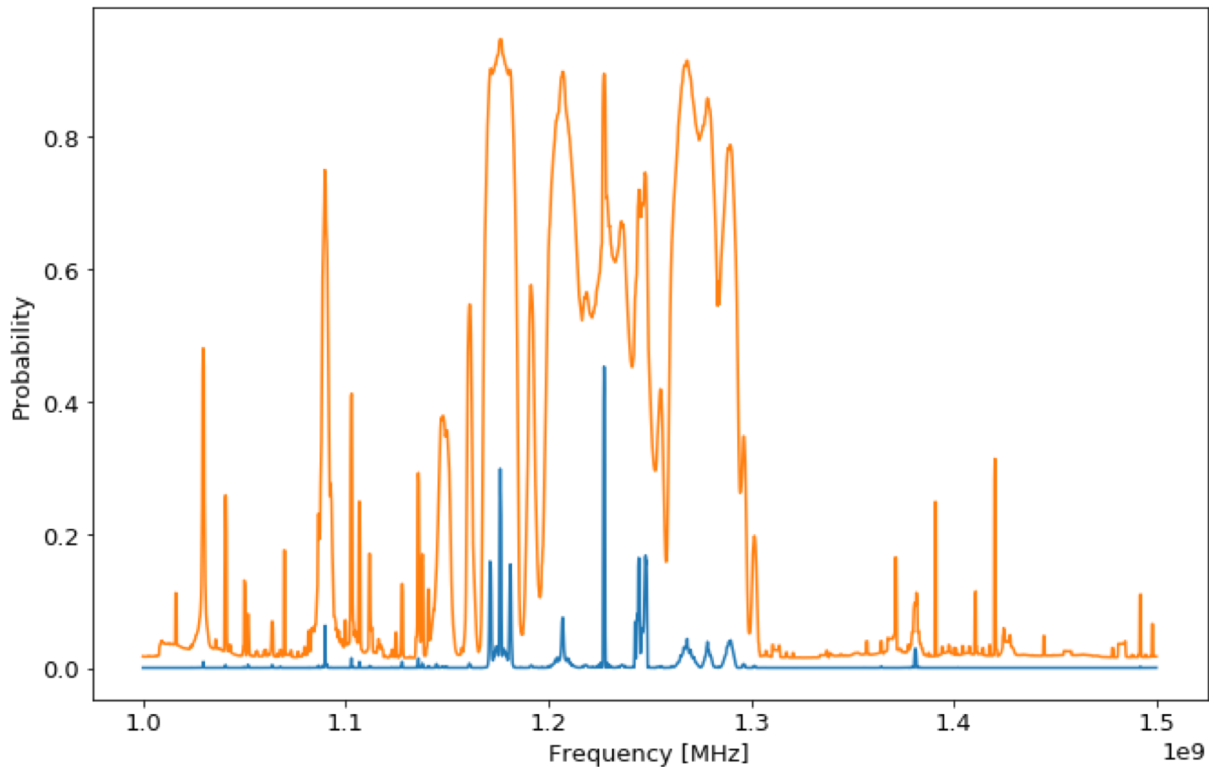


# Network from KAT-7



Model	variables	exe. time(s)
AOFlagger	-	31
R-Net-5	11473	18
R-Net-6	15181	24
Unet-3-16	116770	49
Unet-3-32	465986	53

Frequency Band (MHz)	Shared Service	Scientific Observations
Frequencies below 100MHz		
25.550 - 25.670		RF Bursts (Sun, Jupiter), Interstellar Medium, Steep spectrum sources
37.5 - 38.25	Mobile Communication.	RF Bursts (Sun, Jupiter), Interstellar Medium, Steep spectrum sources
73.0 - 74.6		RF Bursts (Sun, Jupiter), Interstellar Medium, Steep spectrum sources
Frequencies below 1000MHz		
150.05 - 153.0	Mobile Communication except Aeronautical.	RF Bursts (Sun), Interstellar Medium, Steep spectrum sources, Pulsars
322.0 - 328.6	Mobile Communication.	Deutrium line (327.384 MHz), RF Bursts (Sun), Interstellar Medium, Steep spectrum sources, Pulsars
406.1 - 410.0	Mobile Communication except Aeronautical.	RF Bursts (Sun), Interstellar Medium, Steep spectrum sources, Pulsars
608.0 - 614.0	UHF TV Broadcasting.	RF Bursts (Sun), Interstellar Medium, Steep spectrum sources, Pulsars
Frequencies below 5000MHz		
1400.0 - 1427.0	Space Research and Earth Oration-Satellite. (passive)	Galactic and local extragalactic H1 (Hydrogen-1) (1420.406 MHz), Galactic Continuum
1610.6 - 1613.8	Mobile Satellite and Aeronautical Radio Navigation.	OH(Hydroxyl)
1660.5 - 1670.0	Mobile Satellite, Mobile Communication and Space Research.	OH(Hydroxyl)
1710.0 - 1930.0	Mobile Communication.	OH(Hydroxyl)
2655.0 - 2700.0	Mobile Communication except Aeronautical, Mobile Satellite, Space Research and Earth Exploration Satellite.	
4800.0 - 5000.0	Mobile Communication and Space Research.	Formaldehyde, Continuum, VLBI



Outlier detection from Ingest (blue) and flagger outlier(orange) for a RFI free region

- We noticed from the ingest detection that there is around a 17.5% chance of picking an outlier between frequency 1240 MHz - 1250 MHz.
- This corresponds to redshift  $z \sim 0.127 - 0.135$ .
- Any emissions from astronomical sources in this redshift range will inevitably be removed by the ingest and there is no way to undo the excision.

# Conclusion

- Sources of RFI
- How to edit and excise RFI
- Why we need to excise RFI
- What are some of the tools to use
- New algorithm to detect RFI
- The impact of RFI on astronomy