

# Outline

- ▶ Fourier Transforms
- ▶ The Dirac  $\delta$  Function

# Fourier Transform

Fourier Transform of a non-periodic function,  $f(x)$ . If  $g(k)$  is the FT of  $f(x)$ , then

$$g(k) = \frac{1}{2\pi} \int_{-\infty}^{\infty} f(x) e^{-ikx} dx$$

$$f(x) = \int_{-\infty}^{\infty} g(k) e^{ikx} dk$$

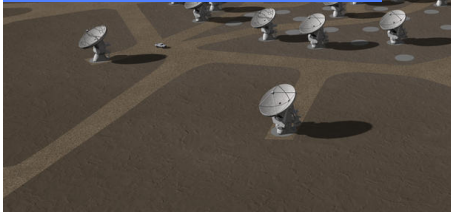
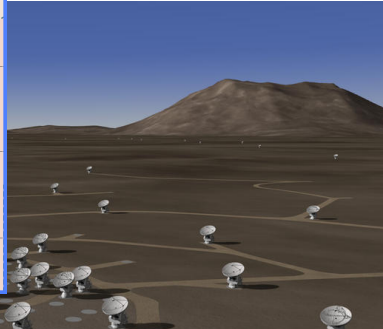
and  $f(x)$  is the inverse transform of  $g(k)$ .

$x$  and  $k$  have inverse units.

If  $x$  is a wavelength,  $k$  is a wavenumber.

If  $x$  is time,  $k$  is frequency.

# ALMA: Atacama Large Millimetre/Submillimetre Array



ALMA



ALMA



[www.alma.ac.uk](http://www.alma.ac.uk)

ALMA



[www.alma.ac.uk](http://www.alma.ac.uk)

# Interferometers

What you measure with two elements of an interferometer is the complex visibility

$$V = \int_{-\infty}^{\infty} I(\underline{\sigma}) e^{i2\pi \underline{b} \cdot \underline{\sigma}} d\sigma$$

where  $I(\underline{\sigma})$  is the distribution of the brightness of the sources on the sky (as described by the direction on the sky by  $\underline{\sigma}$ ).  $\underline{b}$  is vector baseline between the two elements.

$V$  is just the (inverse) Fourier transform of the sky brightness distribution  $I$ . So to determine the structure of the source take the Fourier transform  $V$  measurements.

# Fourier Transform

Transform of a non-periodic function,  $f(x)$ . If  $g(k)$  is the FT of  $f(x)$ , then

$$g(k) = \frac{1}{2\pi} \int_{-\infty}^{\infty} f(x) e^{-ikx} dx$$

$$f(x) = \int_{-\infty}^{\infty} g(k) e^{ikx} dk$$

and  $f(x)$  is the inverse transform of  $g(k)$ .



*Example:*

FT of square pulse is a sinc ( $= (\sin x)/x$ ) function.

*Example:*

Shift theorem for Fourier transforms.

## Dirac $\delta$ -function

$$\delta(x) = 0 \quad \text{for } x \neq 0$$
$$\int_{-\infty}^{\infty} \delta(x) dx = 1$$

By considering the Fourier transform of a Dirac  $\delta$  function and then the inverse of the result

$$\delta(x - x_0) = \frac{1}{2\pi} \int_{-\infty}^{\infty} e^{ik(x-x_0)} dk$$