SADT Technical Information Pack

Rev 2.0, January 2016

Content in this document is approved by SADT and subject to change

Signal and Data Transport Consortium



SIGNAL AND DATA TRANSPORT



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Introducing SKA

SKA Phase 1 has Two Telescopes



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SKA1_Mid 350 MHz – 14 GHz 64 MeerKAT dishes 133 SKA1 dishes.





SKA1_Low 50 – 350 MHz 131,000 aperture array dipole 512 stations of 256 antennas

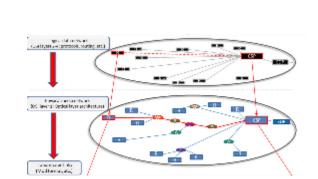
SKA Design Elements



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Low frequency aperture array





Central Signal Processor



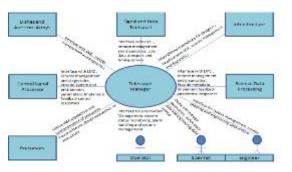
Science Data Processor



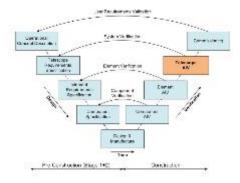
Dish



Infrastructure



Telescope Manager



Assembly, Integration Verification

Signal & Data Transport

SKA Observatory



Australia South Africa SKA Observatory Global Remote Remote Remote stations Headquarters Remote stations Station Station on spiral arms UK Remote Remote Station Station Remote Remote Station Station Host Country Host Country Operations Operations Science Science Central Signal Central Signal SKA1-low Data Processing SKA1-mid Data Processing Processing Processing & Archive & Archive Core Array Core Array RSEC/RDC RSEC/RDC ...

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SKA Uses 3 Network Types



Science Data

CSP-SDP

SDP to world

DDBH

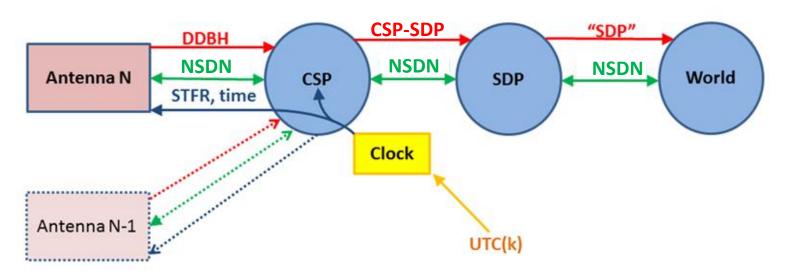
Sync & Timing

- Clock ensemble
- Freq. & Phase
- UTC time

Non-Science Data

- Control & Monitor
- Alarms
- Internet, VoIP

"Spanning" Tasks: Network Architecture; Network Manager; Local Infrastructure



SKA Global Design Consortia



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SKA Global Consortia

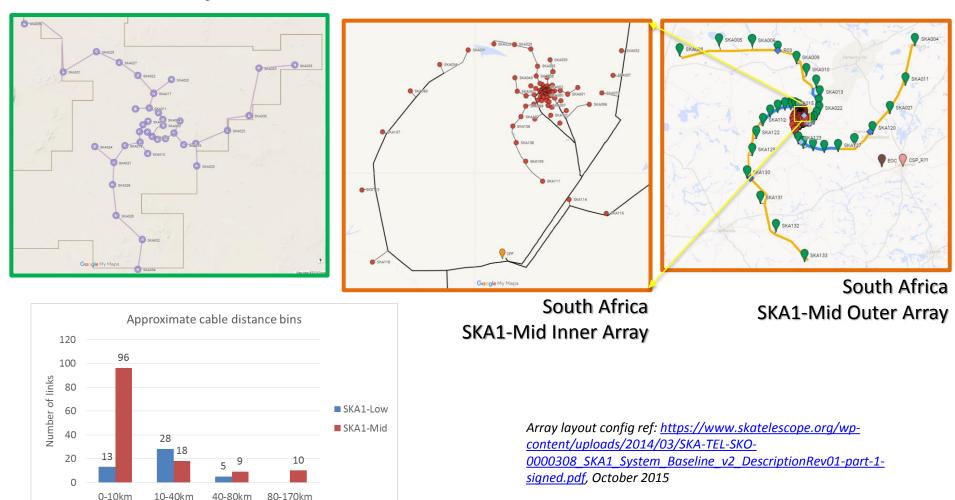


SKA Telescope Sites Remote array stations and cable routing to CPF



Australia SKA1-Low Outer array

Cable distance optimised topology examples

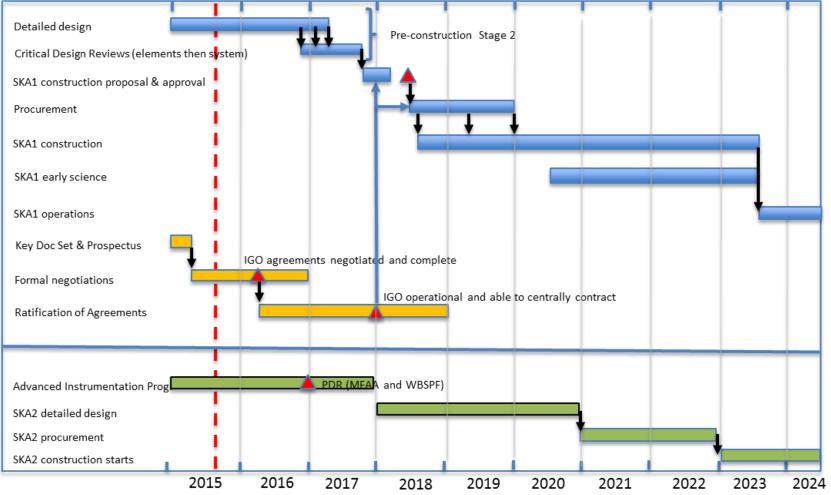


High-level SKA Project Schedule



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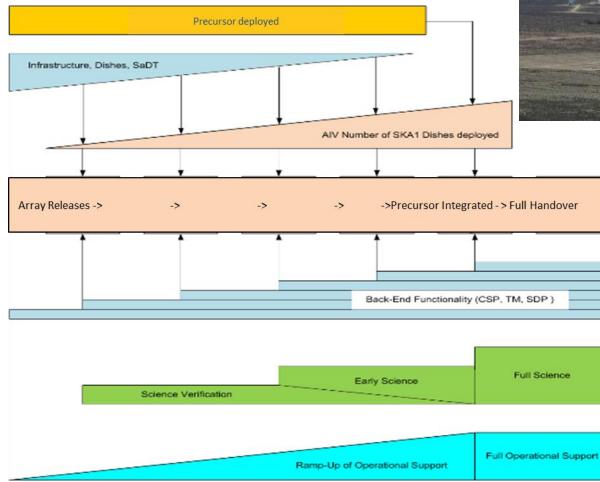


Andrea Casson, August 2015

SKA1 Telescope Roll-out Physical and Functional



AIV roll-out features including consortia input, science experiments, and an operation support



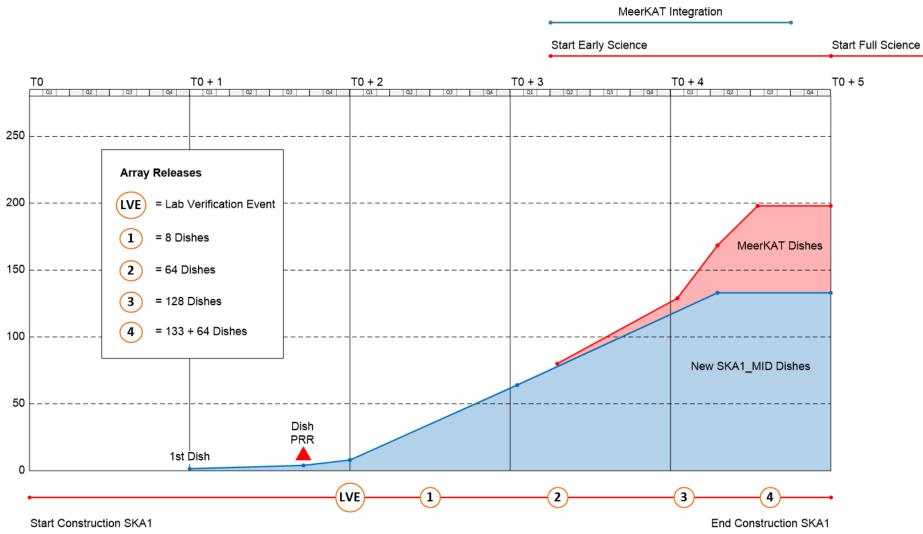


South African core array site with MeerKAT dishes, September 2015

SKA1-MID Dish Roll-Out

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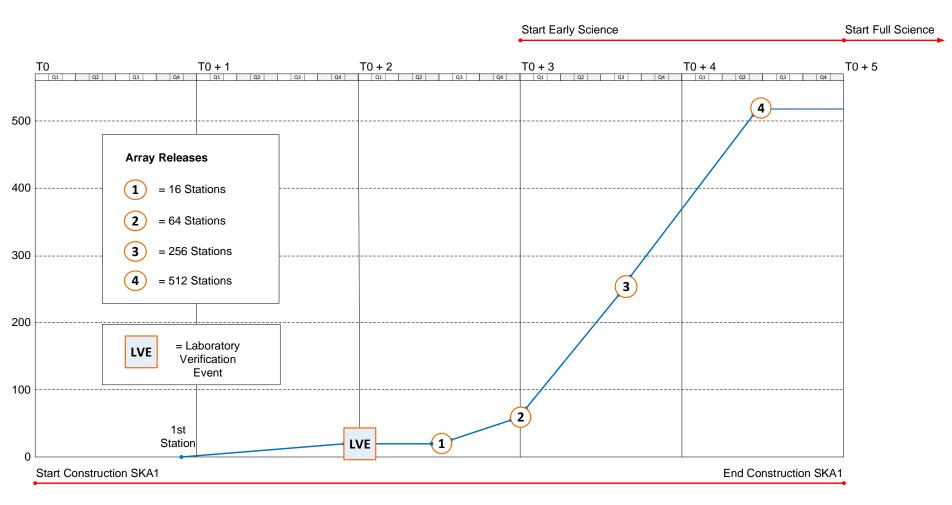
MeerKAT Science Programme



Richard Lord, AIV Consortium, 11 November 2015

SKA1-LOW Station Roll-Out





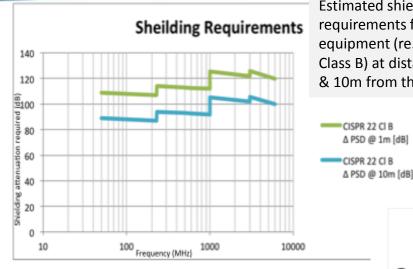
Richard Lord, AIV Consortium, 11 November 2015



- Capital Cost Ceiling, both telescopes: 650M Euro (SADT ~30M Euro per telescope)
- Deployment Period: Years 2018 to 2022
- Power Consumption Cap, SADT Array site equipment: Mid 111kW, Low 51 kW
- Annual Availability
 - Operationally "Available" for 95% of the time
 - Operationally "Degraded" for 50% to 95% of the time
- Environmental Dish & Remote beamforming Station Environmental
 - Outside open air operating temperature -5 C to +50 C
 - Outside open air non-operating temperature -15 C to +60 C
 - Dish pedestal forced air flow, but not air conditioned
 - Operating humidity 40 60%
 - Storage & transport humidity 40 95%
- Max Receptor Baselines Mid 150 km, Low 80 km
 - Approx. max cabled routed distances Mid 160 km, Low 80 km
- Array Synchronisation Clock phase 1ps, Time stamp 10ns

Stringent RFI Shielding Requirements

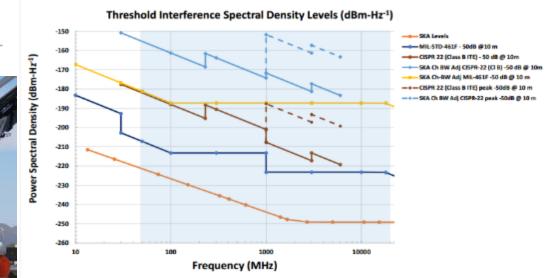


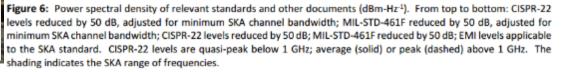


Estimated shielding requirements for COTS equipment (re. CISPR 22 Class B) at distances of 1m & 10m from the RFI culprit

Approx 100dB of shielding requirement to use standard COTS equipment in dish pedestals

Ref: SKA EMI/EMC STANDARDS AND PROCEDURES SKA-TEL-SKO-0000202, Rev 01







Client side Science Observation Data Capacity



 Current <u>client side</u> Ethernet interface estimates, per location and experiment

Data Source Location	Experiment	SKA1-Low	SKA1-Mid
Remote Array Station Egress	All	3 x 40GE**	1 x 100 GE
Central Processing Facility Ingress	All	108 x 40GE	133 x 100GE
Central Processing Facility Egress	Visibilites	48 x 100GE	64 x 100GE
Central Processing Facility Egress	Pulsar Search	8 x 100GE	8 x 100GE
Central Processing Facility Egress	Pulsar Timing	2 x 10GE	2 x 10GE
Central Processing Facility Egress	NSDN	2 x 10GE	2 x 10GE
Science Processing Facilty Egress*	All	1 x 100GE	1 x 100GE

*Network connectivity for science data transport to world-wide regional centres currently out of scope for SKA1

**6 beamformed stations per super-station, calibration beam data also included in client capacity



Introduction to the SaDT The Signal and Data Transport Consortium

Signal and Data Transport (SADT) Consortium

SIGNAL AND DATA TRANSPORT

- Consortium Board, chair: Huib-Jan van Langevelde
- Lead institute: University of Manchester
 - Leader: Keith Grainge
 - Project Manager: Jill Hammond
 - System Engineer: Rob Gabrielczyk
 - Project Engineer: Richard Oberland

Institutions involved in the Signal and Data Transport consortium include :-Commonwealth Scientific and Industrial Research Organisation (CSIRO), Australia

Australia Academic and Research Network (AARNet), Australia University of Western Australia, Australia Tsinghua University/ Peking University, China National Centre for Radio Astrophysics (NCRA) / Tata Consulting, India Persistent Systems, India Joint Institute for VLBI in Europe (JIVE), The Netherlands Instituto de Telecomunicações (IT), Portugal SKA South Africa Nelson Mandela Metropolitan University (NMMU), South Africa Meraka Institute, CSIR, South Africa EM Software and Systems (EMSS), South Africa University of Granada, Spain University of Manchester, UK National Physical Laboratory (NPL), UK GÉANT, UK

• Various industry sub-contractors





SADT Work Packages and Leaders



- SKA.TEL.SADT.MGT, Project management Jill Hammond, Uman
- SKA.TEL.SADT.SE, System engineering Rob Gabrielczyk, Uman
- SKA.TEL.SADT.SAT.CLOCKS, SAT Clock Design David Hindley, NPL
- SKA.TEL.SADT.SAT.LMC, SAT Local Monitoring and Control Rajesh Warange, NCRA
- SKA.TEL.SADT.SAT.STFR, Distribution of UTC and Frequency Paul Boven, JIVE Netherlands
- SKA.TEL.SADT.NWA Network architecture Richard Oberland, Uman
- SKA.TEL.SADT.NMGR Network manager Yashwant Gupta, NCRA
- SKA.TEL.SADT.NSDN Non Science Data Network- Shaun Amy, CSIRO
- SKA.TEL.SADT.DDBH Digital data back haul Richard Oberland, Uman
- SKA.TEL.SADT.CSP-SDP CSP to SDP Data Transmission Richard Hughes-Jones, GÉANT
- SKA.TEL.SADT.SDP-Deliv SDP Interfaces & Data Transmission Richard Hughes-Jones, GÉANT
- SKA.TEL.SADT.LINFRA Local infrastructure Jaco Muller, SKA Africa



SIGNAL AND DATA TRANSPORT

DDBH Digital Data Back Haul



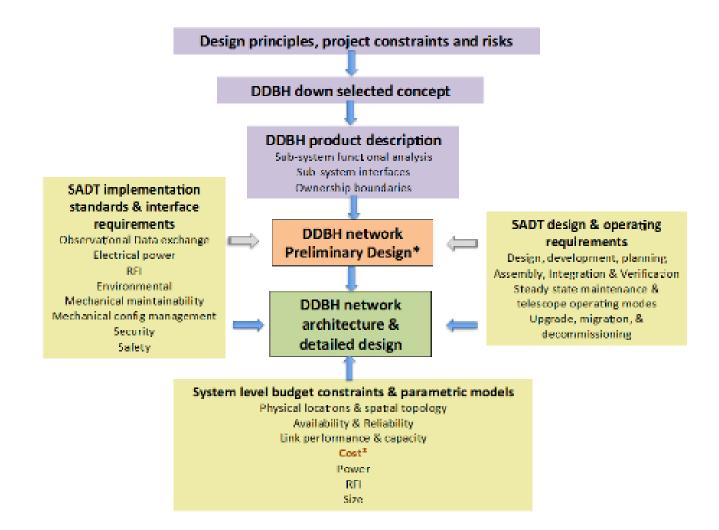
The digital data backhaul networks of the telescope are responsible for transporting digitized science data from the receptor stations to the central processing facility (CPF). Both the SKA1-Low and SKA1-Mid telescopes require their own DDBH node termination equipment and optical fibre infrastructure.

The DDBH will be implemented using Commercial Off The Shelf (COTS) equipment. This approach carries low technical risk, since it builds on the considerable industry experience, however the unique telescope environment requirements must be met.

SADT.DDBH Digital Data Back-Haul DDBH work package design process

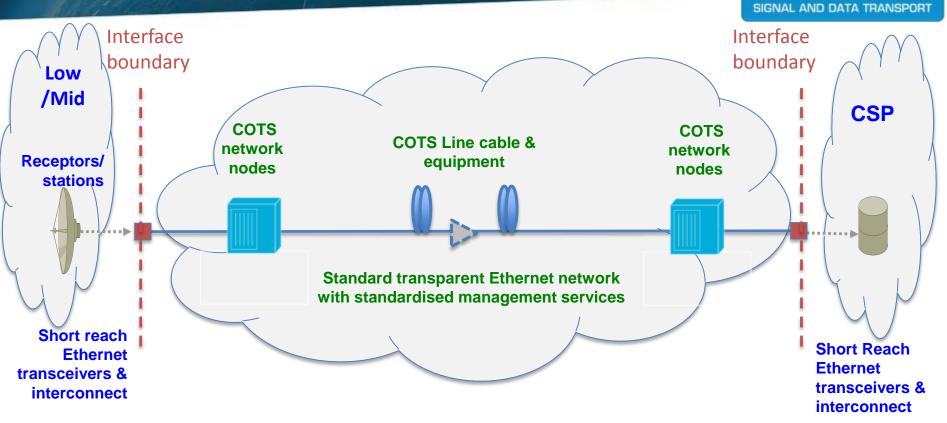


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SADT.DDBH Digital Data Back-Haul DDBH network current design concept



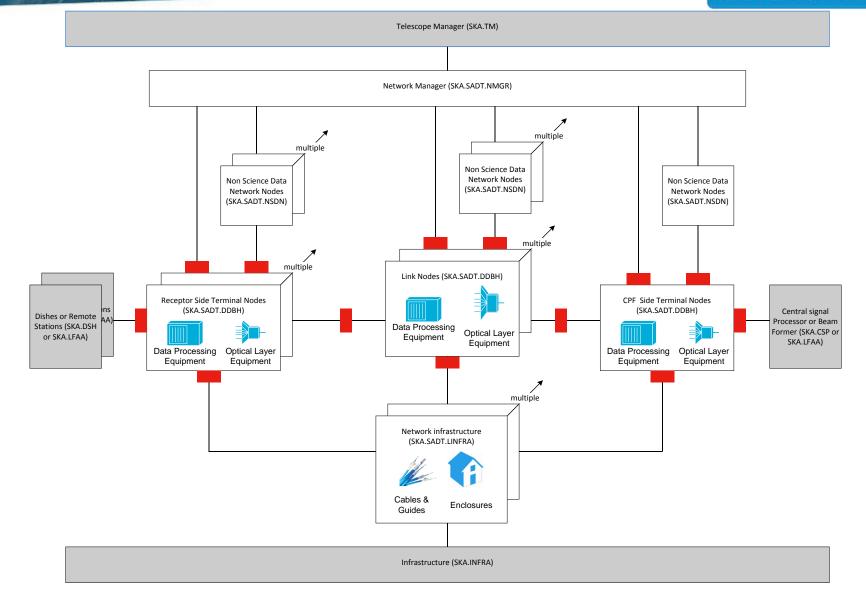


- Fully managed COTS solution vendor agnostic design
- SKA-Mid; 133 dish antennas, 1x100GE transport lanes
 - Passive spans with LR4/ER4 grey optics <u>or</u> amplified/regen spans with DD/Coherent DWDM
- SKA-Low; 36 remote beam formed super-stations, 2 x 100GE transport lanes (3x 40GE clients)
 - Passive spans with DWDM optics

SADT.DDBH Digital Data Back-Haul DDBH network products and interfaces



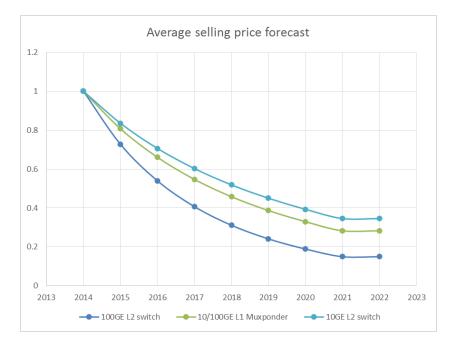
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SADT.DDBH Digital Data Back-Haul COTS cost estimation & market forecasting

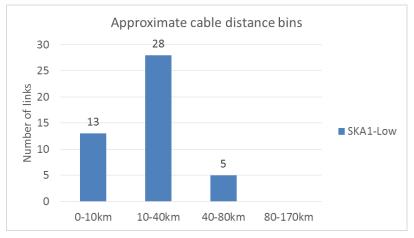


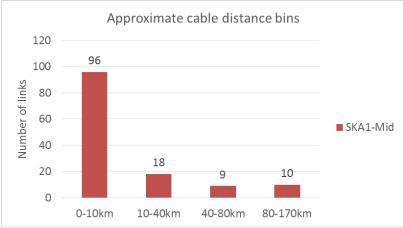
Estimated fractional selling price drop per annum



- Cost and power consumption a function of
 - Year deployed (roll-out)
 - Data capacity
 - Cable distance

Estimated cable routed distance bins





Array layout config ref: <u>https://www.skatelescope.org/wp-</u> content/uploads/2014/03/SKA-TEL-SKO-0000308_SKA1_System_Baseline_v2_DescriptionRev01-part-1signed.pdf, October 2015



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CSP-SDP

Data Transport from Central Signal Processing to Science Data Processing Facilities

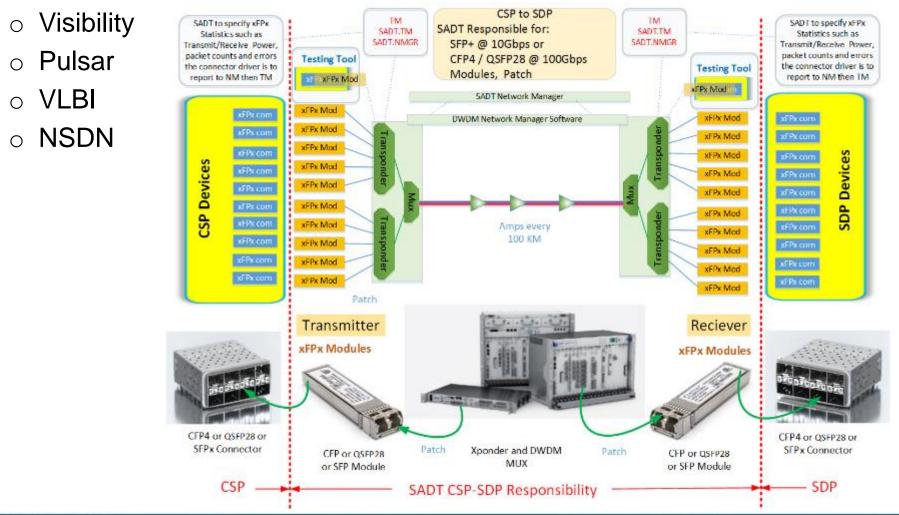


 The SKA 1 CSP-SDP Network provides long distance high bandwidth connectivity between the Central Signal Processor (CSP) locations, and the Science Data Processor (SDP) complex for each of the two observatories. It will carry the science data traffic, and non-science data such as monitoring and control from TM, IP phones, internet access and other auxiliary site traffic

SKA.TEL.SADT.CSP-SDP CSP to SDP transmission CSP-SDP Network Overview

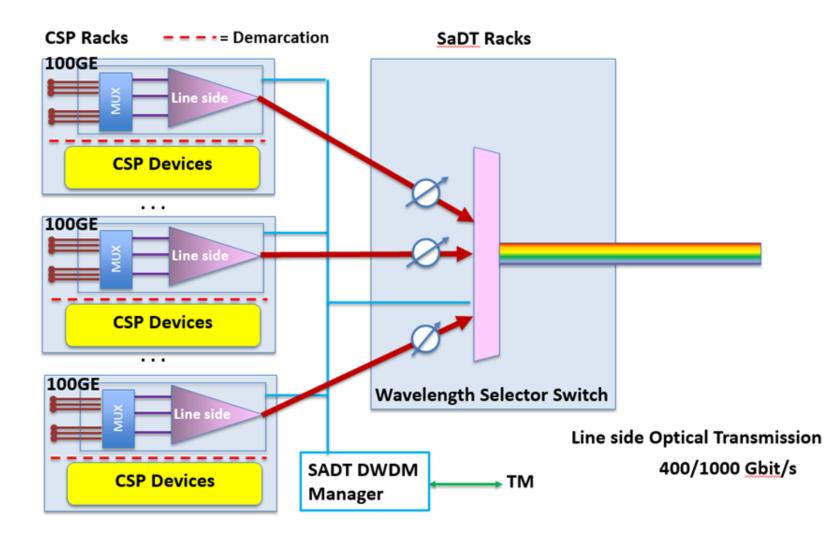


- SIGNAL AND DATA TRANSPORT
- Provides a high bandwidth path from the Correlator to the HPC.
- Carries the following on 100 Gigabit Ethernet channels:



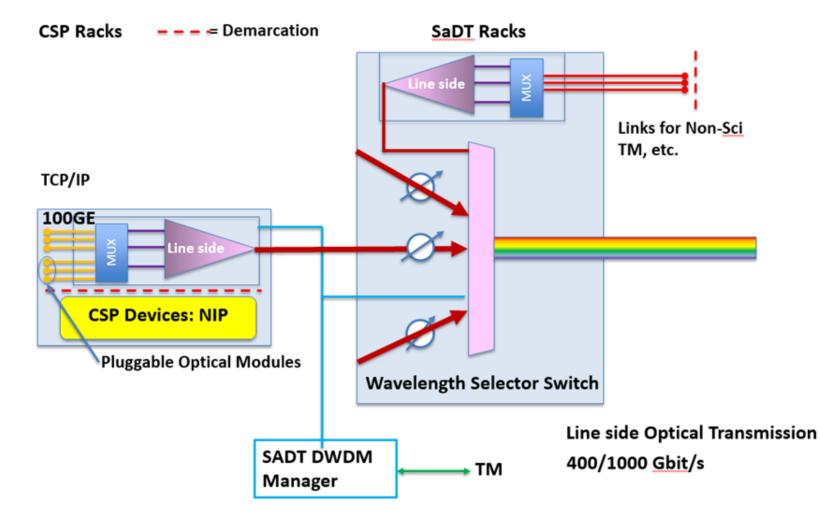
SKA.TEL.SADT.CSP-SDP CSP to SDP transmission CSP Egress: Visibility Data transport concept





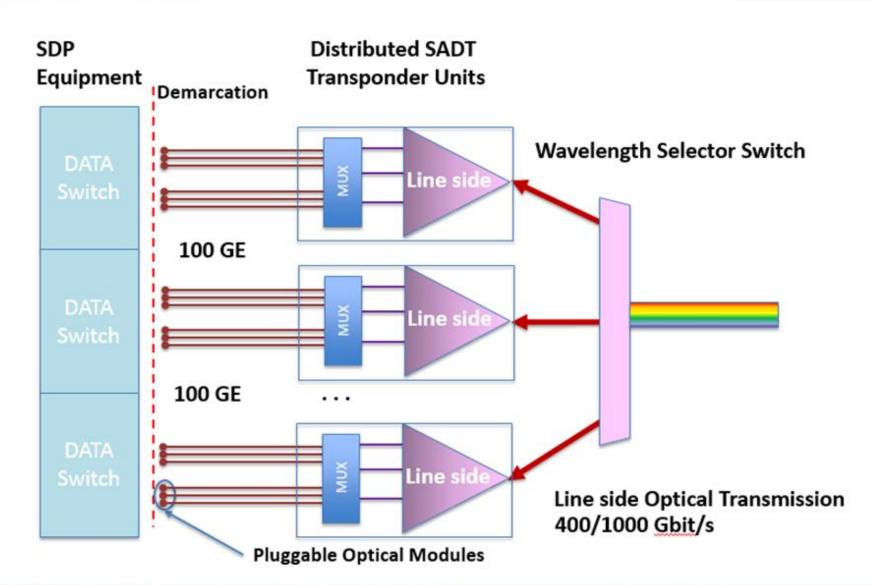
SKA.TEL.SADT.CSP-SDP CSP to SDP transmission CSP Egress: Pulsar & NSDN transport concept





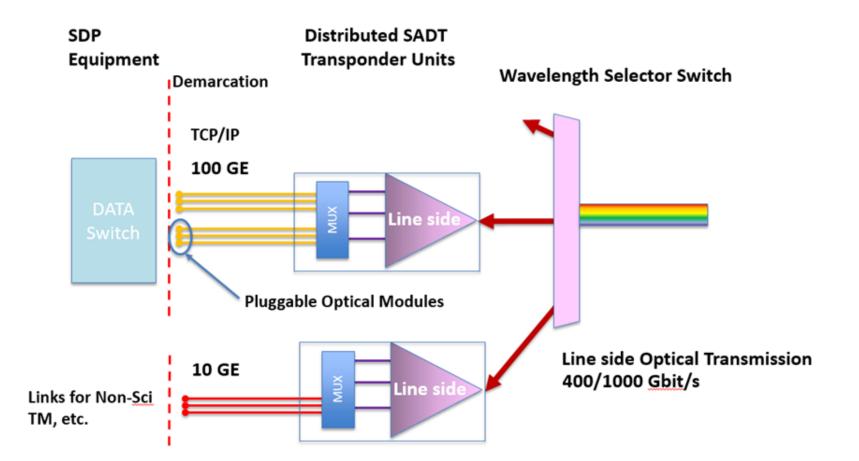
SKA.TEL.SADT.CSP-SDP CSP to SDP transmission SDP Ingress: Visibility data transport concept





SKA.TEL.SADT.CSP-SDP CSP to SDP transmission SDP Ingress: Pulsar & NSDN transport concept





SKA.TEL.SADT.CSP-SDP CSP to SDP transmission Existing MRO to Perth link amplifier huts



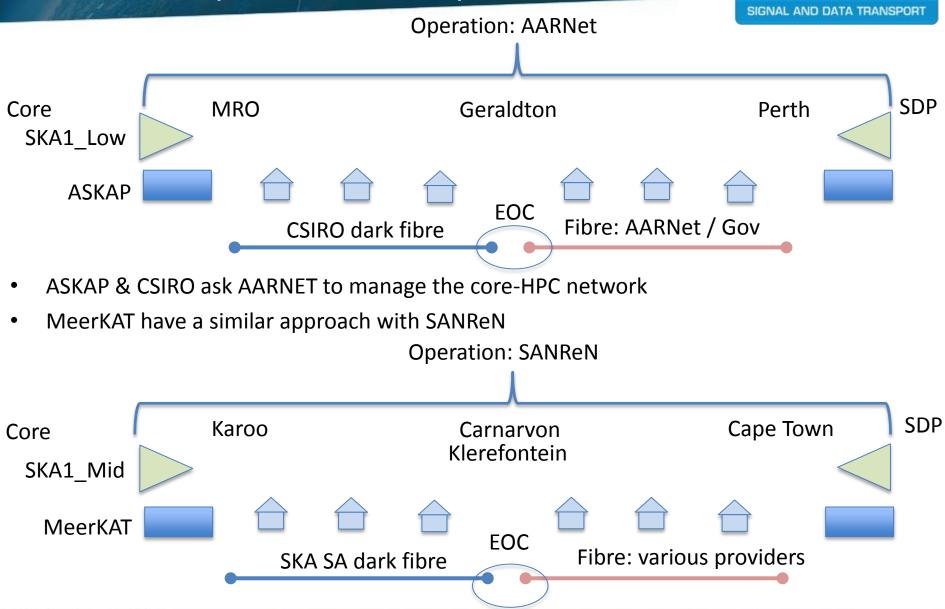
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Solar powered CEV picture (left) and communications rack (right) installed at Geraldton (WA) (courtesy of S. Amy, CSIRO)

- No regeneration required for 820/900 km
- Do need amplifier huts every ~100 km

SKA.TEL.SADT.CSP-SDP CSP to SDP transmission CSP-SDP link operations concept





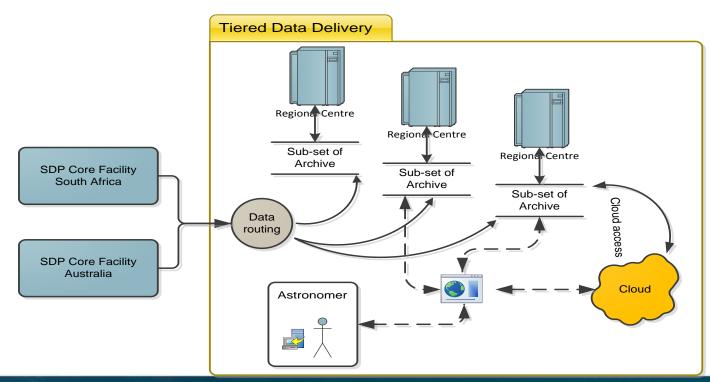
SDP-Deliv Connecting the Observatory to the Astronomy Community



This work package is responsible for the definition of an interface description between the output of the SDP and the Outside World. SADT and SDP consortia propose that this work be extended to include consideration of the full data delivery to the radio astronomy community..



- SDP HPC processing places the SKA data in local archives.
- Model to provide access to the astronomy community:
 - Having a replica is a basic requirement.
 - $\circ~$ Only move the data once.
 - \circ Protocols must be suitable for high bandwidth and real-time transfers.

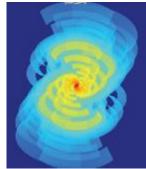


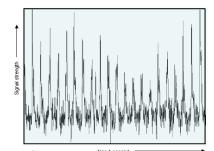


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- Image cubes of sky maps 18PB n*/year ~30Gbit/s
 - $\circ~$ SDP Fourier transform of CSP time series
 - Moderate external compute and model fitting
- Epoch of Re-ionisation 1.6PB/ 6 Hr ~600 Gbit/s
 - Uses calibrated aperture plane
 - Enormous compute of power spectra
- Relativity Gravitational lensing 70PB/2500Hr 60Gbit/s
 - Uses further processed aperture plane data
 - Considerable compute of galaxy elipsciity
- Pulsars 4-5 PBytes/y ≤10Gbits
 - Discovery and in depth study; timing 10 ns in 10 years
 - Large physics compute
- VLBI we do 4 Gbit/s now !
 - \circ Data direct from the correlator ~ 10 Gbit/s UDP

Bottom Line: Lots of data, work in progress to constrain it – what about 100 Gbit/s

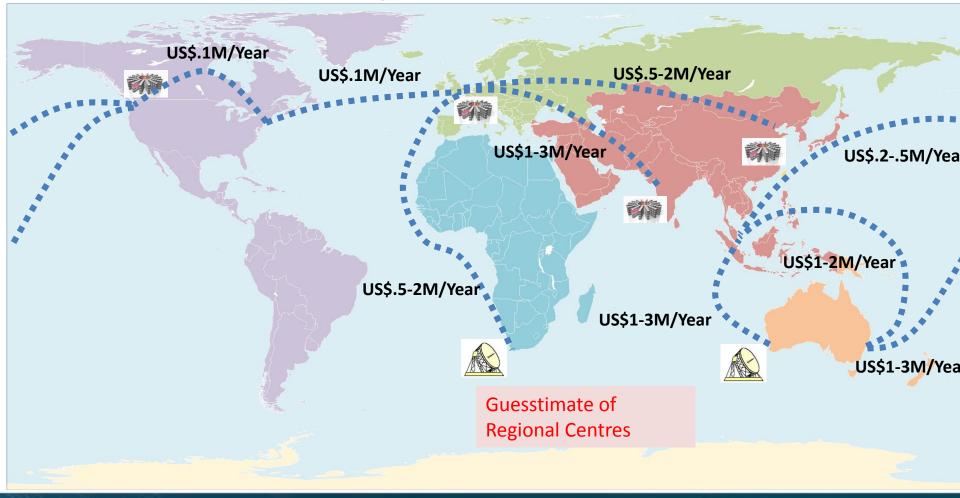




SKA.TEL.SADT.SDP-Deliv Estimated SDP to world costs



- 10 year IRU per 100Gbit circuit 2020-2030
- Guesstimate of Regional Centres locations





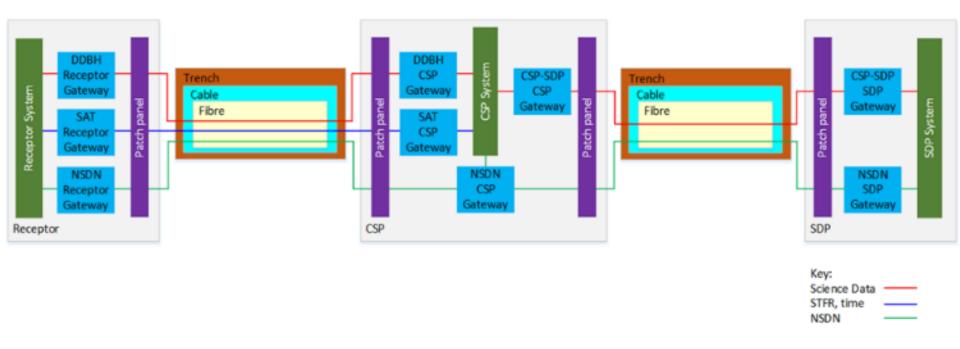
SIGNAL AND DATA TRANSPORT

NWA SaDT Network Architecture

SADT.NWA Network Architecture



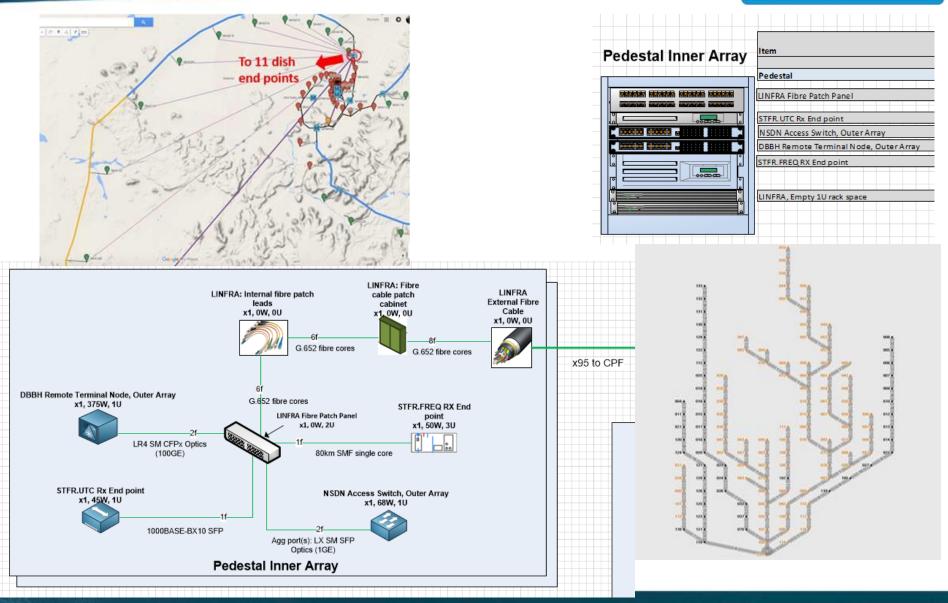
 The signal and data transport network architecture defines the top level physical and functional aspects of the network capable of carrying all the network services required by the telescope.



SADT.NWA Network Architecture Rack configuration and spatial modelling







SADT.NWA Network Architecture Blackbox specifications & allocations



Blackbox Product Types

LINFRA External Cable fibre cores LINFRA Power strip LINFRA Fibre Patch Panel STFR.UTC Tx End point, CPF **STFR.UTC** Repeater STFR.UTC Rx End point STFR.FREQ RX STFR.FREQ RX & Amplifier STFR.FREQ TX, CPF NSDN Access Switch, Outer Array NSDN Access Switch, Inner Array NSDN Distribution switch, Outer Array NSDN Distribution switch, Inner Array NSDN Regeneration/Repeater, Outer Array NSDN Core switch. CPF DBBH Receptor Node, L1 Transponder (up to 300km) DBBH Receptor Node, L1 Transponder (up to 10km) DBBH Receptor Node, L2 Switch (up to 10km) DDBH Link Node (EDFA Amplifier Pair) DDBH Link Node (Optical Coupler/WDM)) DBBH CPF Node, L1 Transponder (up to 300km) DBBH CPF Node, L1 Transponder (up to 10km) DBBH CPF Node, L2 Switch (up to 10km) CLKS Clock System CLKS, NTP Server NMGR M&C System

SAT.LMC PC Central Controller SAT.LMC PC104 CLOCKS Controller SAT.LMC PC104 UTC Controller SAT.LMC PC104 FRQ Controller

Blackbox Product Specs

Total service throughput per optical channel, Gbit/s WDM channels aggregated per fibre Max number of external nodes served, >300m External cable interfacing fibre cores Fibre Type Power consumption, W Rack footprint, U External cable facing port reach, km Quantity of black boxes per rack config (e.g. CPF, Inner array stations)

Location dependent SADT Specifications

Loc ID Config ID Lat Long Direct distance to CPF supplied Nearest node distance supplied Nearest distance node ID Power Demanded (W) Rack Space Demanded (U) Total LINFRA ext cable fibres supplied



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LINFRA Local Infrastructure for SaDT



The fibre optic network is the physical layer across which all of the SKA1 telescope networks operate – Science Data, Telescope Manager, Synchronisation and Timing, Non Science Data network, plus a number of auxiliary networks such as the Building Management System

SKA.TEL.SADT.LINFRA Local Infrastructure



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Installation of optic fibre during construction of the ASKAP antennas and associated infrastructure in 2011. Image courtesy Shaun Amy/CSIRO

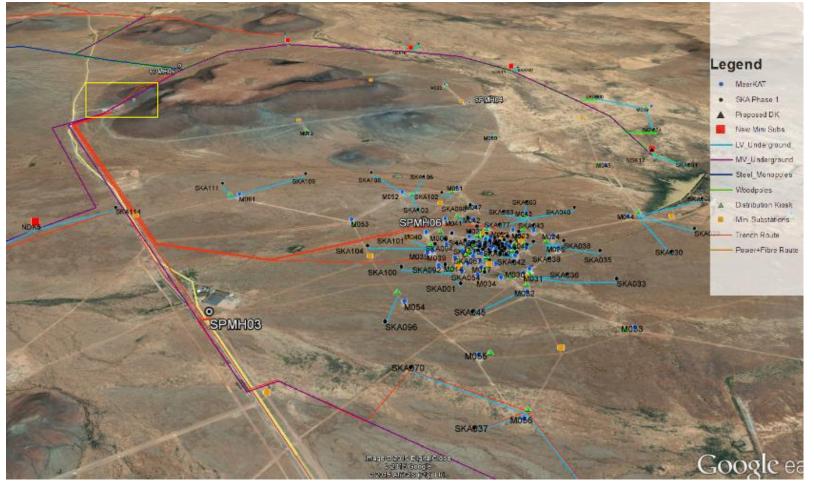


Karoo Array Processing Building and associated assembly infrastructure in 2015. Image courtesy Richard Oberland/Uman

SKA.TEL.SADT.LINFRA Local Infrastructure SKA1-Mid Core array reticulation concepts



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- Fibre cable; power cable; trenching; manholes; ducts; drawpits
- In-house and off-the-shelf spatial modelling tools for spatial optimisation
- Obvious saving combine fibre with power reticulation

SKA.TEL.SADT.LINFRA Local Infrastructure Rack mounted Equipment Specs



- CONFIG NUMBER
- DESCRIPTION
- TRAY HEIGHT (INCL CLEARANCE) REQUIREMENT (U)
- TRAY DEPTH (INCL CLEARANCE) REQUIREMENT (mm)
- MASS (kg)
- EMI Standard
- POWER CONSUMPTION (W)
- THERMAL DISIPATION
- POWER SOCKET TYPE

- FUSE RATING REQUIREMENT (A)
- AIRFLOW DIRECTION REQUIREMENT (COLD TO HOT)
- FIBRE CONNECTORS LOCATION
- POWER CONNECTOR LOCATION
- EARTH STRAP LOCATION
- PRODUCT LABEL LOCATION
- STATUS INDICATORS LOCATION
- MOUNTING DIRECTION (INSERT)
- MOUNTING SUPPORT
- MOUNTING METHOD





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NSDN Non-Science Data Network



The purpose of the Non-Science Data Network (NSDN) is to transmit and receive throughout the system the monitoring and control information from the Telescope Manager (TM) element, and the general communications (Auxiliary Network, AUX) traffic. The NSDN is an observatory-site wide infrastructure that carries a number three sets of services to the various locations required to operate and maintain the telescopes and observatory. These services carry 'live' observation critical data; testing, diagnostic and commissioning data; all other monitor and control information, and the general purpose communications traffic (e.g. IP telephony).



- Deliberately location agnostic:
 - clearly there will be some implementation differences but where possible strive to implement the same design in both Australia and South Africa.
- Utilise in-depth knowledge of the existing precursor networks.
- Core, Distribution (where applicable), Access
- MPLS Core/Backbone:
 - standard industry practice (e.g. campus/enterprise/ISPs)
 - VPNs (layer 2 and layer 3)
- Avoiding "specials":
 - e.g. every pedestal is the same



- Have taken a Total Cost of Ownership view.
- Industrial Ethernet switches system trade-off required.
- Operational Model needed to consider maintenance (and what level) and self-sparing
- Network operations overhead with multiple vendors. Need to minimise!
- Security:
 - pragmatic view
 - "use" cases
- Learn from the precursors

SADT.NSDN Non-Science Data Network Service and location Matrix



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TM Related Flows												
System Configuration data												
States & Modes												
Alarm Signals												
Calibration Data												
Execution Status												
Schedule												
MeerKAT Proxy data												
Science calculation / data												
Calibration Transforms												
Data Products												
Pipeline												
Administration												
Security signals												
safety signals												
QA												
Asset Information												
Internet / Enterprise												
External Input												
Guest services (hard & soft)												
Enterprise services												
Videoconferencing												
Voice												
Internet services												



No data flow required

General data communication services

Production and Engineering Service Data Flows

Engineering Service Data Flow only

SADT.NSDN Non-Science Data Network Services, Characteristics, Locations



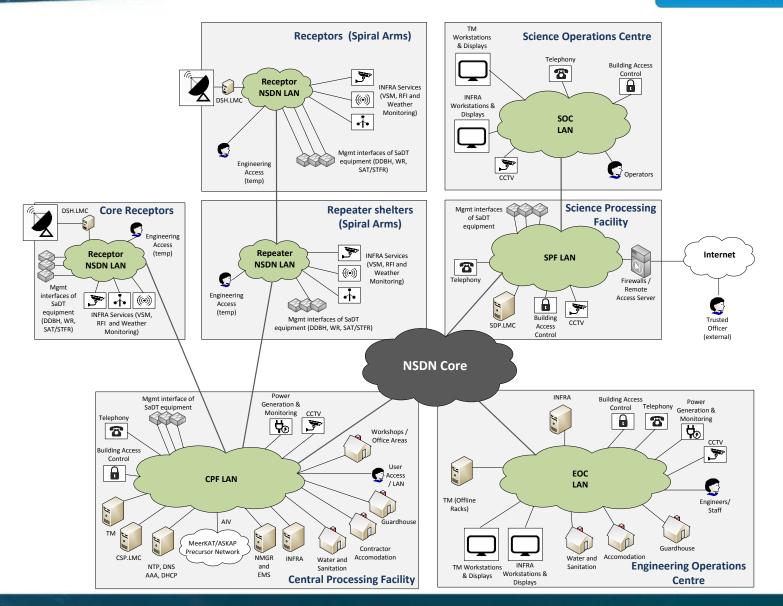
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Network Locations	Service Characteristics	Service	Description					
Power system locations	Service Name	System Configuration data	Monitor and control transactions with specific associated values. Quasi-static data associated with Observatory assets. This will also include serial numbers.					
, Water & sanitation locations	Service End points	Asset Information						
SADT intermediate network	Service Description	Safety signals	Commands and associated values that directly affect or influence safety.					
locations Dish / LFAA Remote Station	Does the service have any critical redundancy requirements	States & Modes	Monitor and Control information altering the state or observing mode of the instrument.					
Locations	requirements	Alarm Signals	Alarm data is a warning signal indicating abnormal conditions exist in the system.					
Site Processor Facility	What time to repair does the	Security signals						
Trusted Offices	service require		Physical security monitoring and control. E,g, Building security door monitors All data associated with the MeerKAT precursor telescope handed to SKA					
Science Proc Centre	What is the average service bandwidth requirement	MeerKAT Proxy	Telescope systems via defined interfaces.					
Science Ops Centre	What is the peak service bandwidth	Calibration Data	The process whereby the output of a measurement is related back to the value of the measure and, in order that absolute measurements are possible.					
Eng Ops Centre	Physical Port Speed required Number of Ports	Calibration Transforms						
Accomodation	Is VLAN trunking required to the service end point	Execution Status Schedule	Defined observing times and objects					
Contractor Accomodation	Number of Vlan per port - if	QA	Quality Assurance					
Admin Buildings	required Port Type	Data Products	A dataset, which when combined with other datasets providing spectrally, temporally and/or spatially resolved measurements of phenomena of					
Site Monitor locations	Transceiver package cable connector		astronomical interest or of sectors of the celestial sphere					
	Distance to Local Element Multicast required	External Input	Data entering the SKA System from trusted external sources.					
	Power over Ethernet (POE)	Pipeline	Pipeline code for the processing of astronomy data.					
	IP Version IP Addressing Protocol	Guest Services	Data associated with guest services provided by the obsertvatory. These include both services associated with custom experiments and visitors to Observatory sites.					
	QOS - What Packet Loss can the service tolerate What is the end to end latency	Enterprise Services	Observatory central functions that apply across the telescope and host country facilities.					
	requirement	Video Conferencing	Packet based Enterprise Video & Audio conferencing capabilities and facilities					
	Port Security requirements What is the maximum packet payload the network needs to	Voice	Voice data and signalling carried over a packetised network. Commonly referred to as VoIP.					
	support	Internet Services	External data accessed from the global network.					

SADT.NSDN Non-Science Data Network NSDN Spatial Network Distribution

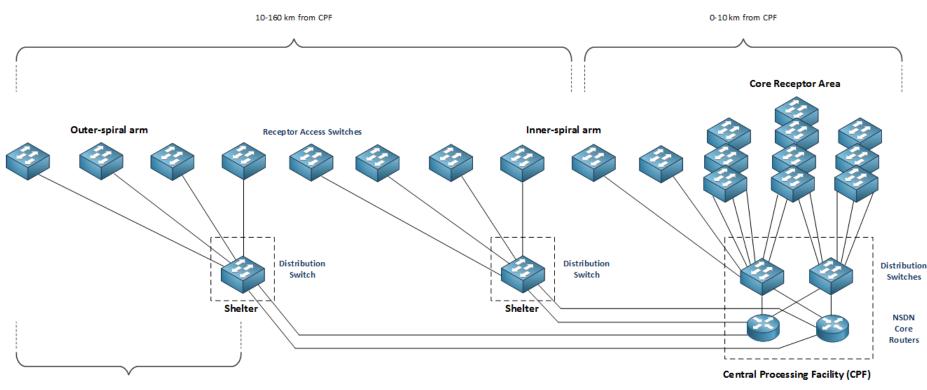


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SADT.NSDN Non-Science Data Network Array Architecture concept



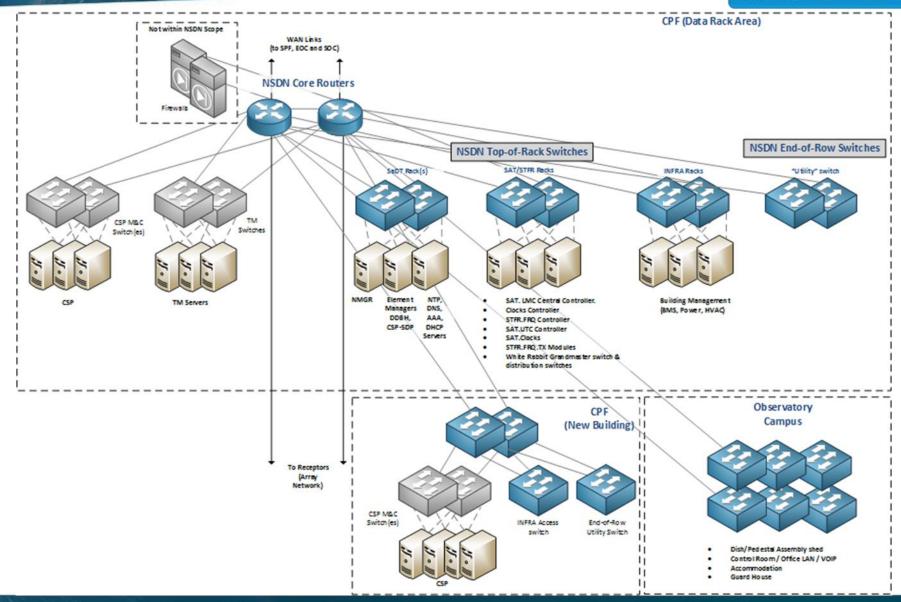


Max 80 km

SADT.NSDN Non-Science Data Network Preliminary Detailed Design concept at CPE

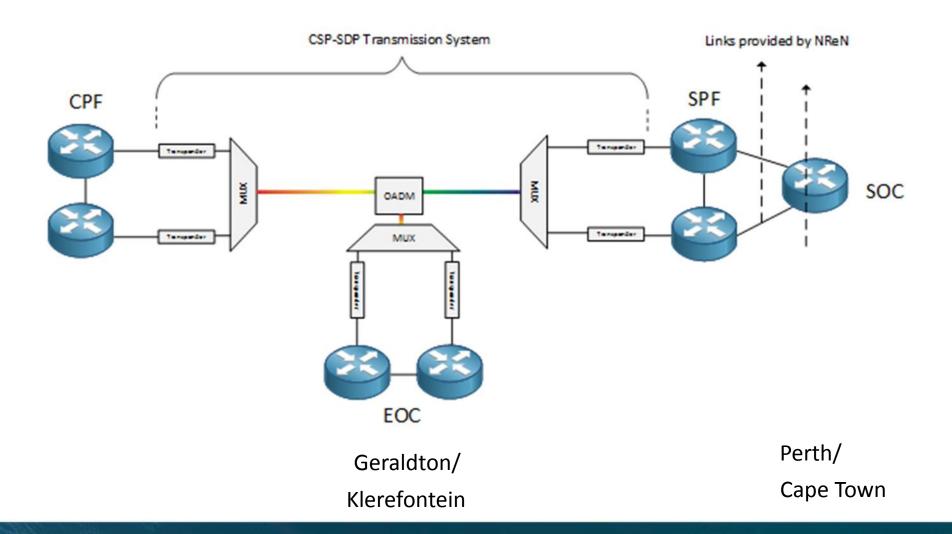


SIGNAL AND DATA TRANSPORT





Similar approach for Australia and South Africa





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NMGR SADT Network Manager

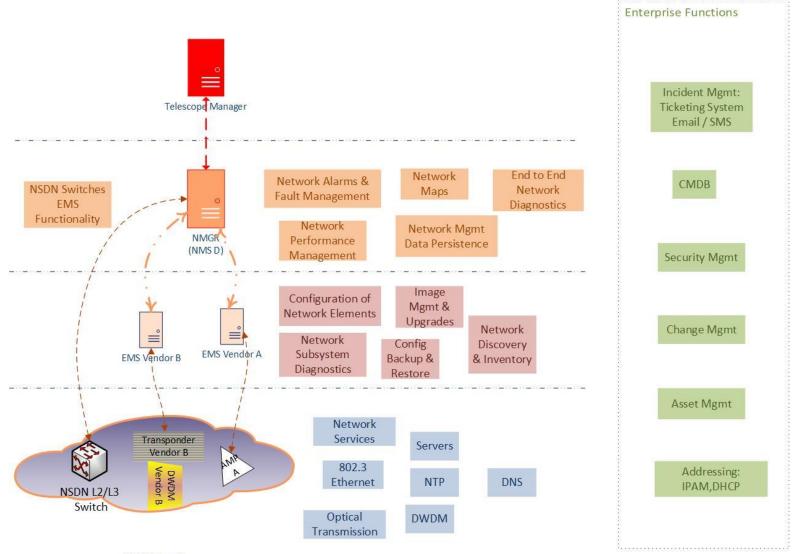


The SADT.NMGR sub element is responsible for monitoring and controlling the DDBH, NSDN, CSP-SDP networks and their associated infrastructure at the receptor array, CPF, EOC, SOC and SDP.

The role of network management includes the configuration and setup of all the equipment (including hardware and software elements), monitoring health and performance parameters, optimizing the behaviour of the network and detecting and responding to adverse situations. Network monitoring and control interfaces are provided to operators and engineers, as well as to the Telescope Manager element. Engineers are provided with additional interfaces to facilitate troubleshooting, reconfiguration and upgrades.

SKA.TEL.SADT.NMGR Network Manager Management Layers

SIGNAL AND DATA TRANSPORT



SADT Network

SADT.NMGR Network manager Functionality In Scope for SADT NMGR



- Fault Management
 - Reception and processing of SNMP traps, Syslog messages, ICMP pings, etc.
 - Performing diagnostics
 - Performing monitoring for network devices, servers and applications.
- Performance Management
 - Polling regularly by SNMP (or other standard protocol) for performance counters.
 - Rolling up counters at predefined intervals
 - Storage of performance counters

Configuration Management

- Inventory management
- Network maps
- Configuration backup and restore
- Configuration audits & compliance reports
- Configuration templates & bulk configuration
- Change automation
- Change notifications
- Scheduled tasks
- Power cycle network element
- Startup / shutdown card / port
- Image file management
- Software / firmware upgrade
- Other Requirements for NMGR
- Interfacing between SADT network and Telescope Manager in accordance with the SADT TM ICD & the LMC Common Interface Guidelines
- Performance requirement as per the L1 alarm latency requirement
- L1 availability requirement for SADT and the Telescope



SIGNAL AND DATA TRANSPORT

SAT.CLOCKS Synchronisation & Timing: Clock Design



This work-package covers provision of the SKA reference timescale . There will be two similar realisations of the SKA reference timescale, one in South Africa and one in Australia. A robust design is proposed for each timescale, based on the use of three active hydrogen masers as the reference clocks.

SKA.TEL.SADT.CLOCKS Clocks design Clock System overview





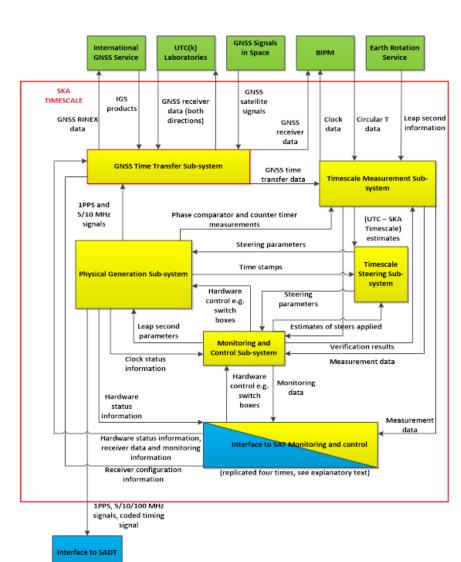
Hydrogen Maser clock example ,at NPL UK

- Requirements:
 - Phase coherence of array
 - \rightarrow accuracy = 1ps
 - Long-term timing for pulsars
 - \rightarrow 10ns over 10 years
 - "3 cornered hat" H-masers
- GNSS link to UTC
- \rightarrow SKA time
- Baseline: 3 instances
- Stringent environmental requirements
- Incorporate existing site masers (TBC)

SKA.TEL.SADT.CLOCKS Clocks design SKA Timescale block diagram



SIGNAL AND DATA TRANSPORT



SKA TIMESCALE: BROKEN DOWN INTO SUB-SYSTEMS



SIGNAL AND DATA TRANSPORT

SAT.STFR Synchronisation & Timing : Station Time and Frequency Reference



The SAT.STFR (Synchronisation and Timing – Station Time and Frequency Reference) carries out the function of delivering standard frequency references and timing signals derived from the SKA frequency and timescale reference (SAT.CLOCKS) to the receptors and other locations. The reference signals must be delivered with sufficient precision to enable each of the telescopes

i. to be phase-coherent

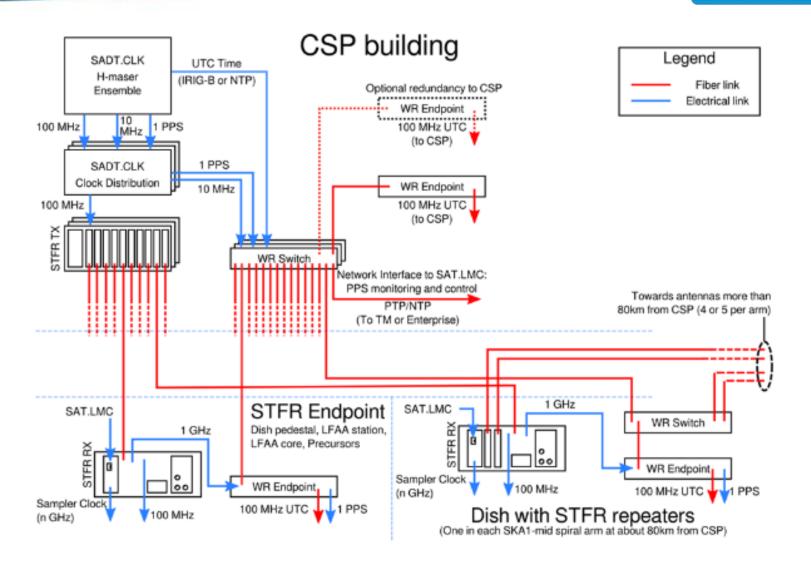
ii. to support the demanding requirements on absolute time for scientific observations such as pulsar timing on timescales of decades
iii. to distribute absolute time for system management, antenna pointing, beam steering, time stamping of data and producing regular timing ticks

- iv. to provide frequency standards for digitizer clocks
- v. To provide time and frequency accurate enough for VLBI observations.

SADT.SAT.STFR Station Time & Freq Reference Frequency & time distribution overview



SIGNAL AND DATA TRANSPORT





DIGINAL AND DATA THANGPORT

SAT.STFR Synchronisation & Timing : STFR.FRQ Frequency Distribution

SADT.SAT.STFR.FRQ Frequency Distribution Concept A: System diagram & performance



System schematic

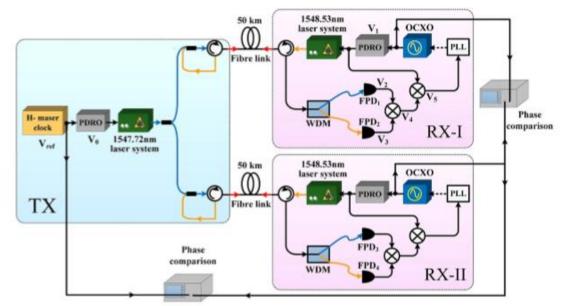
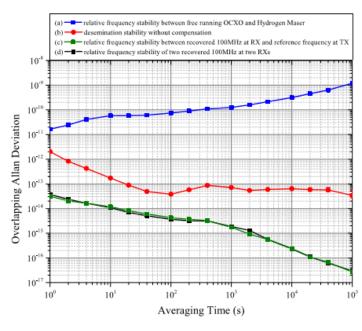


Figure 1. Schematic diagram of the client-side, 1f-2f actively compensated frequency dissemination system. PDRO: phase-locked dielectric resonant oscillator. OCXO: oven-controlled crystal oscillator. WDM: wavelength-division multiplexer. FPD: fast photodiode.

Ref: "Square Kilometre Array Telescope—Precision Reference Frequency Synchronisation via 1f-2f Dissemination", 2015, www.nature.com/scientificreports/

Figure 2. Results of relative frequency stability measurements. (a) Relative frequency stability between the free-running OCXO and H-maser clock. (b) Measured frequency stability of the dissemination system without compensation. (c) Relative frequency stability between the recovered 100-MHz signal at RX and V_{ref} at TX with the PLL closed. (d) Relative frequency stability of the two recovered 100-MHz signals at the two RX sites with both PLLs closed. Curves (c,d) were measured simultaneously.

Performance

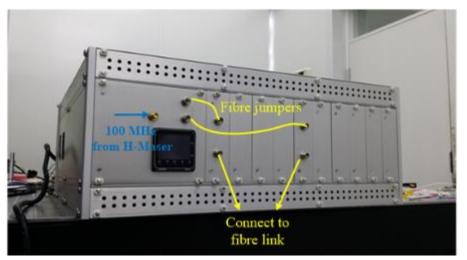


SADT.SAT.STFR.FRQ Frequency Distribution Concept A: STFR Prototypes for field testing

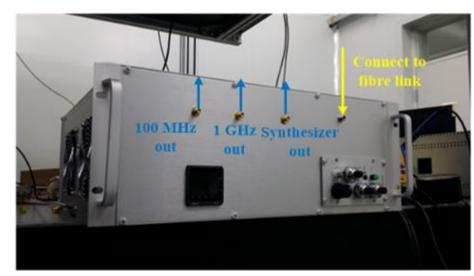


SIGNAL AND DATA TRANSPORT

Concept Prototype Transmitting Module



Concept Prototype Receiving Module



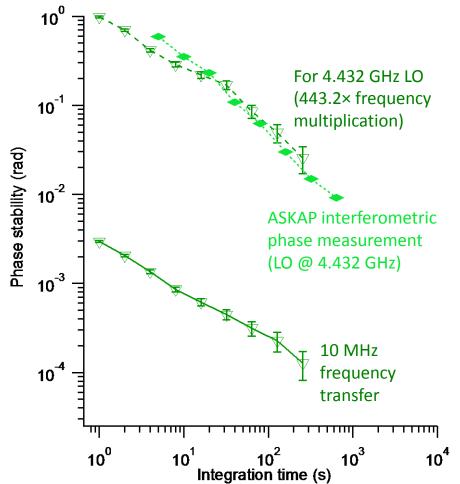
SADT.SAT.STFR.FRQ Frequency Distribution Concept B: Astronomical Verification with ASKAP

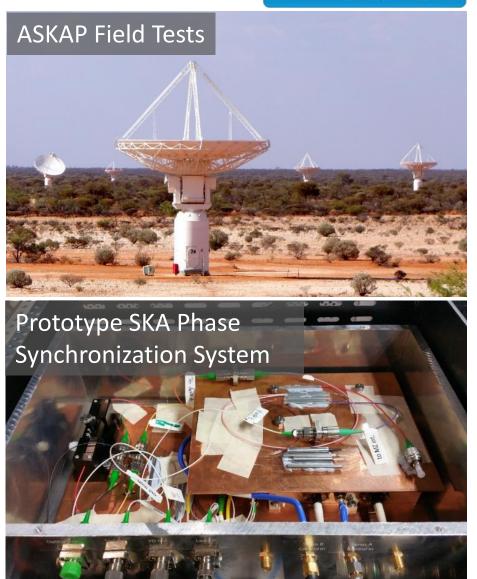


SIGNAL AND DATA TRANSPORT

Measured SKA phase synchronization system optical fibre transfer stability

Astronomical verification of stability





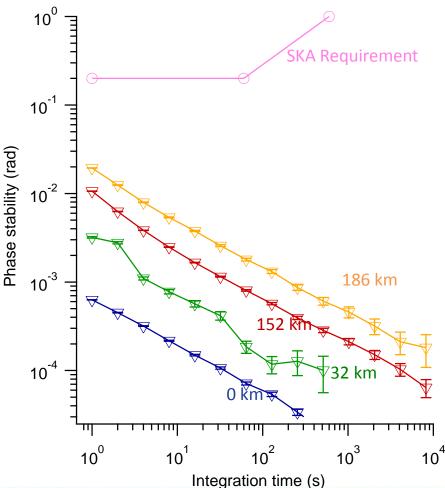
SADT.SAT.STFR.FRQ Frequency Distribution Concept B: South Africa Overhead Fibre Tests

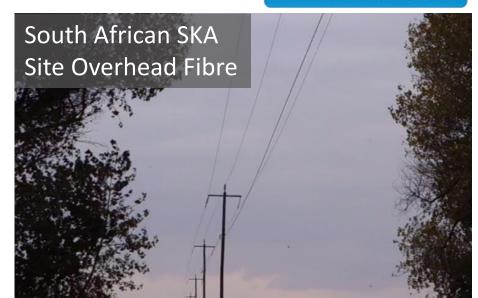


SIGNAL AND DATA TRANSPORT

Measured SKA phase synchronization system optical fibre transfer stability

• On overhead optical fibre links





Prototype SKA Phase Synchronization System

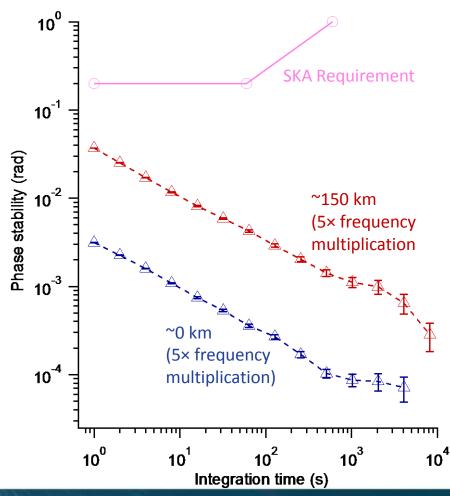
SADT.SAT.STFR.FRQ Frequency Distribution Concept B: Design to Manufacture Development

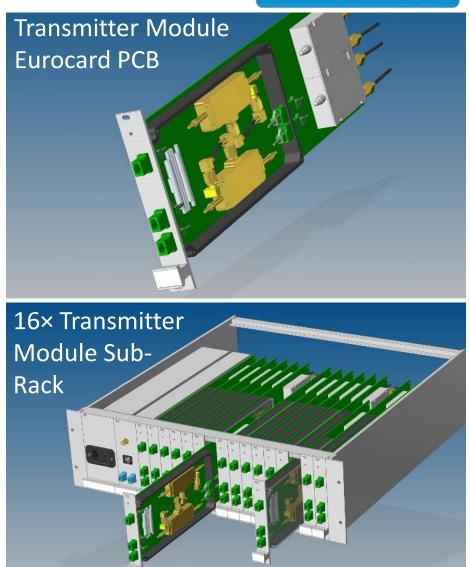


SIGNAL AND DATA TRANSPORT

Expected SKA phase synchronization system optical fibre transfer stability

• Short and long link examples







SAT.STFR Synchronisation & Timing : STFR.UTC Time Distribution

SADT.SAT.STFR.UTC Time Distribution

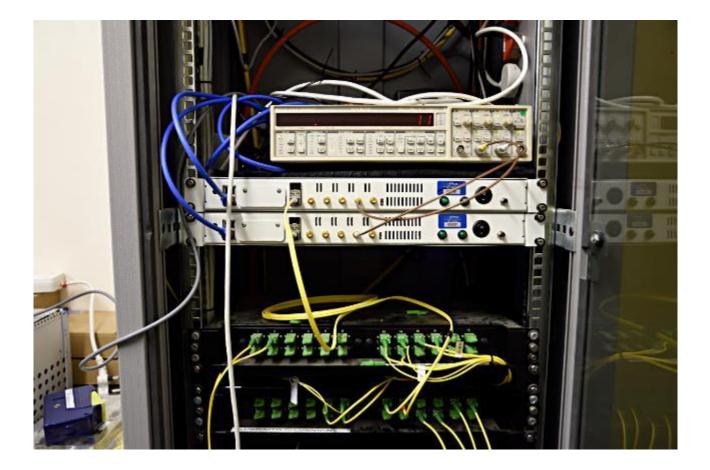


- Absolute time at antennas and RPFs
- Selected the 'White Rabbit' standard for distribution of absolute time
 - o Open Hardware design
 - o Mostly off-the-shelf hardware
 - Single fibre round-trip measurement and compensation
 - Sub-ns accuracy
- Sub-ns accuracy demonstrated to hold even on overhead fibre in South Africa
- Can meet SKA timing requirement even at the most remote dishes.



SADT.SAT.STFR.UTC Time Distribution Prototype equipment for field testing



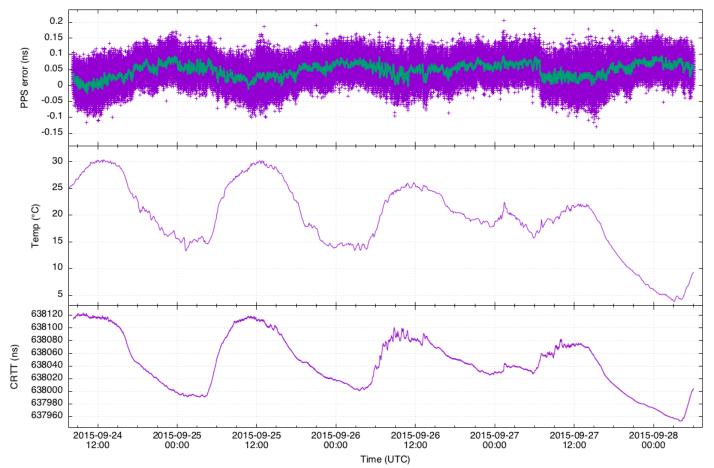


• Time distribution prototypes (White Rabbit equipment) used for overhead cable field testing in South Africa

SADT.SAT.STFR.UTC Time Distribution Time distribution field test results



64km fiber Klerefontein - Carnarvon - Klerefontein, 80km BiDi SFP, 2x WR-Zen



Performance logged over several days. 64km overhead cable

Compensated PPS variations are well below 1ns

Correlation between uncompensated Round Trip Time (CRTT) and outside temperature

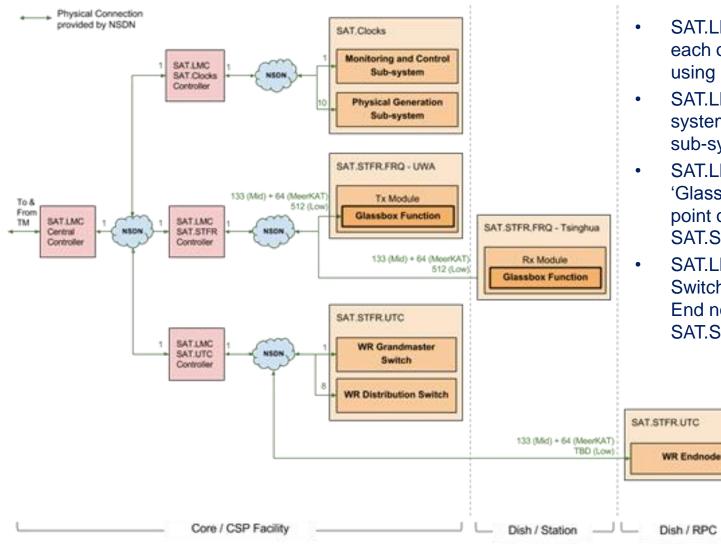


SAT.LMC Synchronisation & Timing : Local Monitoring & Control



 The SAT.LMC will provide monitor and control functionality to the SAT.CLOCKS, SAT.STFR.FRQ and SAT.STFR.UTC sub-systems of the SAT system. This includes the configuration and setup of all the equipment (including hardware and software components), monitoring health and performance parameters, and detecting and responding to adverse situations.

SADT.SAT.LMC SAT Local Monitoring & Control Overview

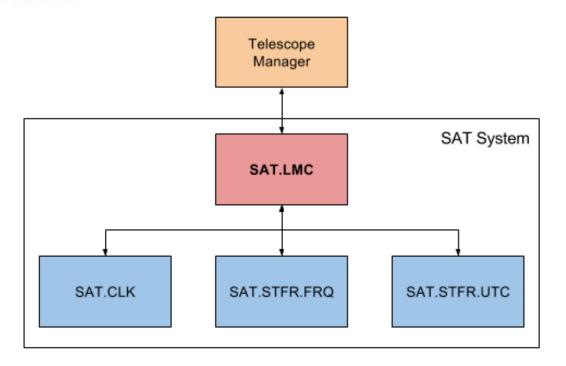


- SAT.LMC components connect to each other and SAT sub-systems using NSDN switches
- SAT.LMC connects to M&C subsystem and Physical Generation sub-system of SAT.CLOCKS
- SAT.LMC connects to the
 'Glassbox 'Function as a single point of contact for
 SAT.STFR.FRQ sub-system.
- SAT.LMC connects with the WR Switches at the core and WR End nodes at the Dish / RPC of SAT.STFR.UTC sub-system

SADT.SAT.LMC North & south bound interfaces



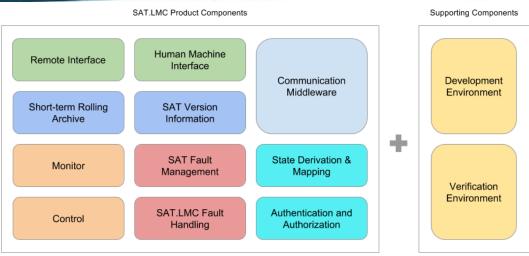




- Monitors and Controls SAT sub-systems (SAT.CLOCKS, SAT.STFR.FRQ and SAT.STFR.UTC)
- Receives command requests and sends monitoring data to TM
- Derives SAT operating state and translates to states understood by TM
- Hierarchical monitoring and control structure
- Coordinates alarm handling across SAT sub-systems
- Uses modular and compact PC104 industrial PCs as SAT sub-system controllers
- Software a major component and not hardware

SADT.SAT.LMC Product component breakdown





- Monitor : Receive and send monitor data
- Control : Receive control commands from TM and send to SAT sub-systems
- Remote Interface : Interface to log into the SAT.LMC controllers from TM
- Human Machine Interface : Interface for the operator to login to the SAT.LMC controller
- Short-term Rolling Archive : Stores logs for the last 24 hours
- SAT Version Information : Stores information on the versions of S/W, H/W and F/W for SAT sub-systems
- SAT Fault Management : Handles faults across SAT sub-systems.
- SAT.LMC Fault Handling : Handles faults generated by SAT.LMC
- State Derivation & Mapping : Derives SAT state from SAT sub-systems and translates to TM state
- Authentication & Authorization : Authorizes and authenticates users for SAT.LMC
- Development Environment : Contains tools to develop SAT.LMC
- Verification Environment : Contains tools, scripts etc. to verify SAT.LMC





Other Information



Re-Baselining & Past/Future events

Re-Baselining Statement (RBS), March 2015 Impacts on SADT



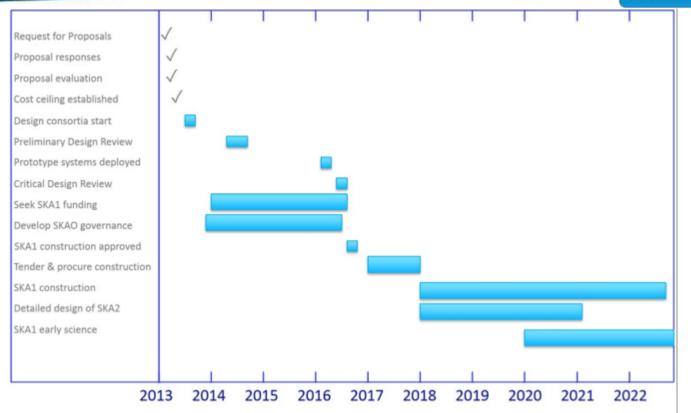
ltem	SKA1 Mid		ltem	SKA1-Low		
	BD V1	BD V2	Item	BD V1	BD V2	
		00 12	Number of			
Number of dishes	190	133		262144	131072	h
Number of disnes	190	133	dipoles Number of	202144	131072	
				1024	F10	•
			stations	1024	512	•
			Antennas per			
			station	256	256	
Receiver bands	1,2,3,4,5	2,5,1				
	0.35-13.8	0.35-13.8				•
Frequency range	GHz	GHz		50-350 MHz	50-350 MHz	
		150 (120)			80 km (70	
Baseline lengths	150km	km	Baseline lengths	80km (70 TBC)	твс)	
			Remote stations			
			built first to			
			demonstrate			
			calibration			
			Remote Station		More uniform	
			distribution	spiral	(random?)	
			Spectral	Spiral	(random:)	
Spectral channels		65536	channels		65536	
Spectral channels		05550	Channels		05550	

Impacts from RBS

- Deferring SKA1_Survey
- Reducing the number of stations and dishes
- Reducing the number of frequency bands
- CSP-SDP required metadata now included in Re-Baselining
- Pulsar beams in SKA1_Low & SKA1_Mid

SKA1 Schedule and Recent Events - original





- Completed Preliminary Design and Costing
 - Elements had external reviews end of 2014
- SKAO made a re-baselining exercise to fit the € 650 M cost cap.
 - SKA Board approval 9th March 2015
- Critical Design Reviews to commence late 2016

