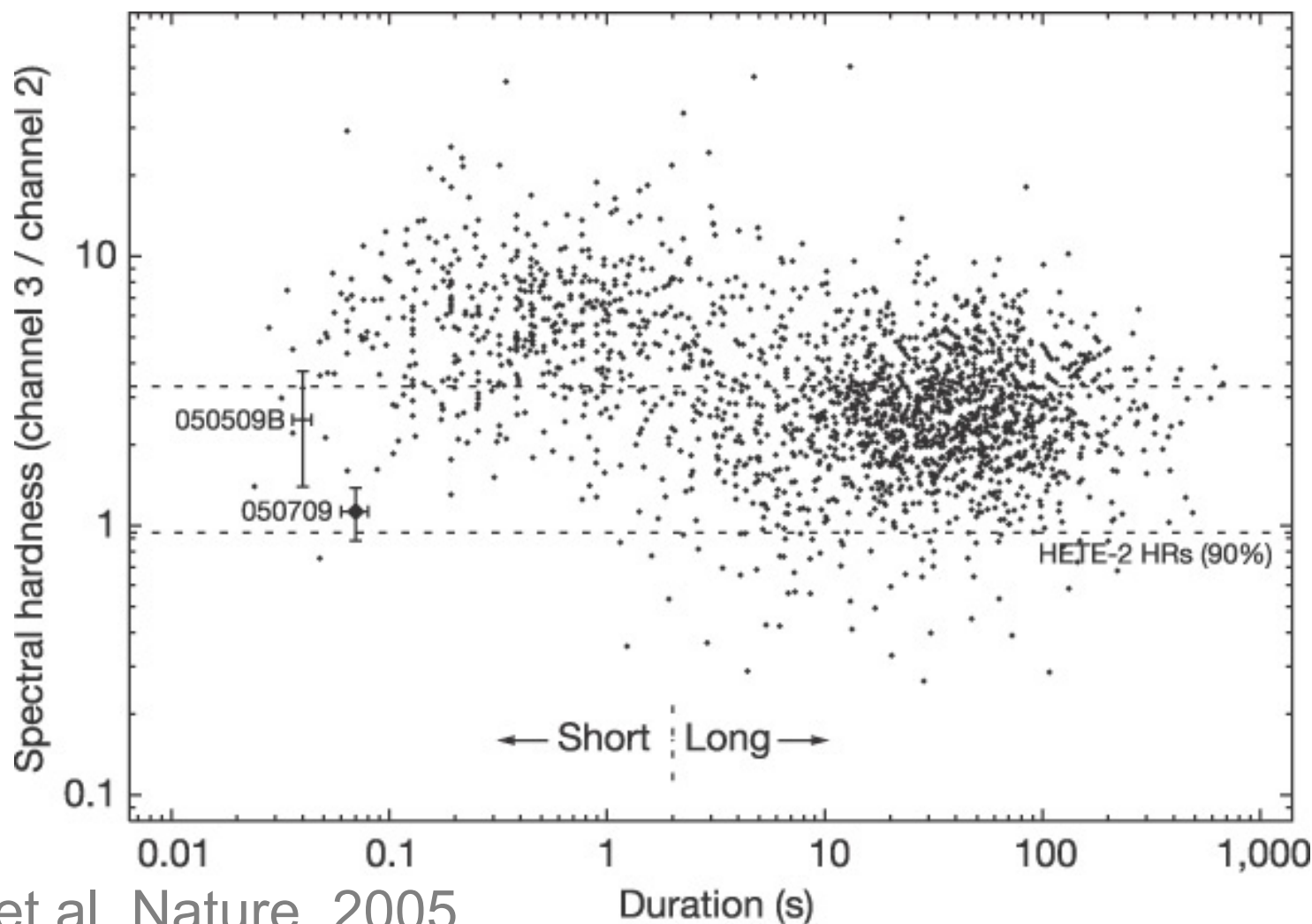


Searches for Gravitational Wave Signals Associated to Gamma Ray Bursts

Stephen Fairhurst

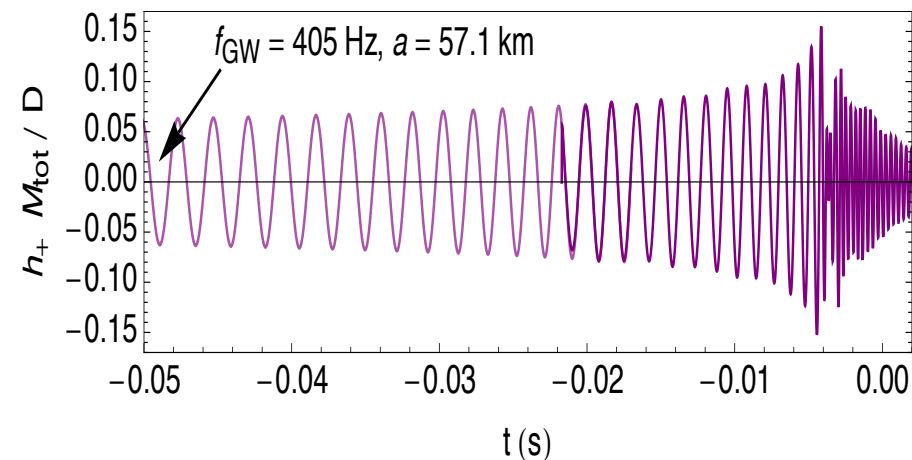
Cardiff University
for the LIGO Scientific Collaboration
and the Virgo Collaboration

Gamma Ray Burst Classification



Gravitational Wave Emission

Short GRB: Merger of
Neutron Star-Neutron Star
or Neutron Star-Black Hole
known GW emission

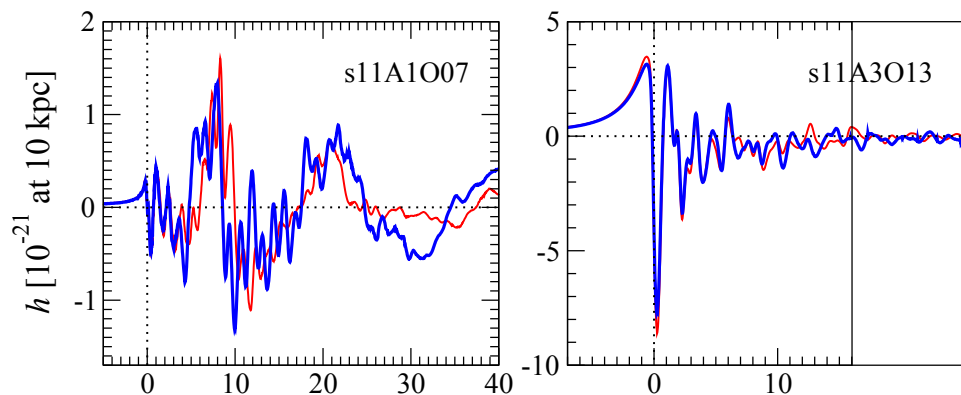


J Read/YITP (In Prep)

Long GRB: Collapsar
uncertain GW emission

- Emission from $10^{-2} M_{\odot} c^2$
Piro and Pfahl 2007,
Davies et al 2002.

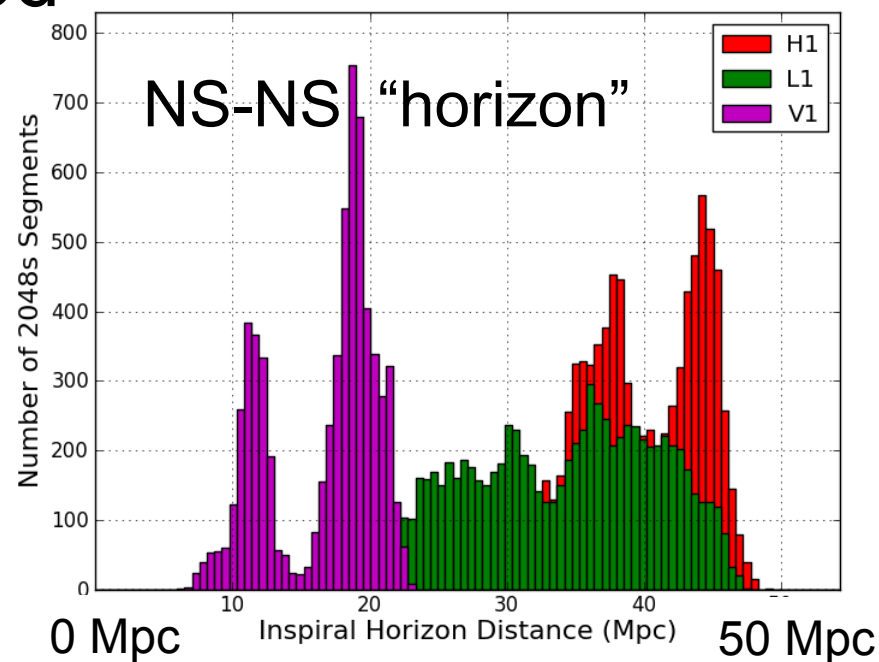
- To $10^{-8} M_{\odot} c^2$ Ott CGQ (2009)



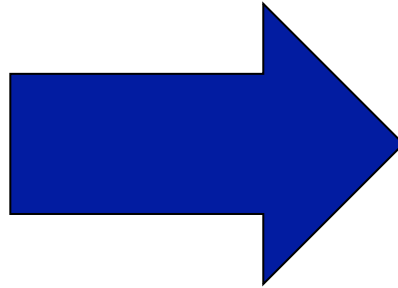
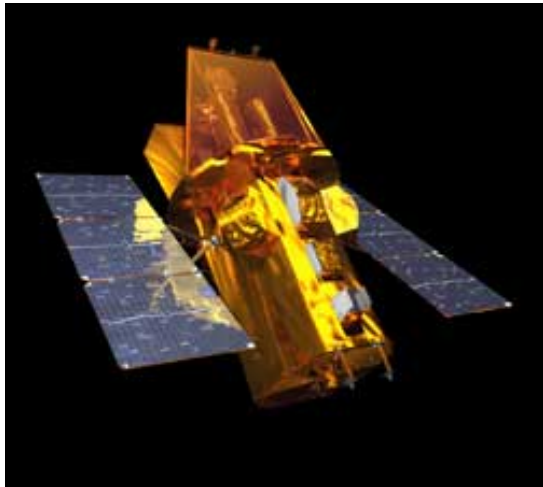
- Or less

Gravitational Wave Detectors

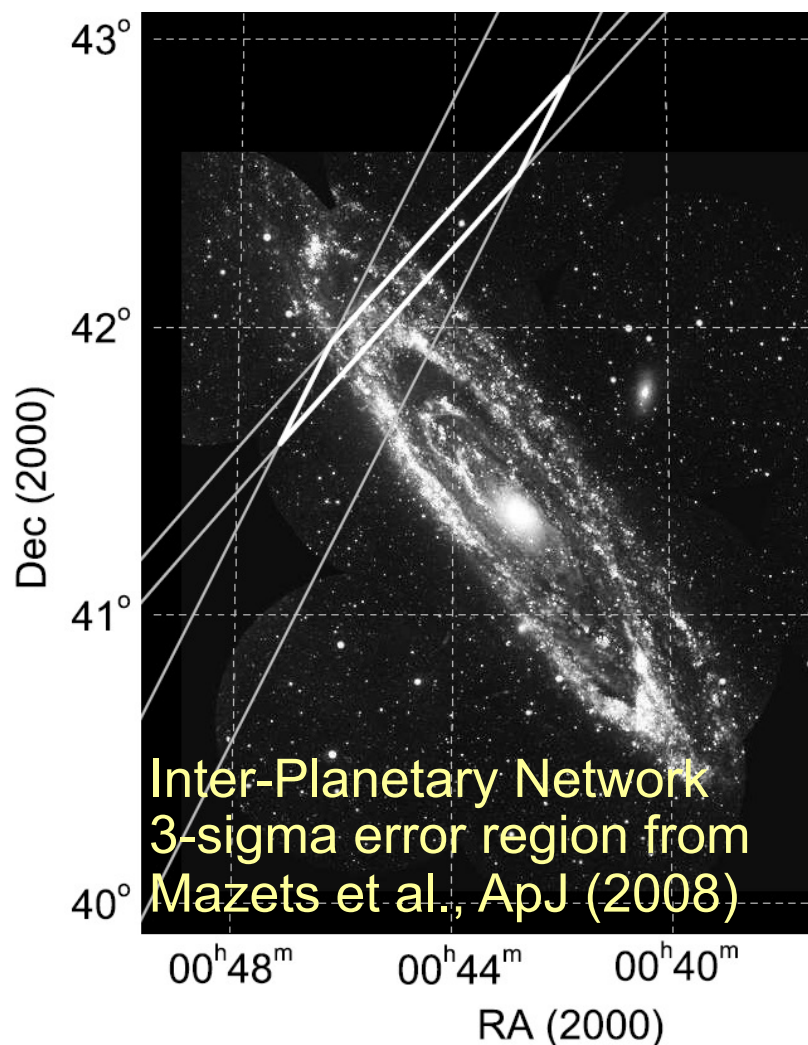
- Operated from Nov 2005 to Sept 2007 & July 2009 to Oct 2010
- Currently being upgraded
- From 2015, advanced detectors
 - 10x increase in (distance) sensitivity



Triggered Search



GRB 070201

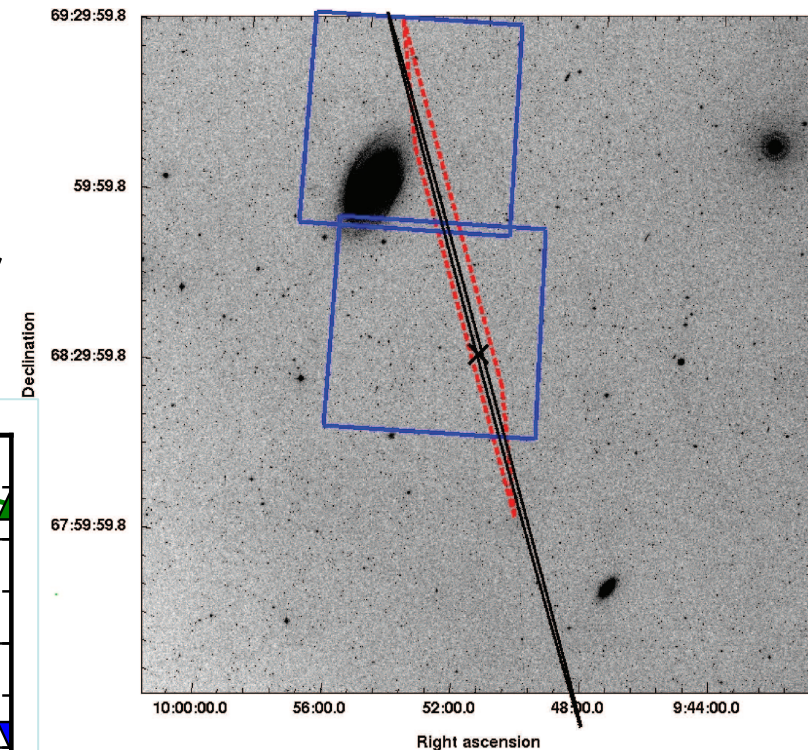
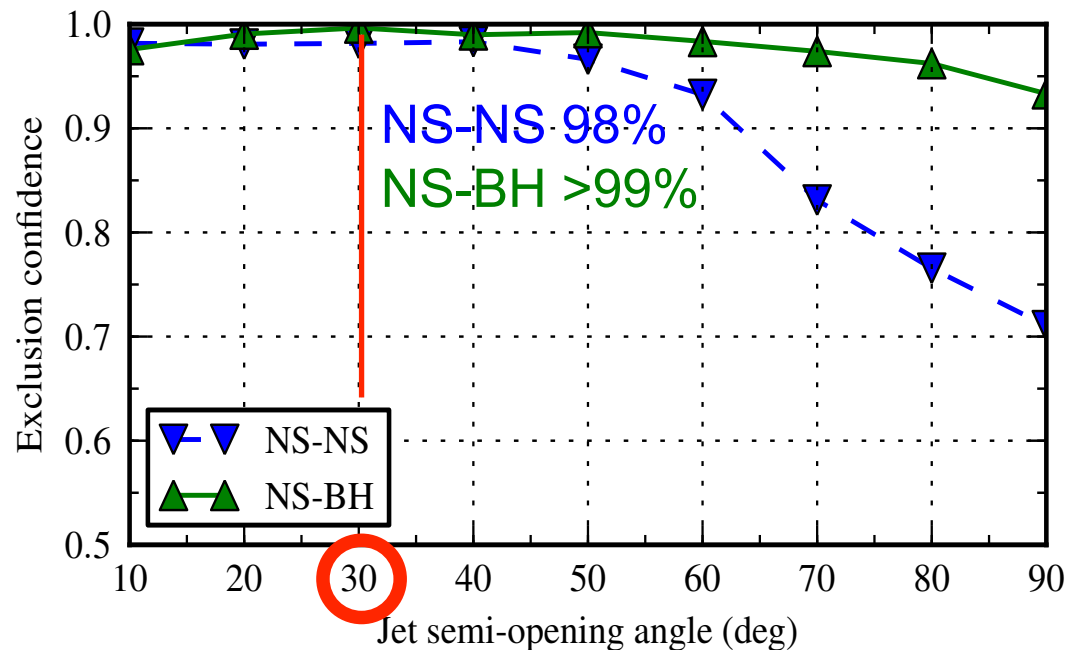


- Localization overlaps M31 (at 770 kpc)
- No GW signal observed
- Exclude NS-NS and NS-BH merger in M31 with 99% confidence
- Indirect support for hypothesis of soft gamma repeater in M31

Abbott et al. ApJ 2008

GRB 051103

- Localization overlaps M81 (at 3.6 Mpc)
- No GW signal observed
- Exclude binary merger progenitor as function of opening angle



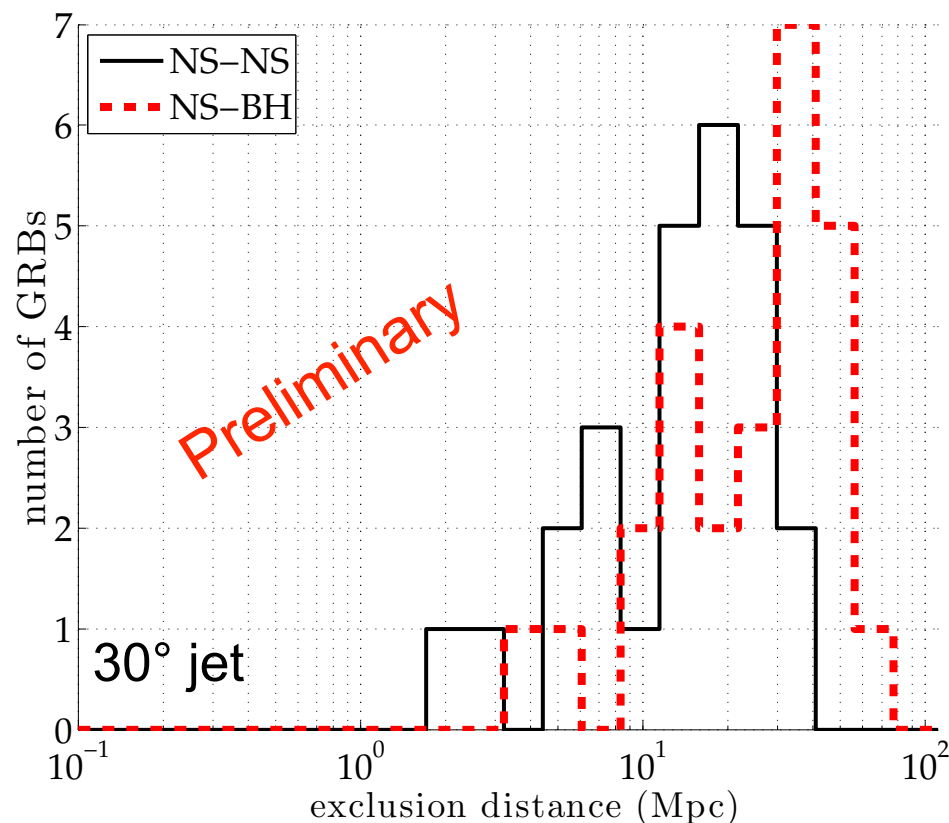
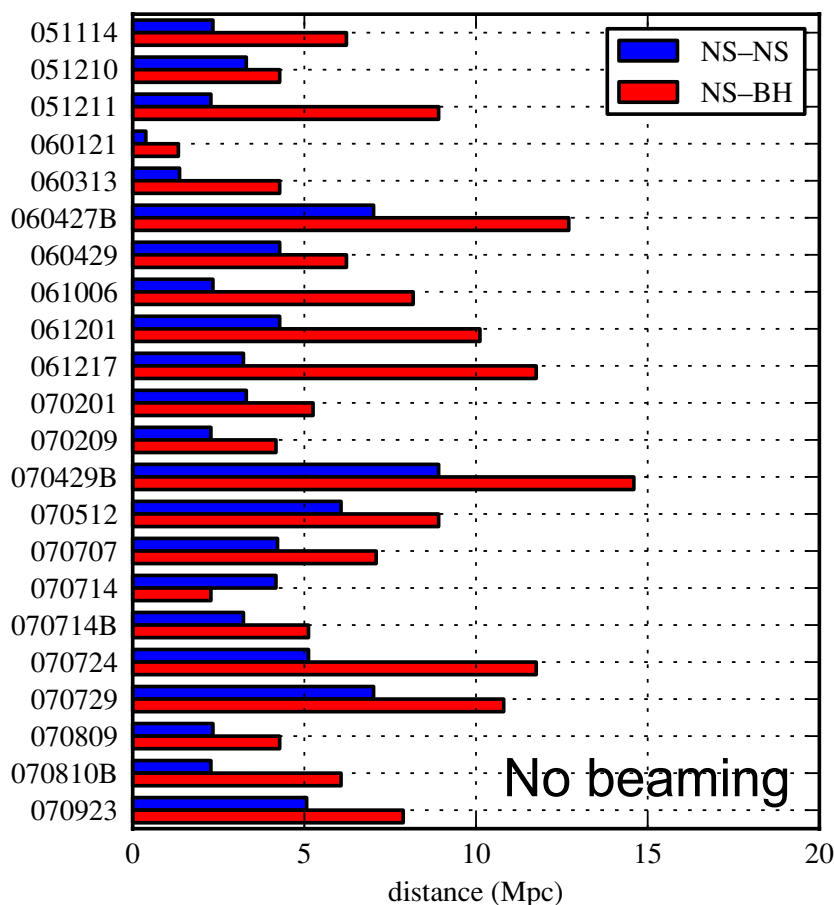
Abadie et al. 1201.1163

Short GRB exclusion distances

2005-7 run

Binary Merger Model

2009-10 run



Exclusion distances for all GRBs

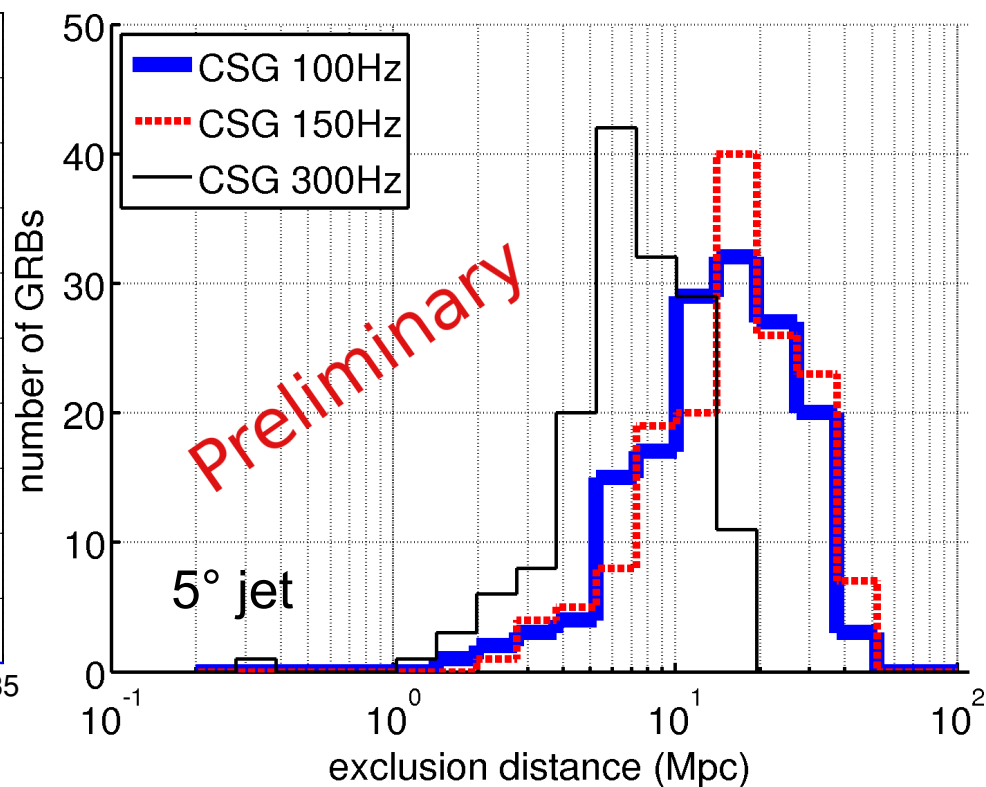
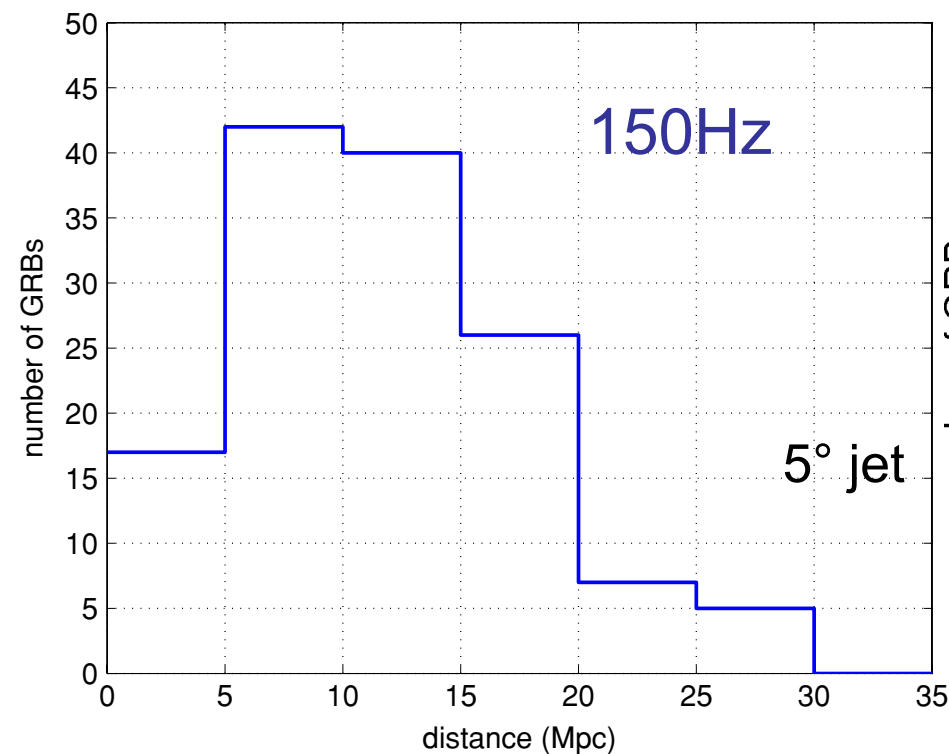
Unmodelled GW Burst

Assuming $10^{-2} M_{\odot} c^2$ in GW;

2005-7 run

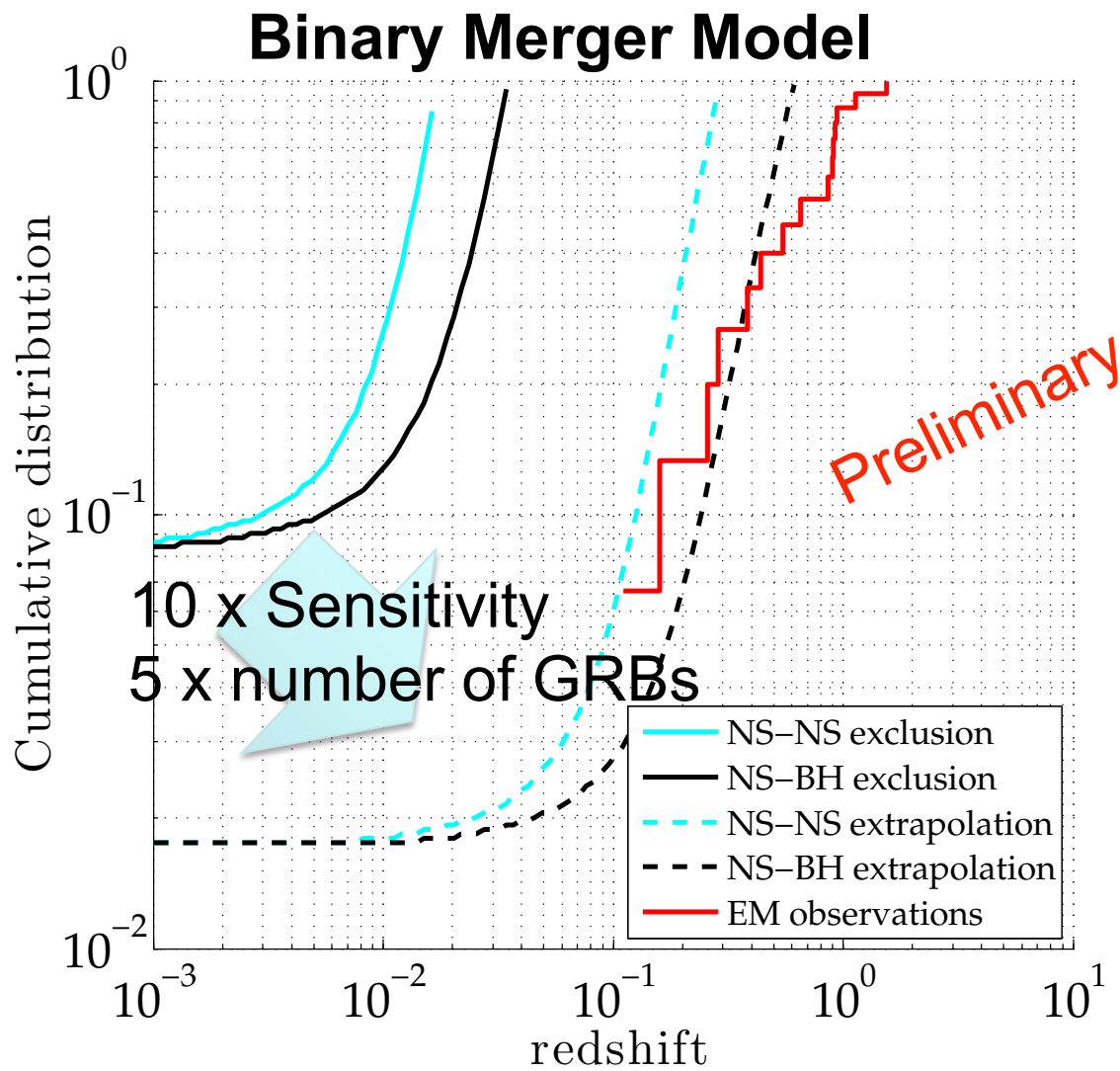
emitted in small frequency band

2009-10 run

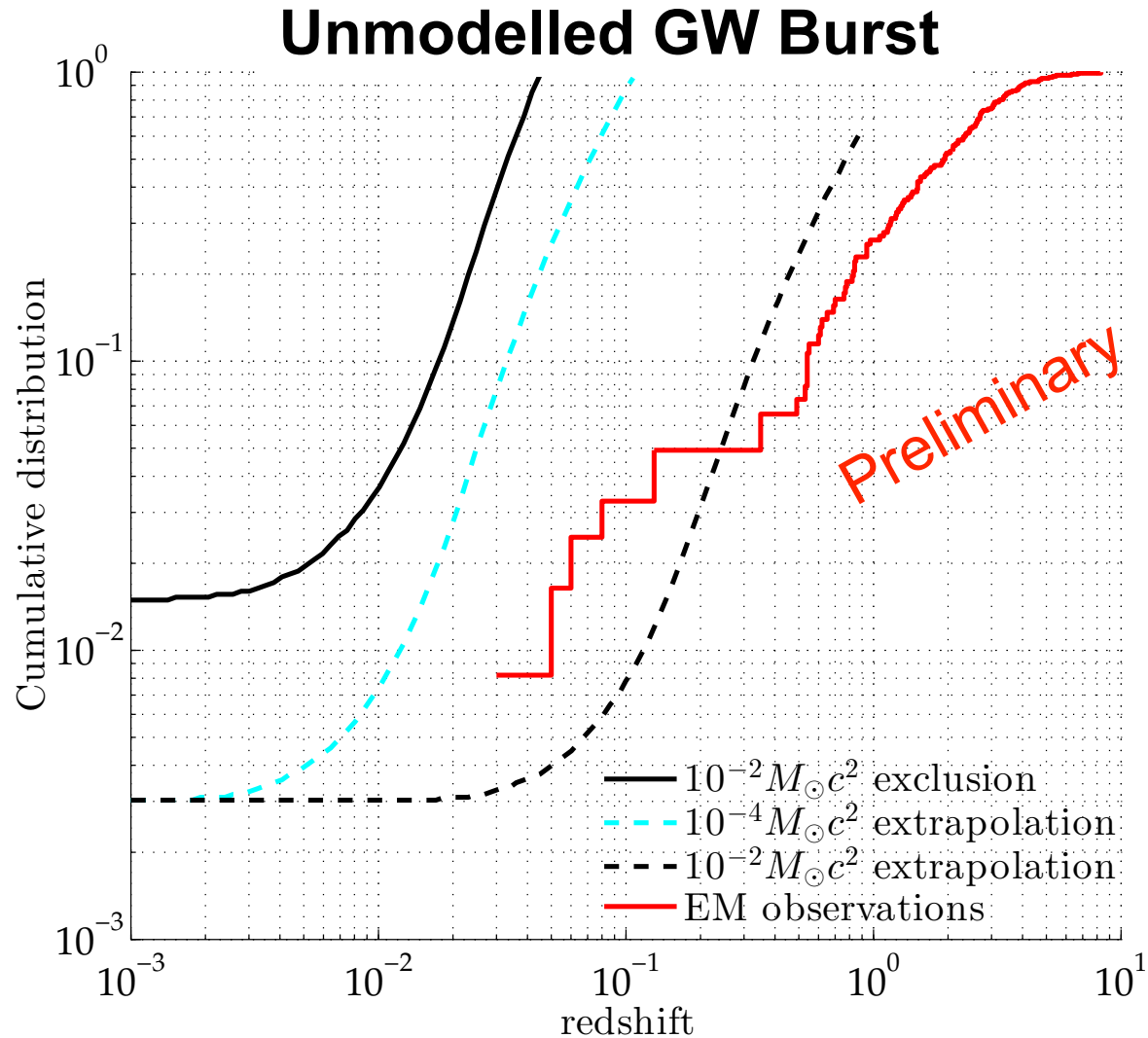


Abbott et al ApJ (2010)

Future Prospects

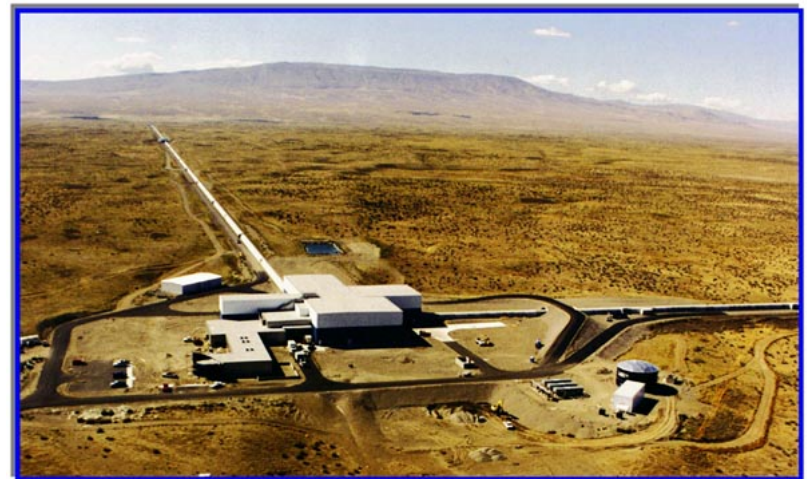


Future Prospects



Future Prospects

- Advanced detectors may detect GW associated to GRBs
 - Confirm (or rule out) progenitor models



- Prospects are strongly dependent on number of GRBs observed electromagnetically.