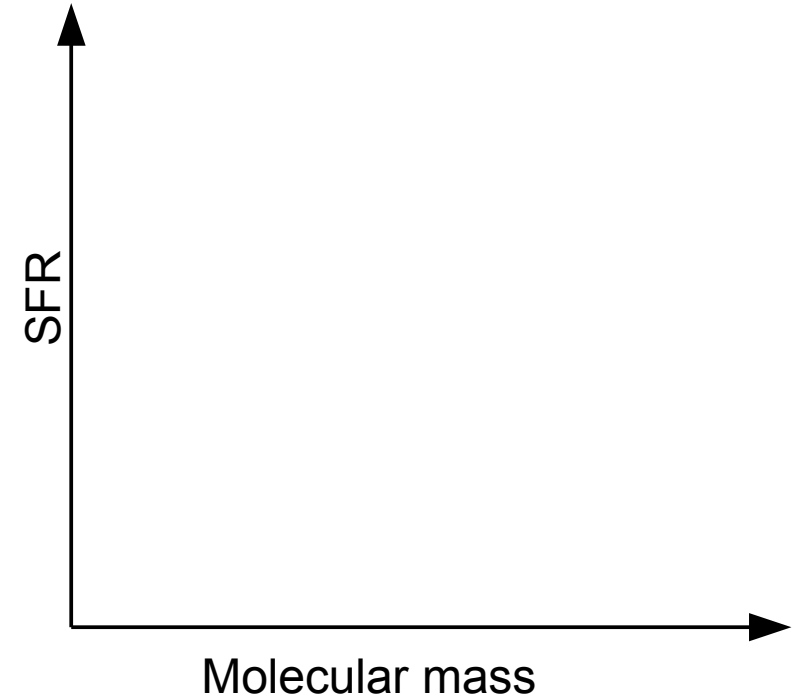
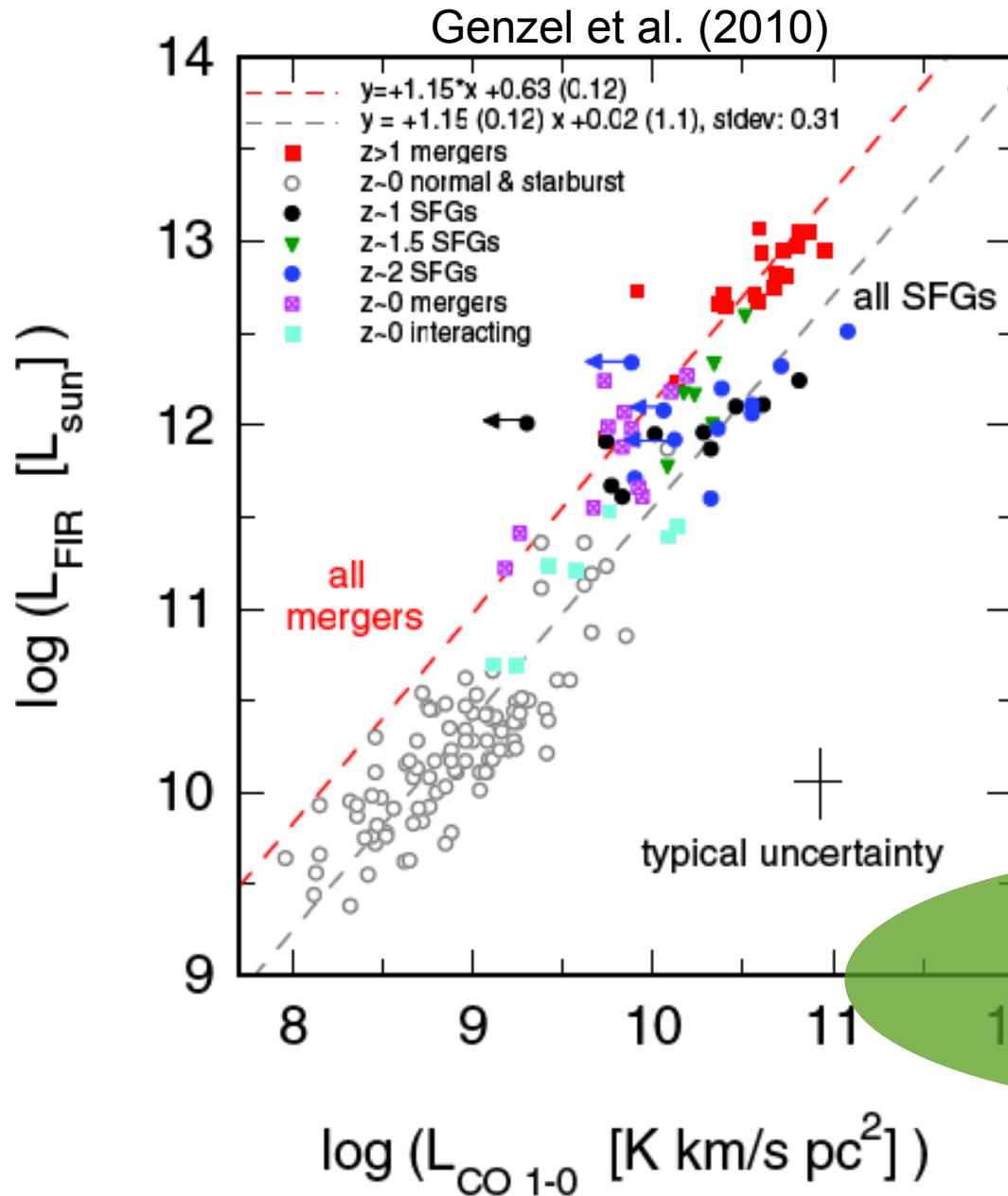


**A new technique to
understand the relation
between star formation and
molecular gas**

Claudia Lagos, Carlton Baugh, Cedric Lacey, Estelle Bayet
(Oxford), James Geach (McGill), Tom Bell (Caltech), Nikos Fanidakis
(MPIA)

Observationally we study IR vs. CO



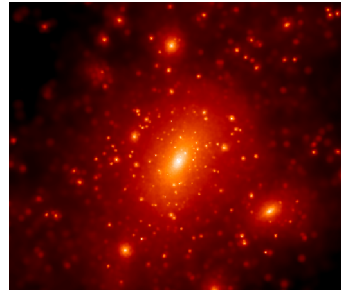
Need to assume a
CO(1-0) to H_2
and a
CO($J \rightarrow J-1$) to CO(1-0)

The paradigm of galaxy formation

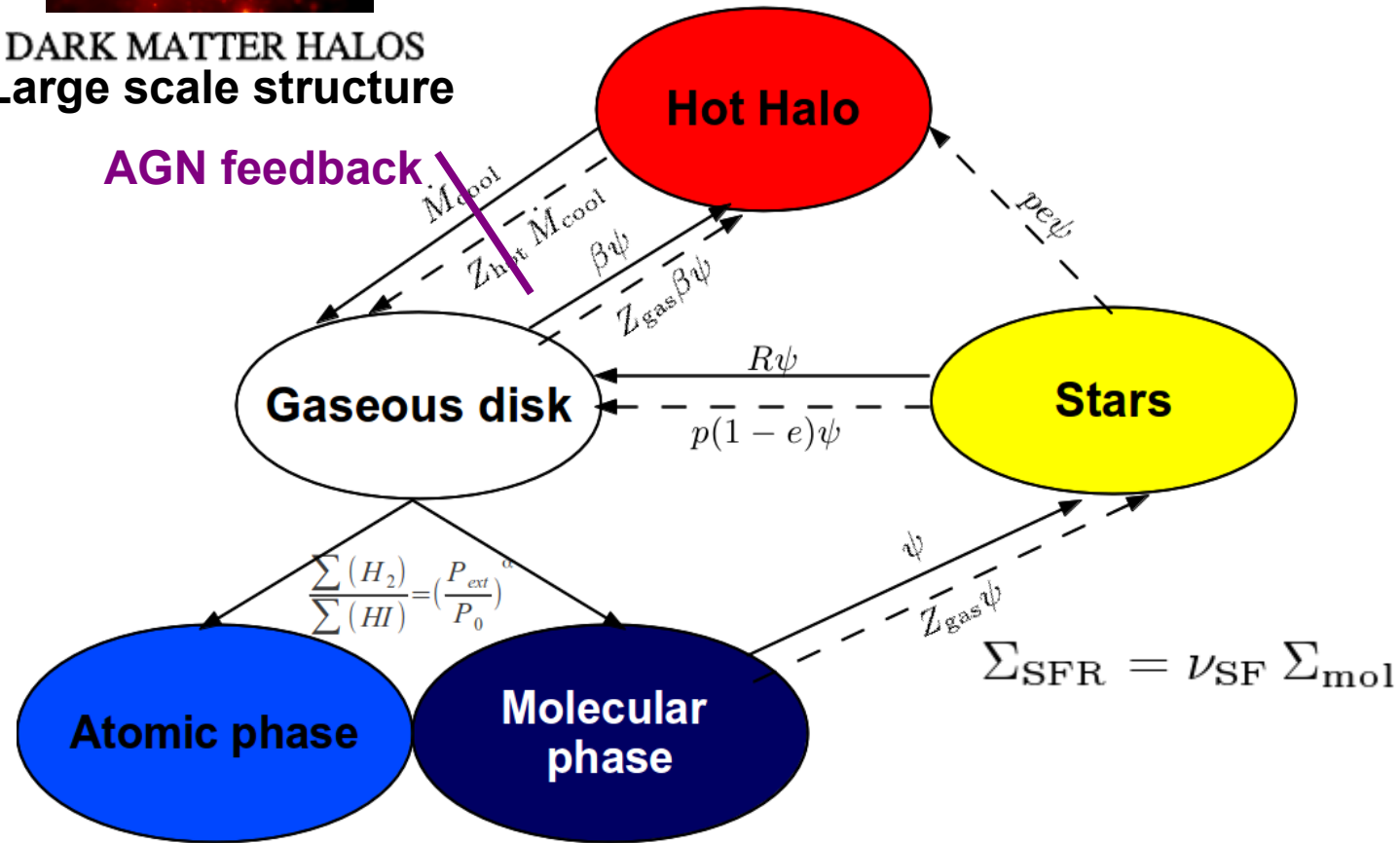
(Lagos et al. 2011a,b)

Cosmological model

$$\Omega_{\nu}, \Lambda_{\nu}, \sigma_8, h, P(k)$$



DARK MATTER HALOS
Large scale structure



AGN feedback

————— Mass exchange
- - - - - Chemical exchange



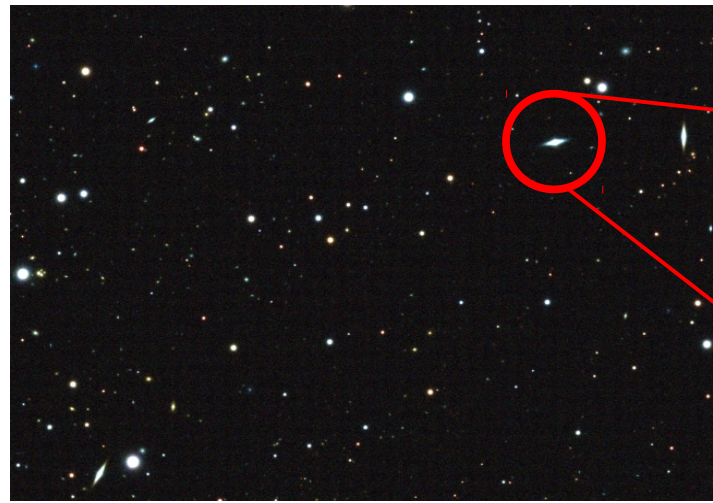
@Daniel Farrow, Pan-STARRS

Predicting the CO emission of galaxies: combining GALFORM with the UCL_PDR RT

(Lagos et al, 2012, arXiv:1204.0795 & Bayet et al., 2011)

→ A novel approach to estimate the CO emission of a large sample (millions) of simulated galaxies.

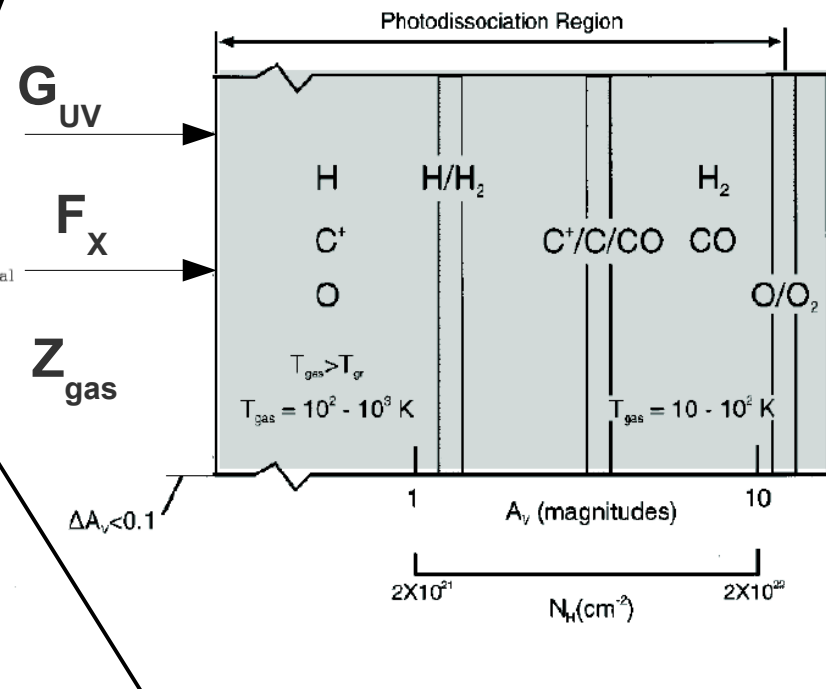
GALFORM



Sizes, luminosities, molecular, atomic gas masses, UV flux, gas metallicities, AGN luminosity

Allows 2 SF gas phases

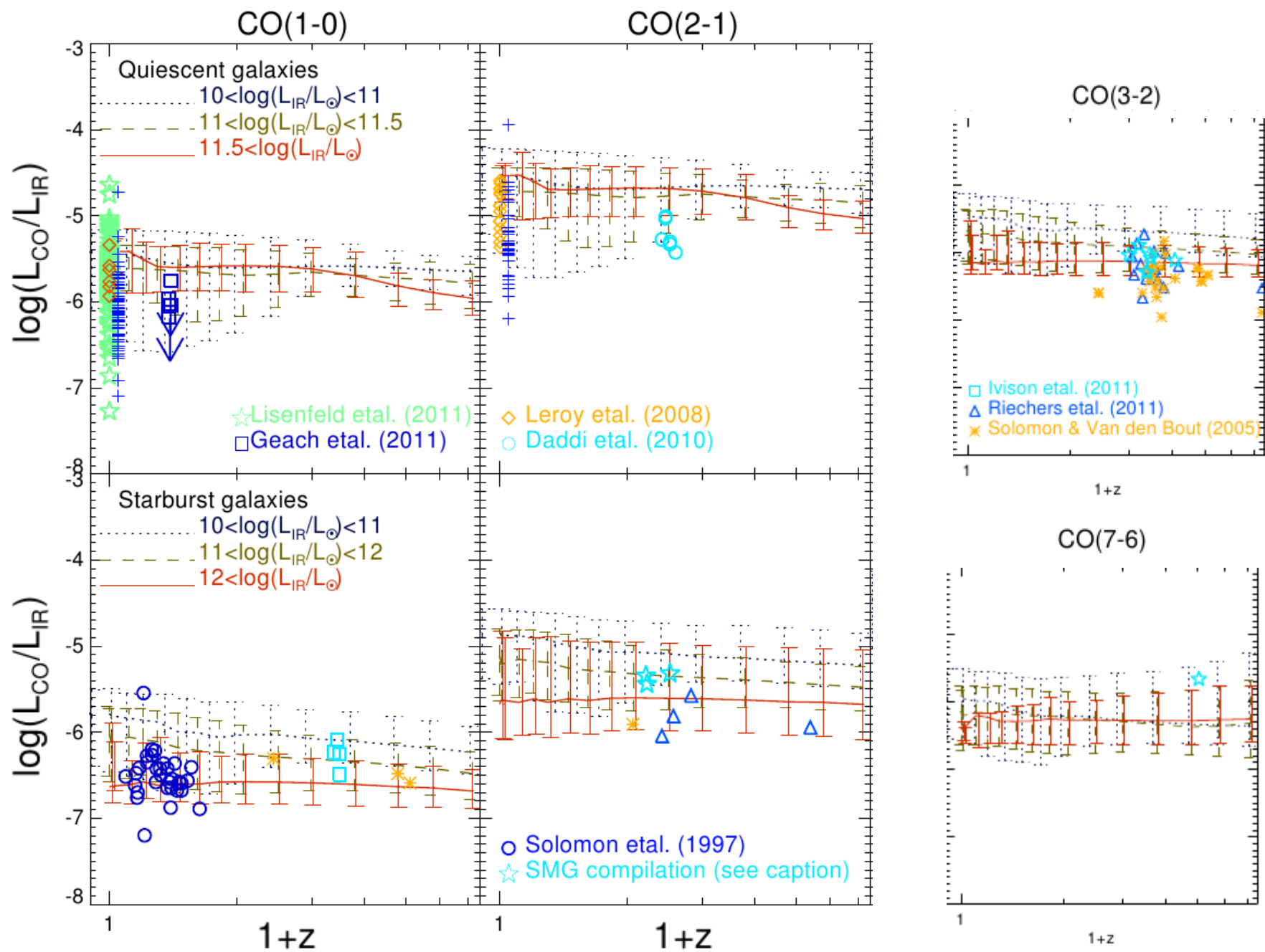
UCL_PDR



X_{CO} T_{gas}

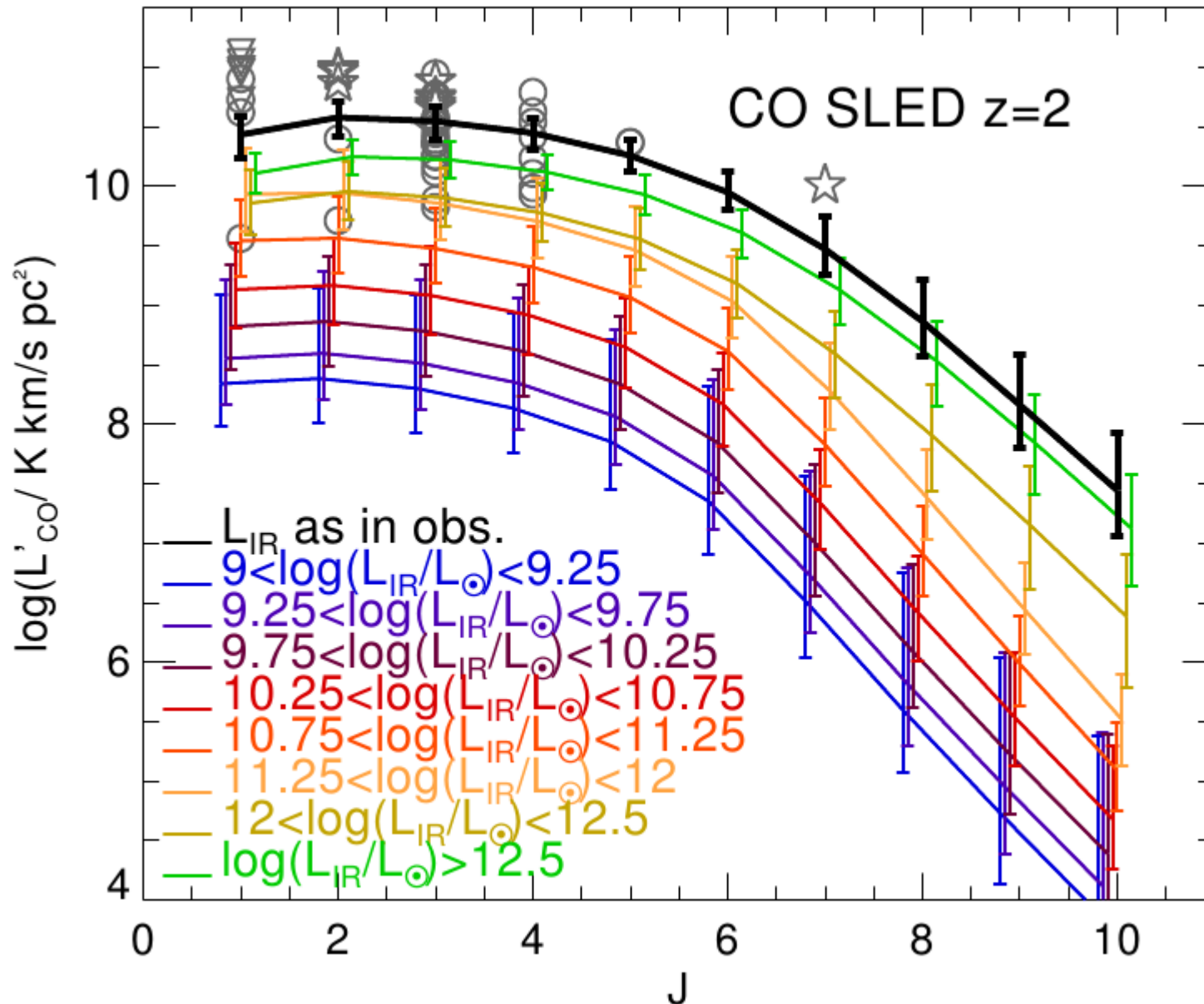
The CO-to-IR luminosity ratio evolution

(Lagos et al, 2012, arXiv:1204.0795)



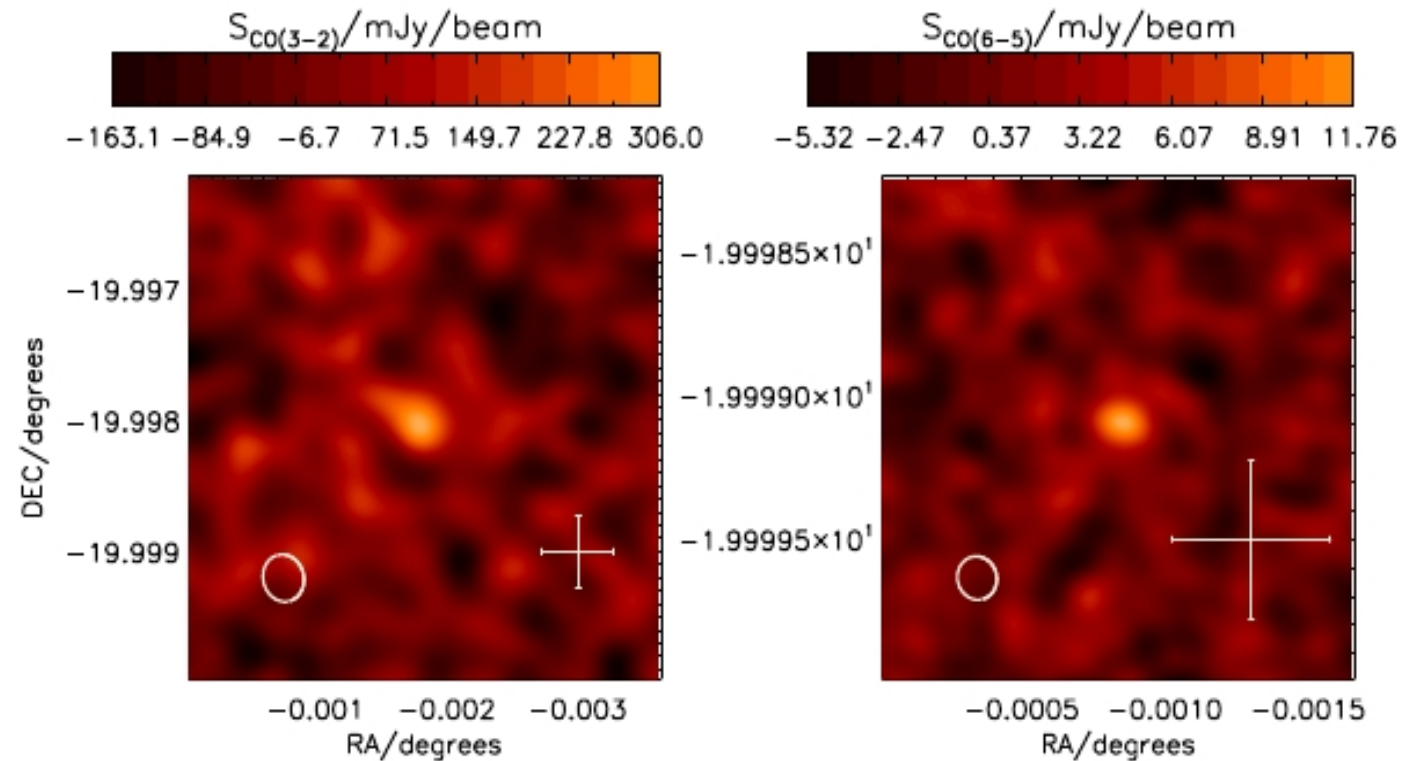
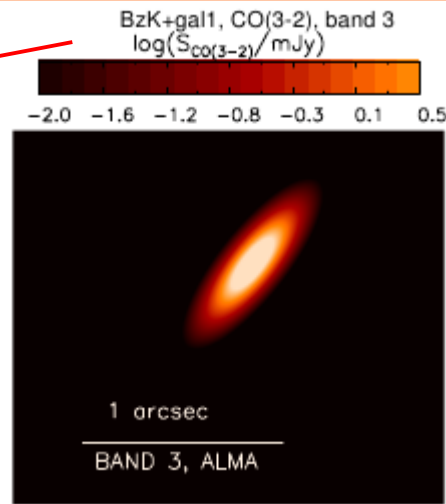
The CO SLED of high-z SMGs

(Lagos et al, 2012, arXiv:1204.0795)



Simulating observations for ALMA

(Lagos et al, 2012, arXiv:1204.0795)



Conclusions

Lagos et al. (2011a), Lagos et al. (2011b), Lagos et al. (2012, arXiv:1204.0795), Bayet et al. (2011, 2012, in prep.)

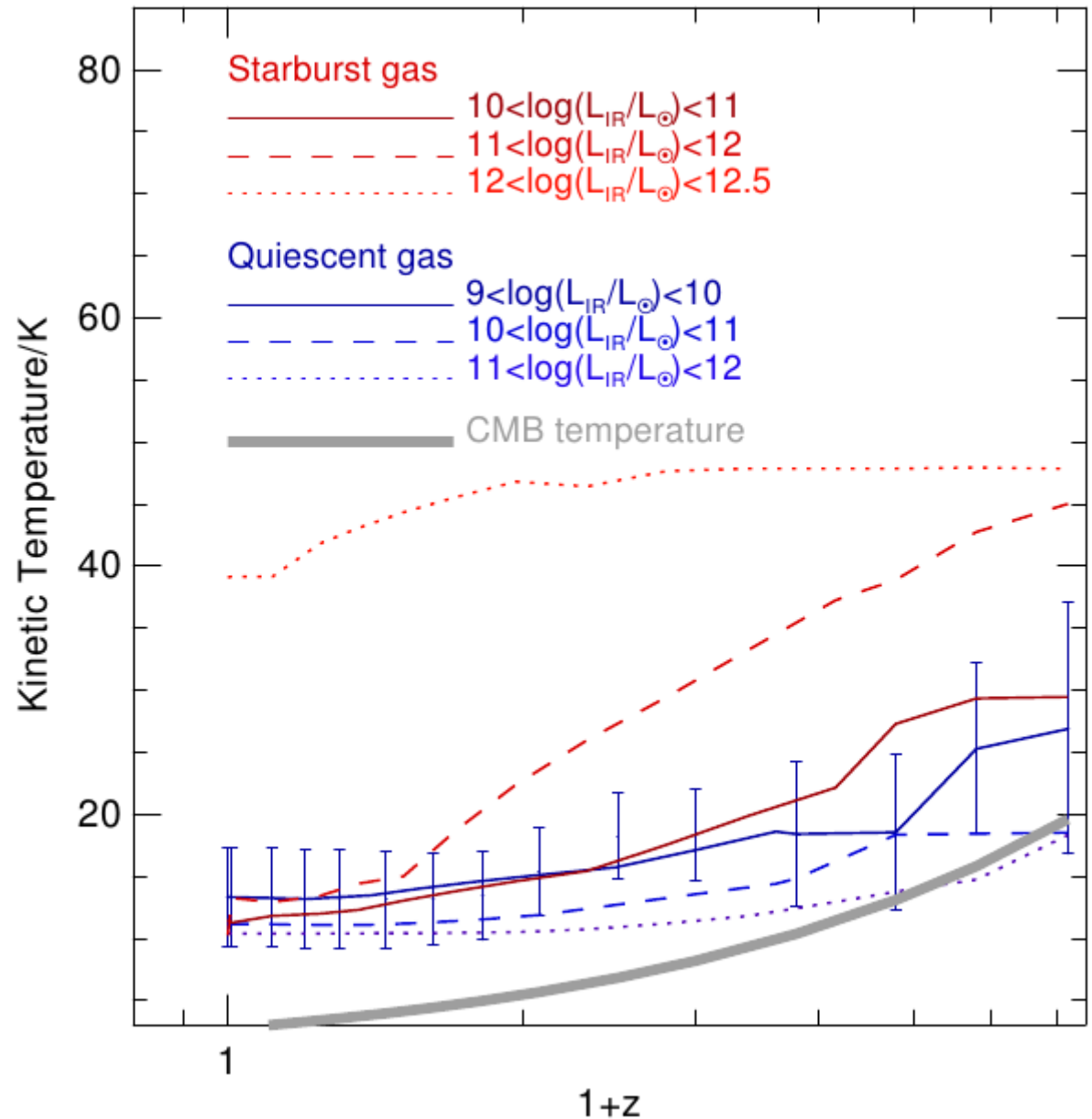
- SAM: Powerful tool to study the connection SF/H₂/HI. Self-consistent use of parameter free SF law.
- **GALFORM+UCL_PDR**: predictions in good agreement with obs.
 - CO(1-0) luminosity function
 - CO-IR luminosity relation of local and high-redshift galaxies
 - CO SLEDs of LIRGs and SMGs.
- **Predictions for ALMA: simulate CO observations**
 - CO catalogues of galaxies to simulate obs.
 - *Diagnose the observability* of e.g. BzKs, LBGs, etc.

Gas kinetic temperature evolution

(Lagos et al, 2012, arXiv:1204.0795)

→ Starbursts tend to have temperatures higher than normal star-forming galaxies

→ Temperature tends to increase with redshift (effect of lower metallicities and higher SFRs)



Evolution of molecular gas fractions (Geach et al. 2011)

- Strong molecular fraction evolution explained by higher ISM pressure (Lagos et al. 2011b)

