

Galaxy Clustering with Pan-STARRS1

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the PS1 team*



Outline

- What is Pan-STARRS1?
- Status of Pan-STARRS1.
- Sample selection, masks, star galaxy separation.
 - Early galaxy clustering results

Pan-STARRS1

- 1.8m on Haleakala, Maui. Built & run by the University of Hawaii.
- 5 optical/NIR bands (roughly SDSS griz + y)
- Large area survey - '3 pi' – ideal for ISW
- Smaller deeper fields - 'Medium Deeps'.
- PS1 Surveys completed October 2013



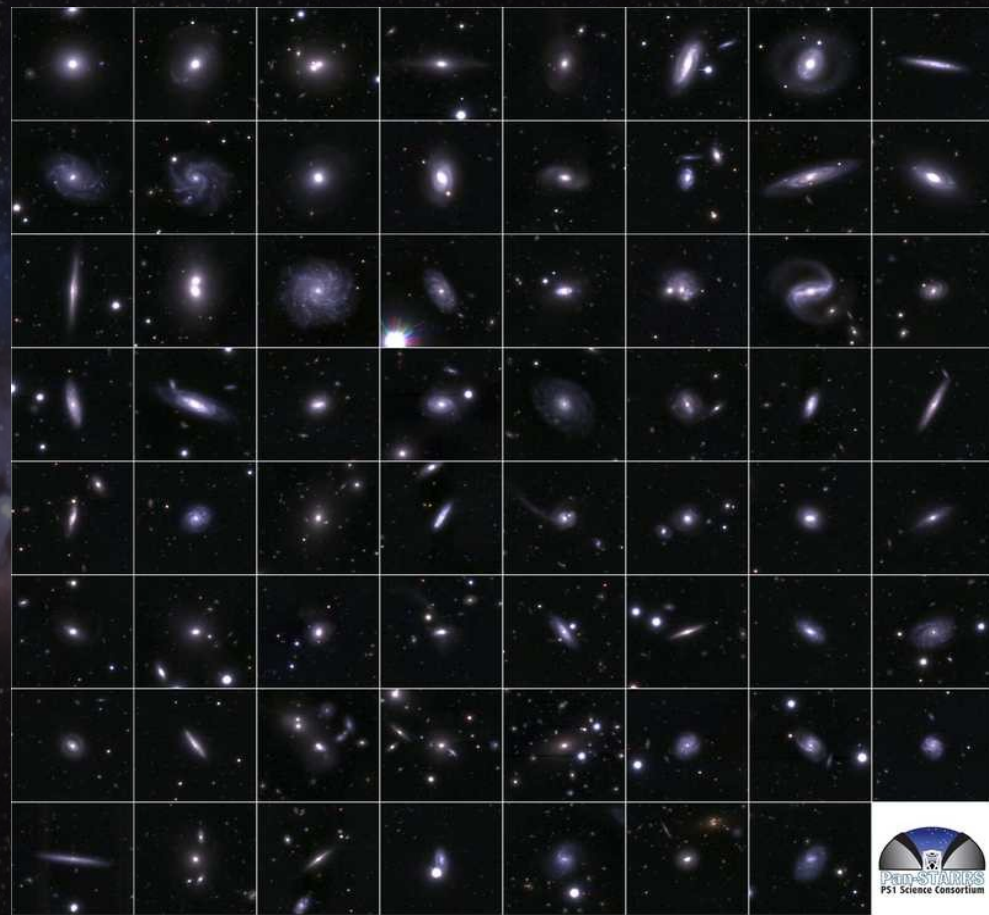
Photo by Rob Ratkowski © 2010 PS1SC.



Pan-STARRS1



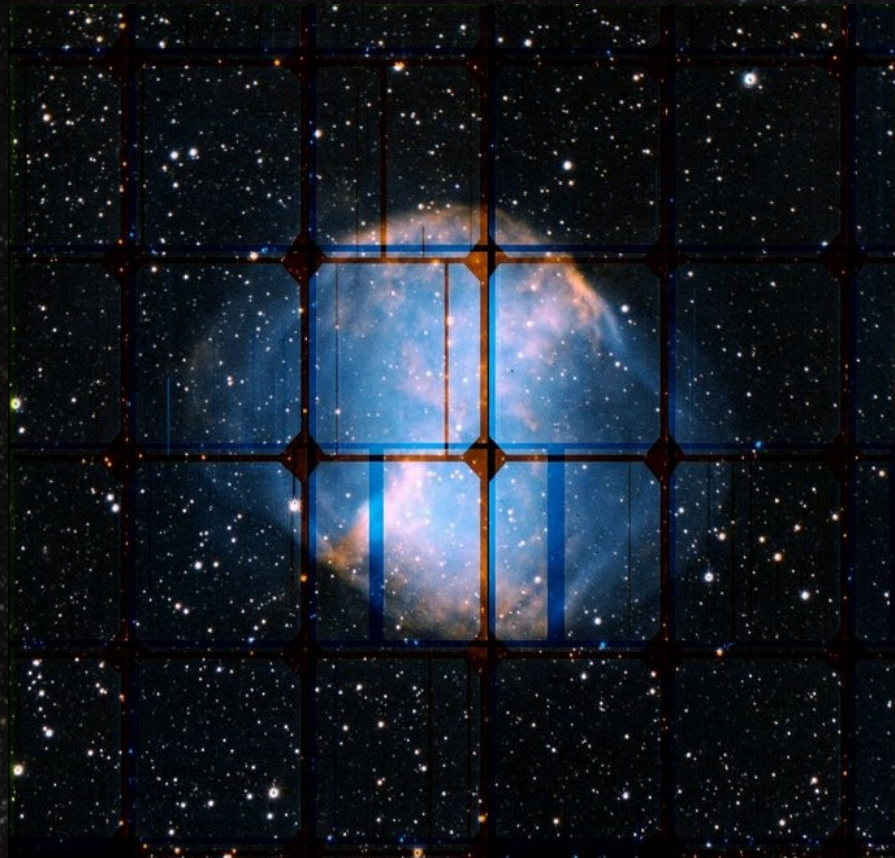
(Background is a zoom of this image)



Credits: Eugene Magnier (UH IfA), Peter Draper & Nigel Metcalfe (Durham University), ©PS1 Consortium

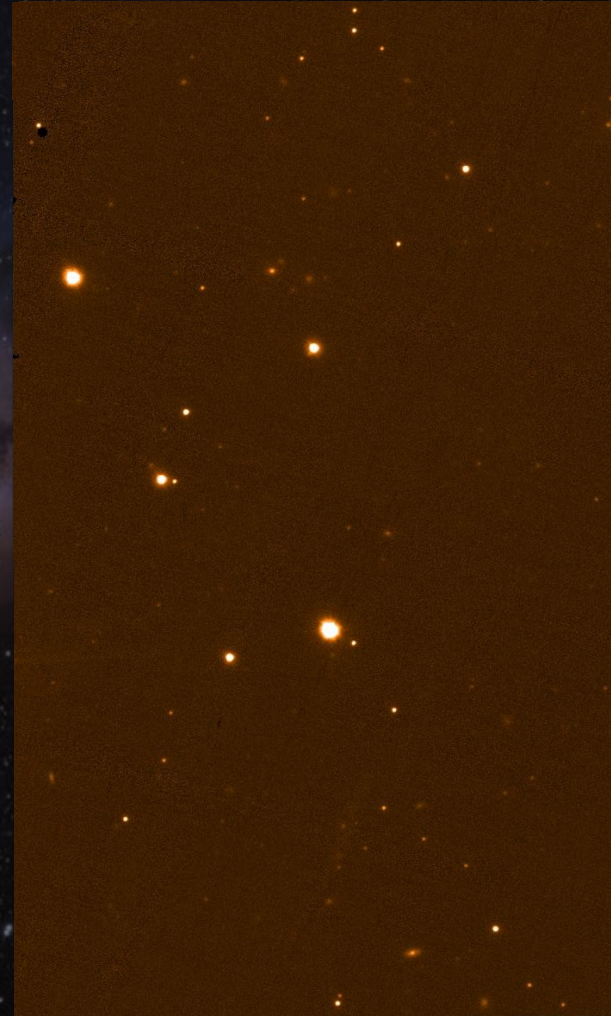
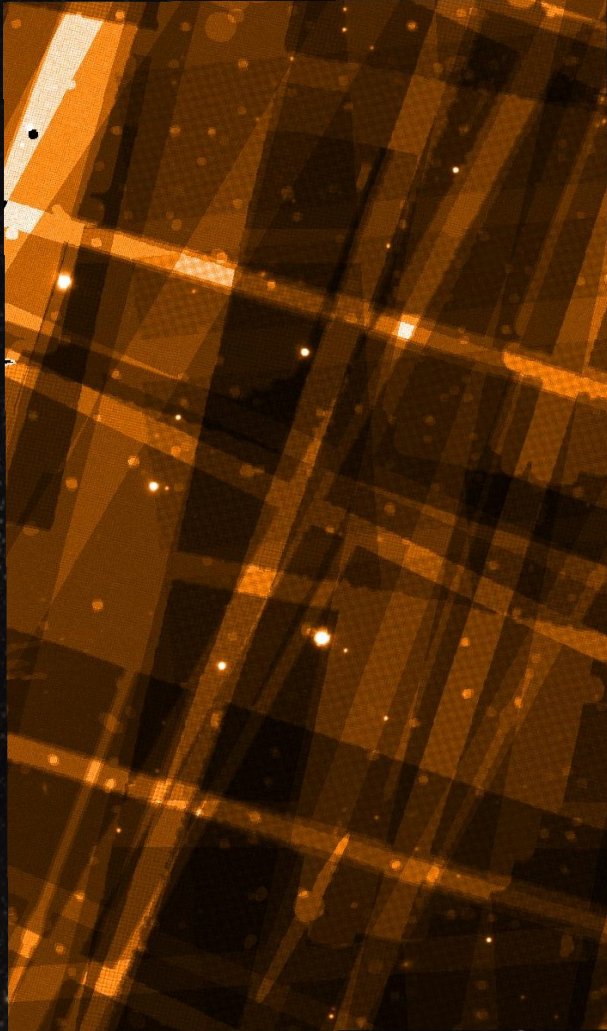
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Observing Strategy



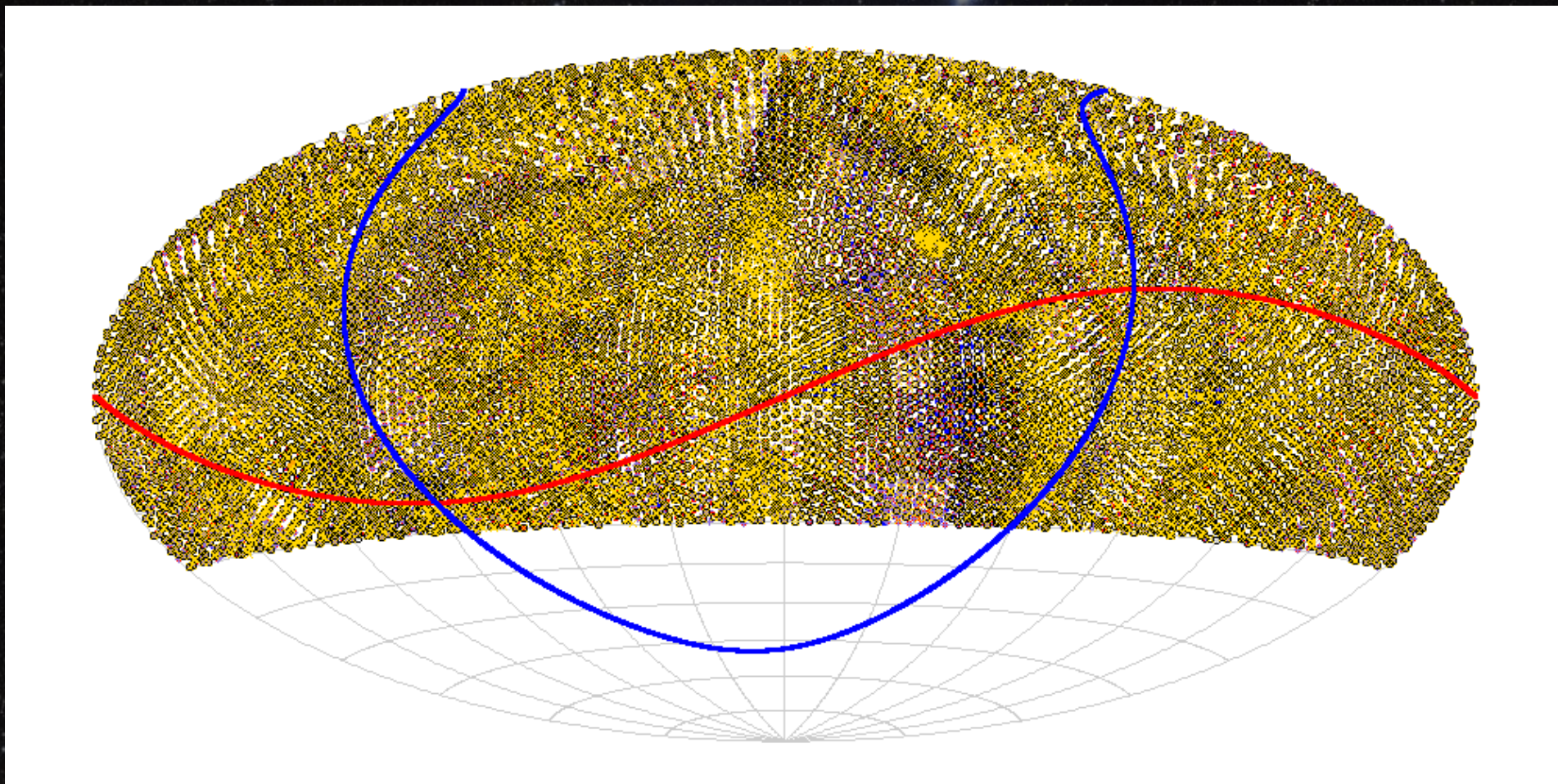
Stack images (in the same band or different bands) with different rotations & different centres to form the final image

Variance Map



Pan-STARRS1 Status

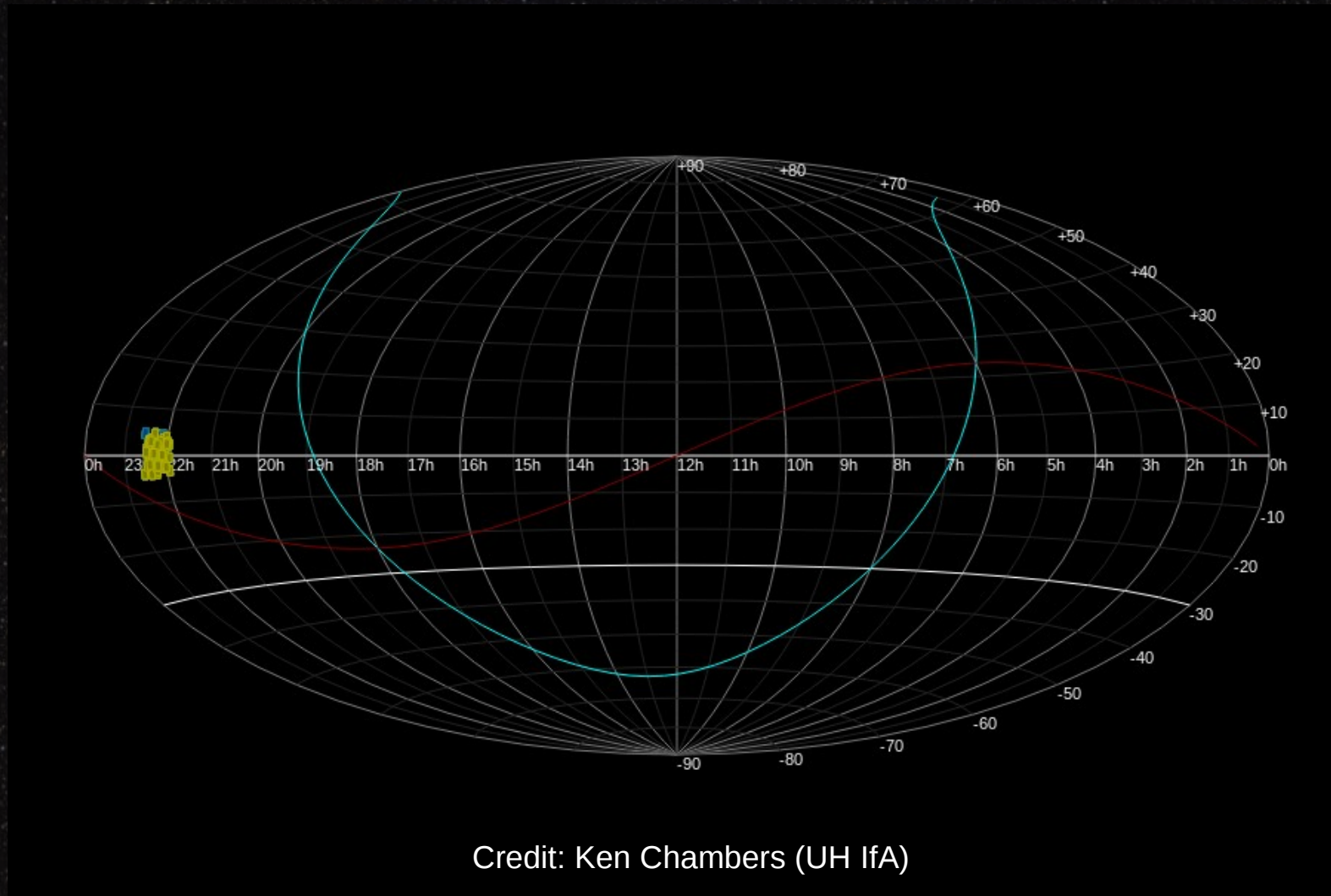
Large area – good for ISW in photo-z slices:



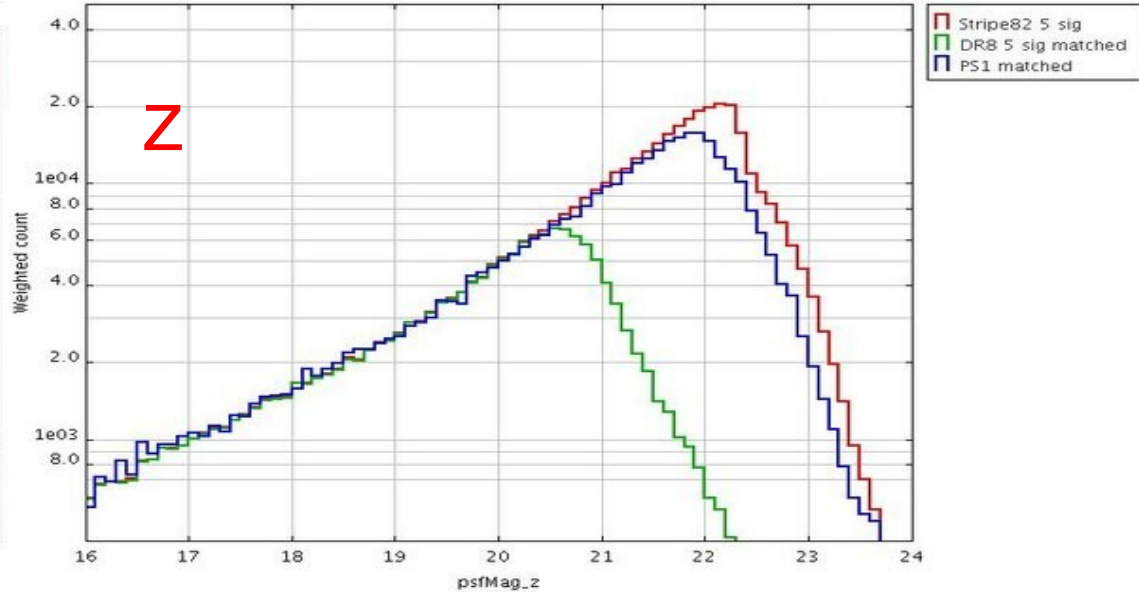
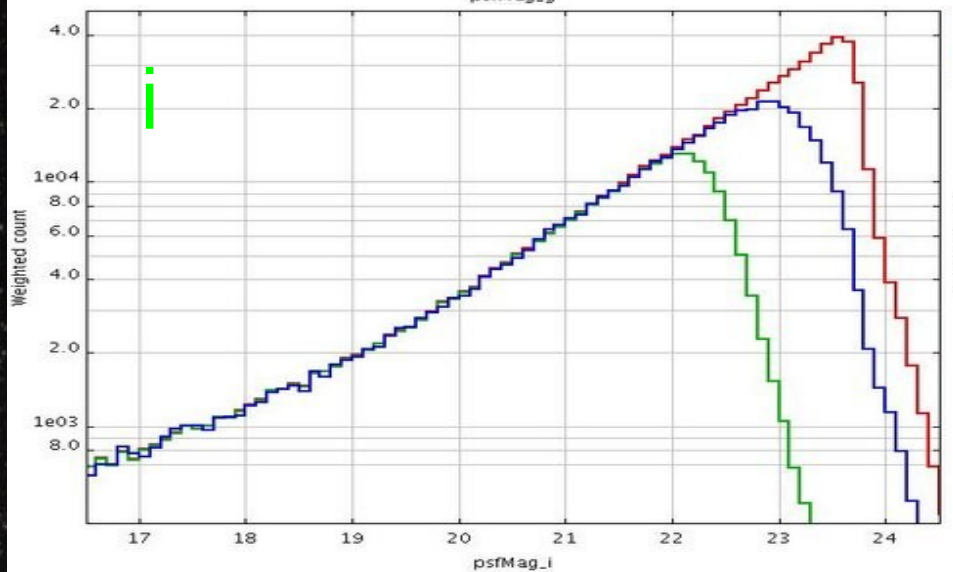
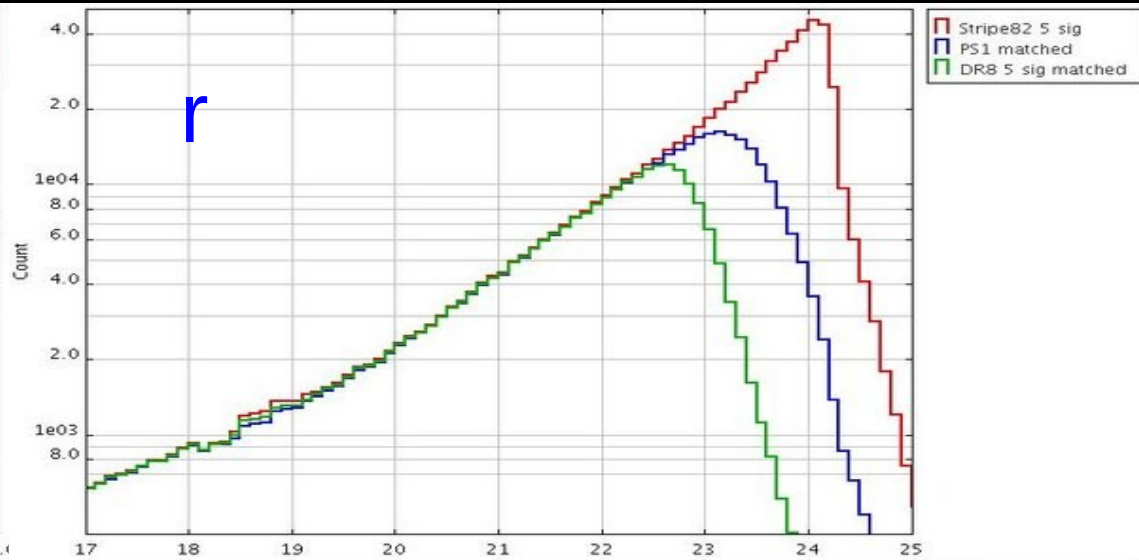
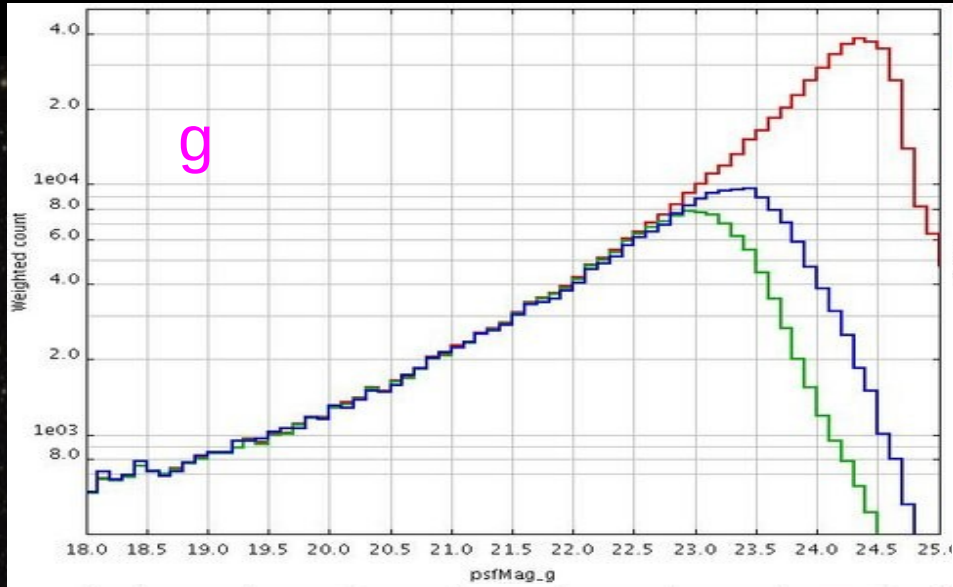
Credit: Ken Chambers (UH IfA)

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Pan-STARRS1 – Small Area Survey 2



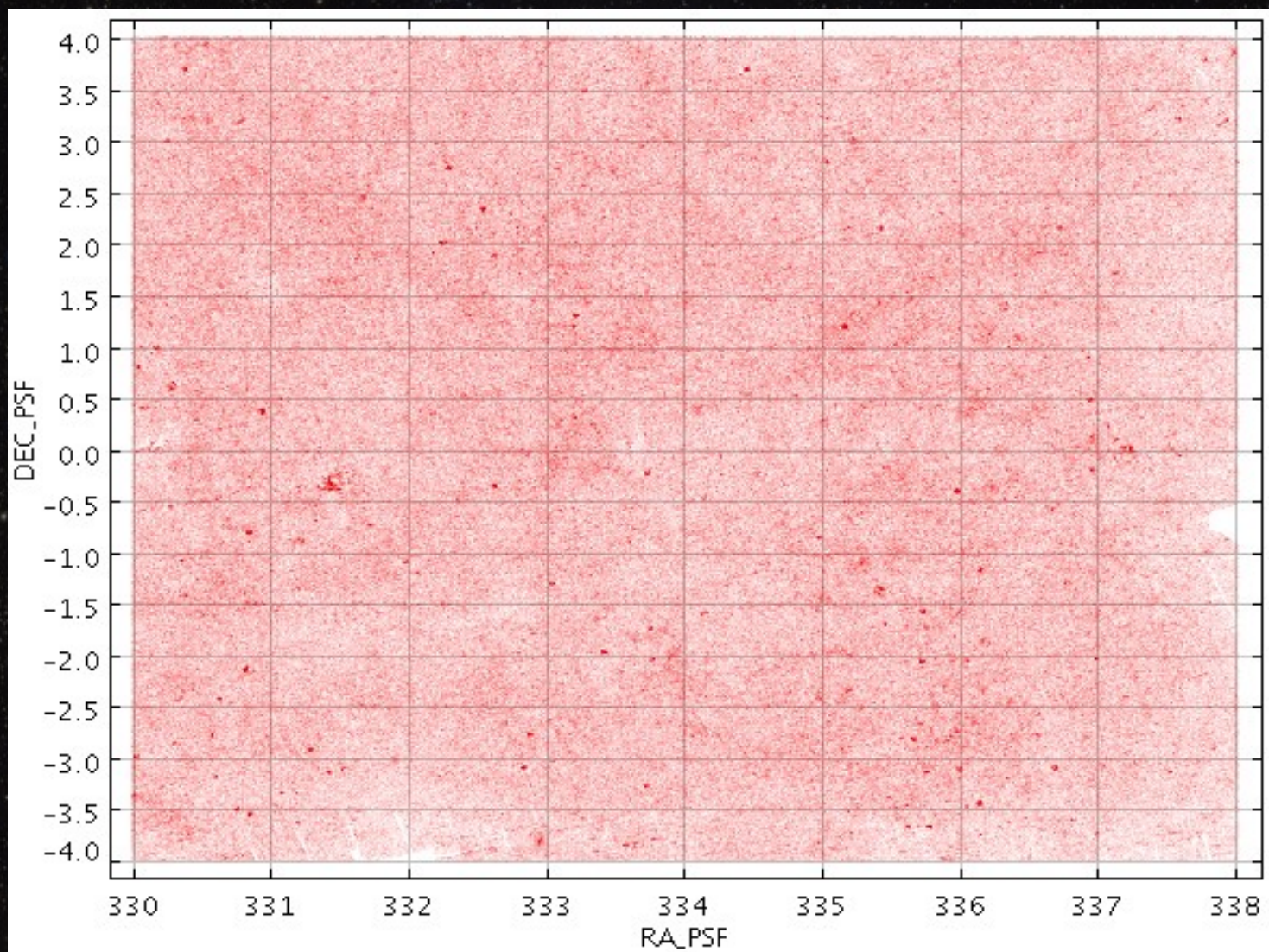
Depth



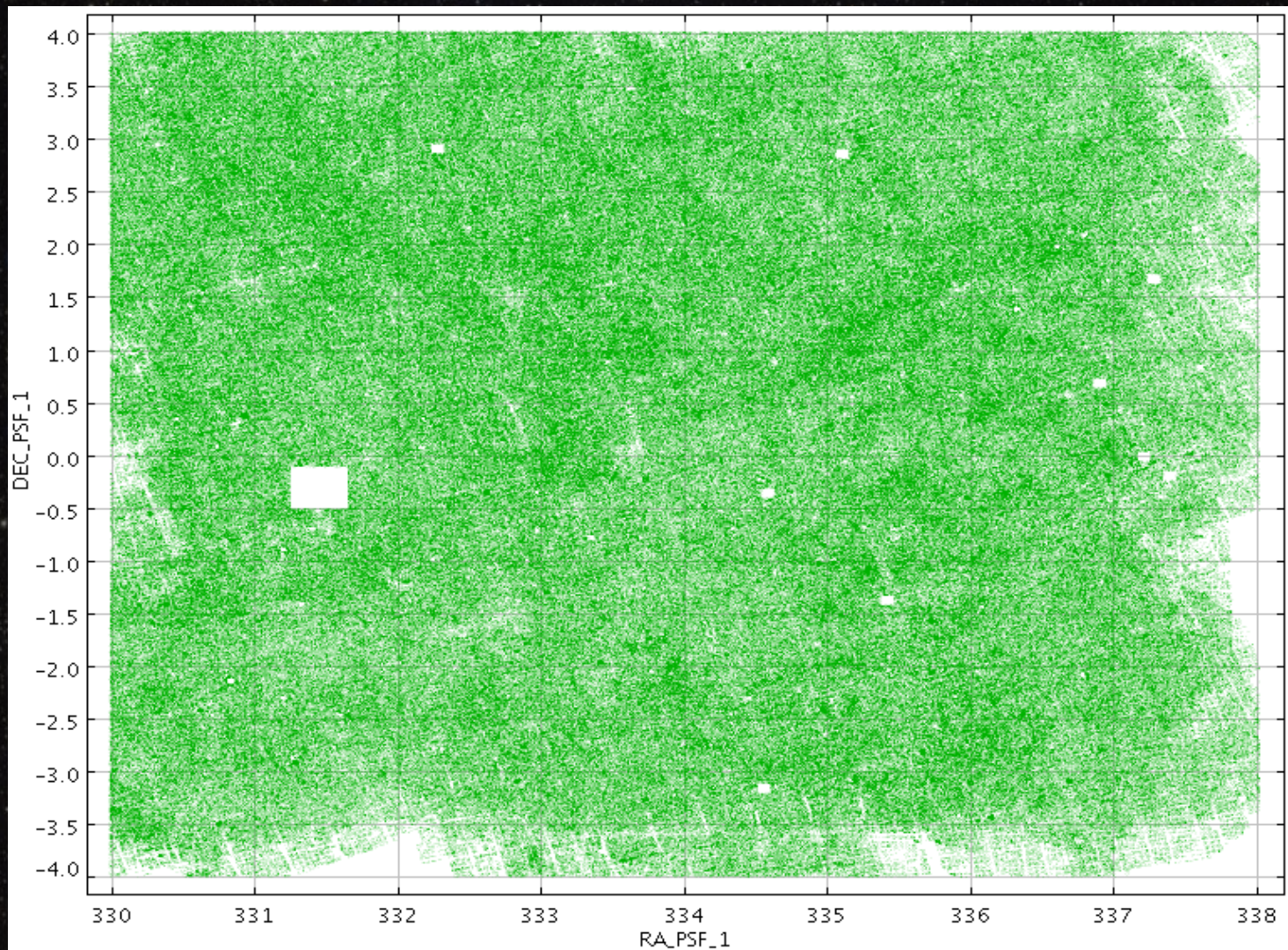
Credit: Nigel Metcalfe

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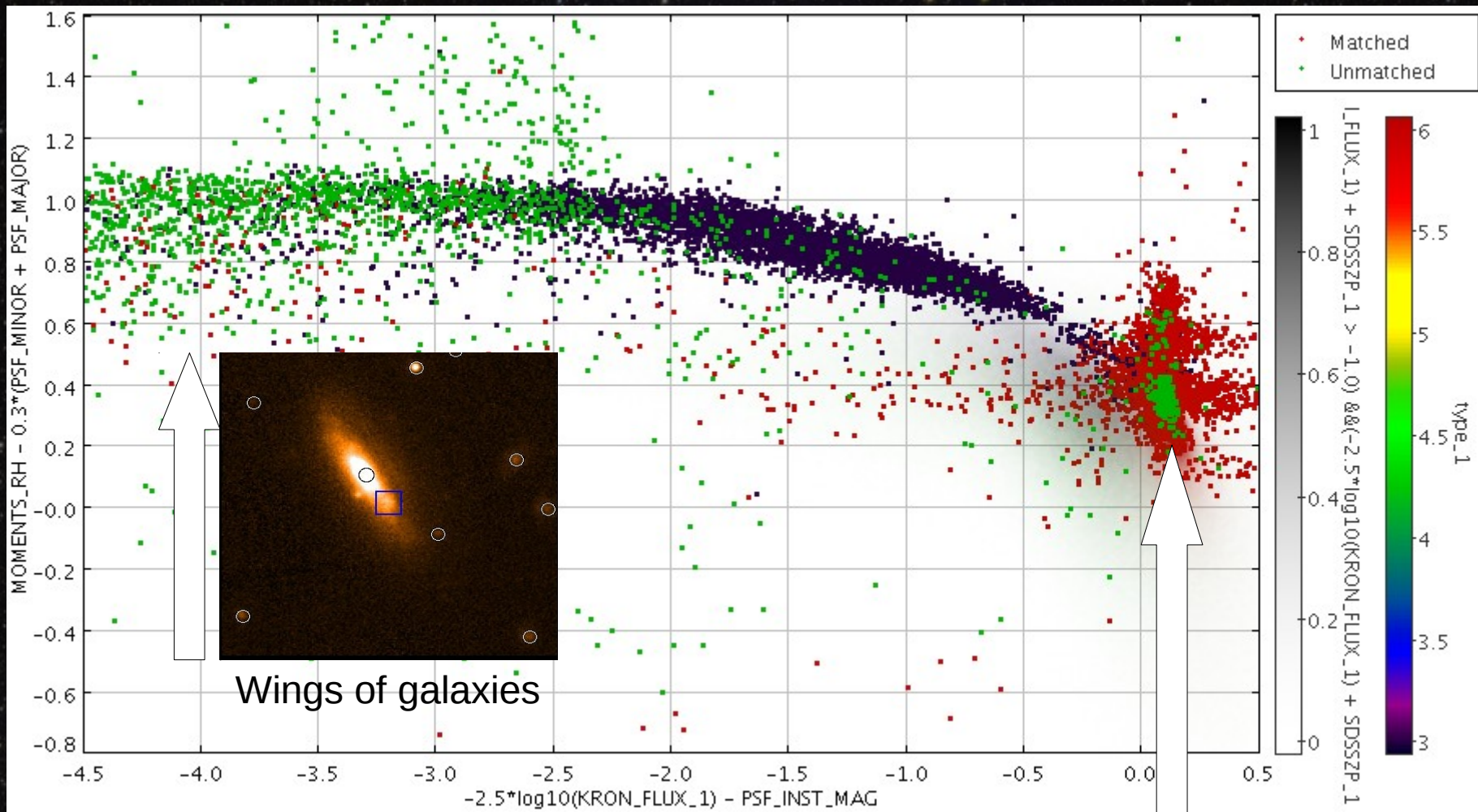
Sources



Sources in our Mask



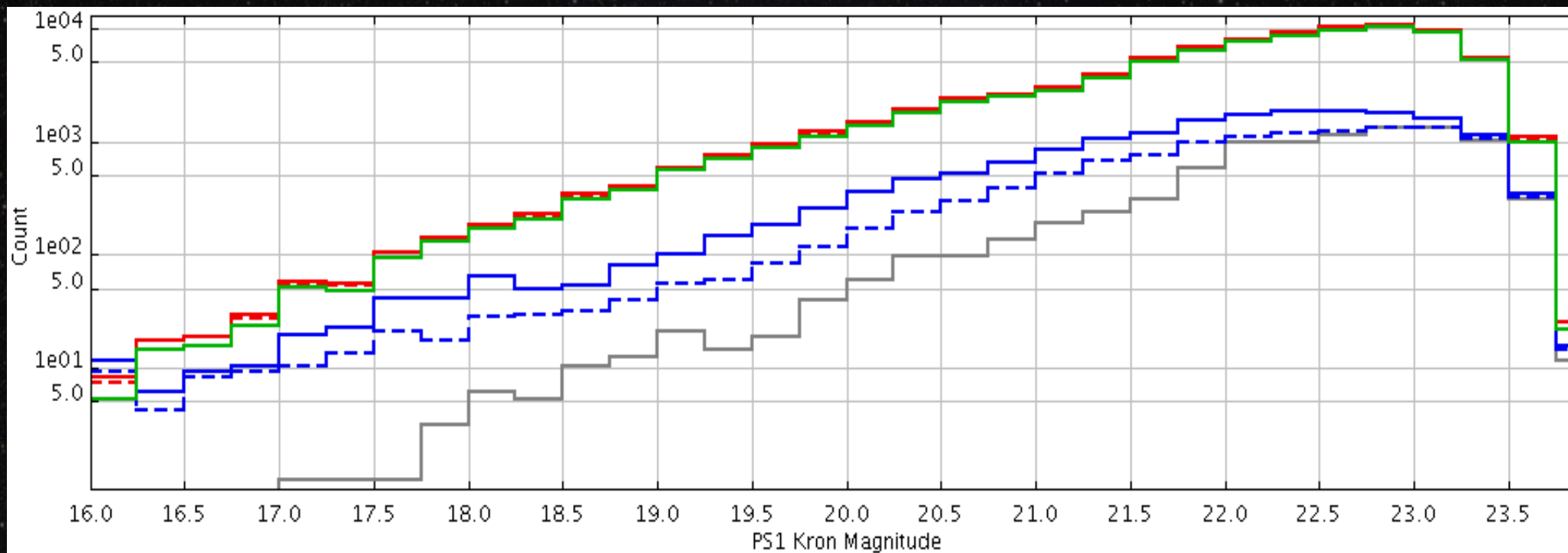
Star/Galaxy Separation



$r < 18$

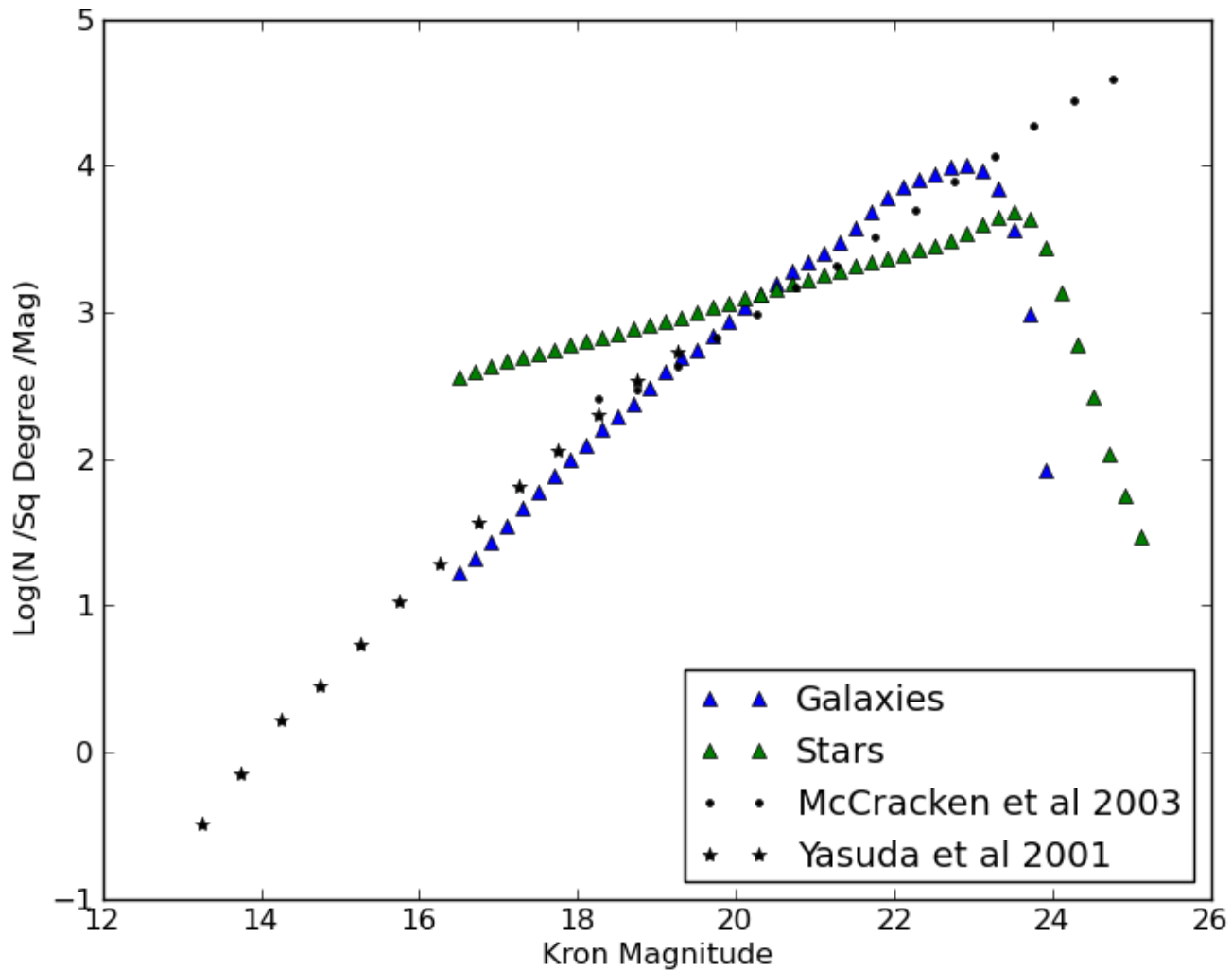
Real objects in DR8 mask

False Positives Rate

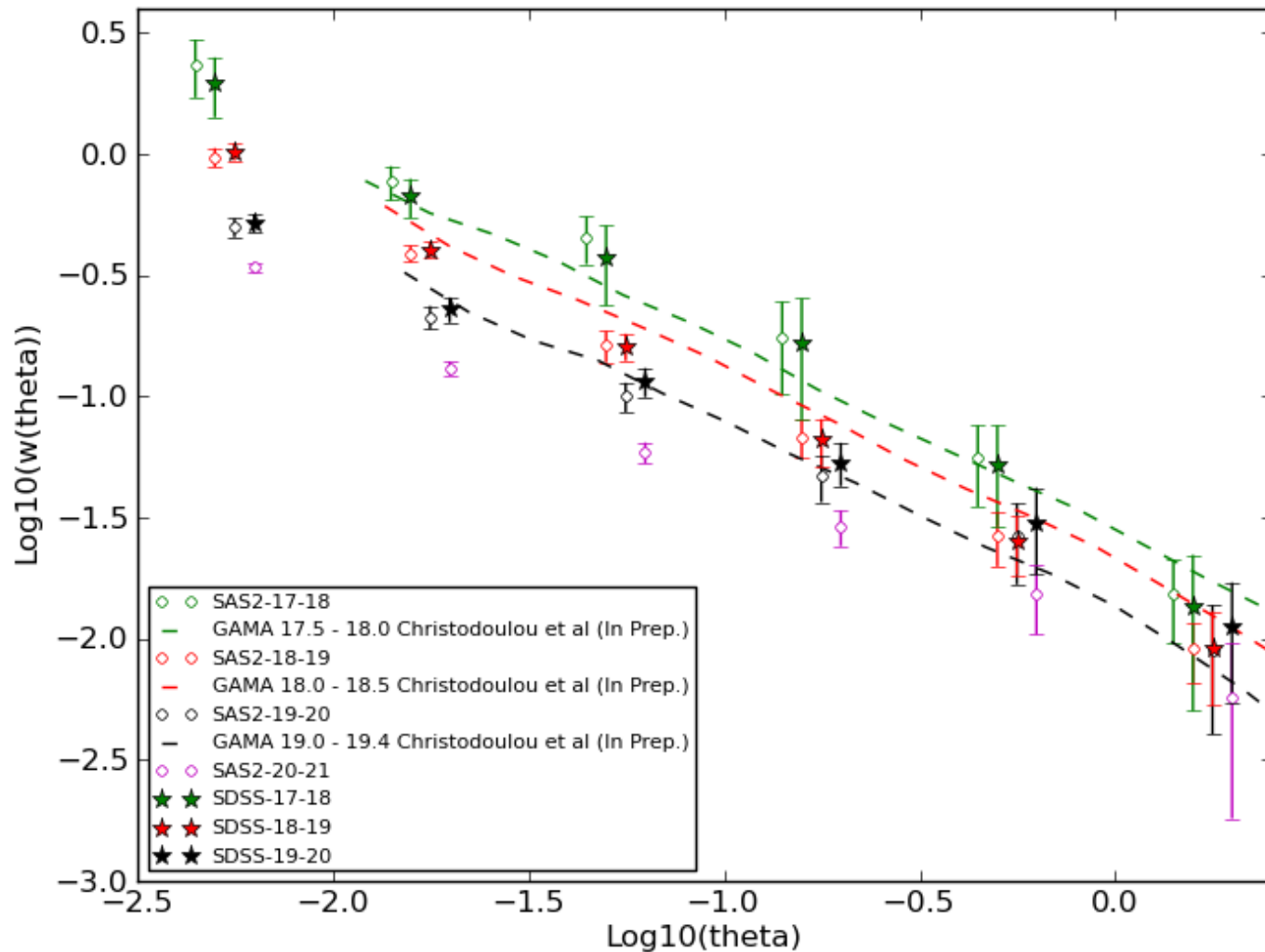


Red – Matched Galaxies
Red Dashed – Basic deblending
Blue – Unmatched Galaxies
Blue Dashed – Basic deblending
Grey – Unmatched with 2D cuts in s/g plot
Green – Matched with 2D cuts in s/g plot

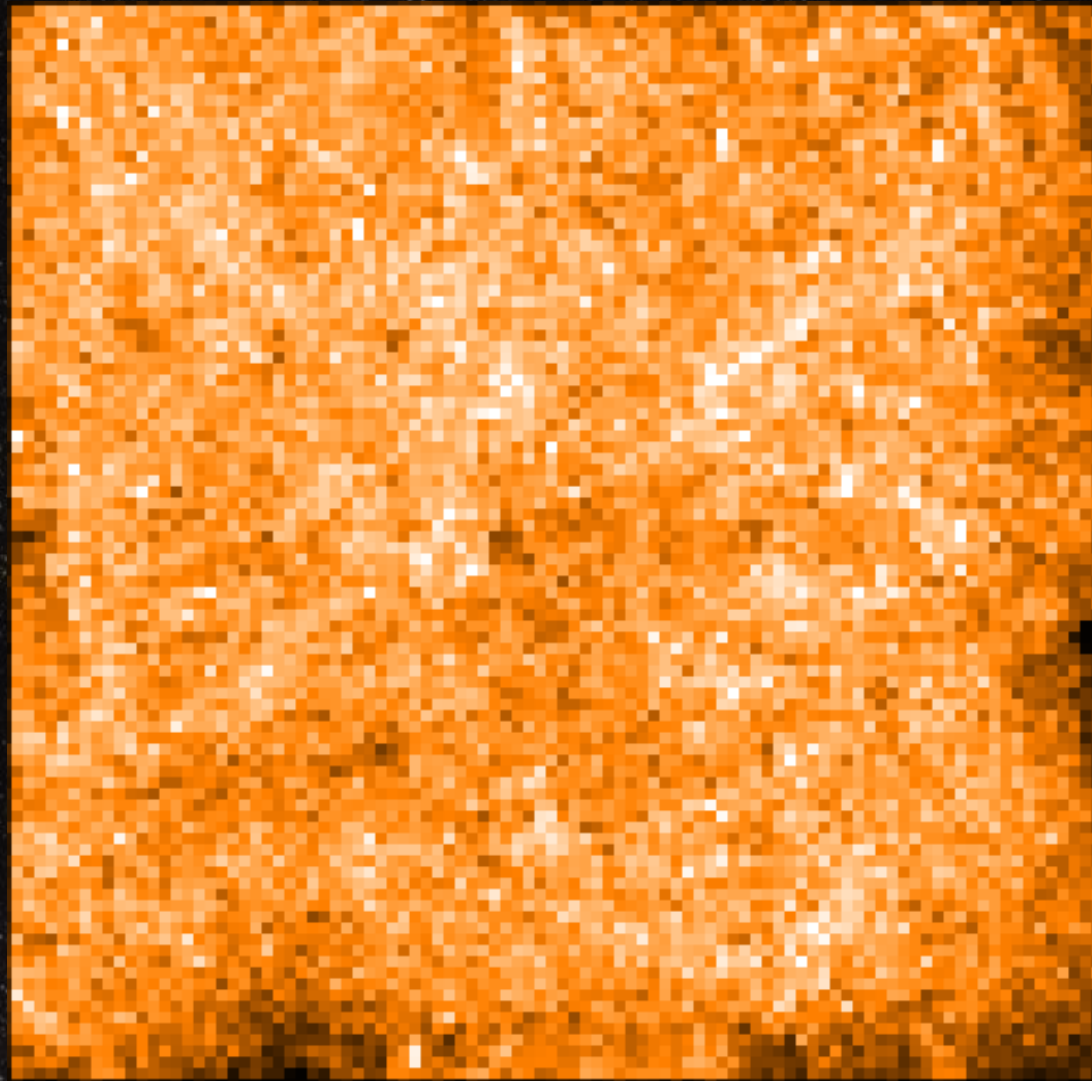
Number Counts



Clustering Results

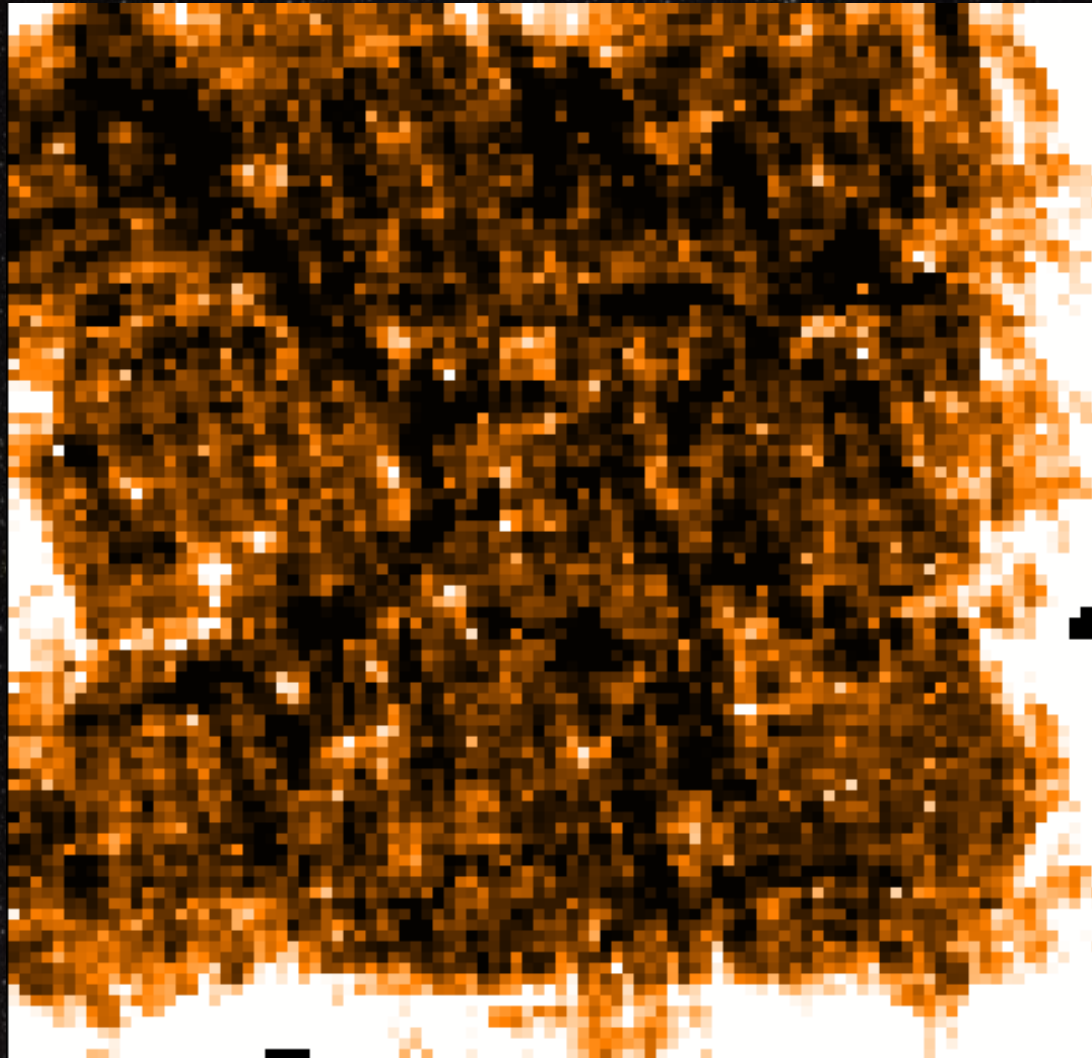


Variance Maps vs. Detections



$22 < \text{mag} < 23$

Variance Maps Vs. Detections



Conclusions

- Pan-STARRS is on track to deliver 5 band photometry over most of the sky
- We've developed star/galaxy separation & masks
- We've measured angular clustering
- Need to add detection efficiency effects to the randoms
- Scale our work to the whole 3π

