

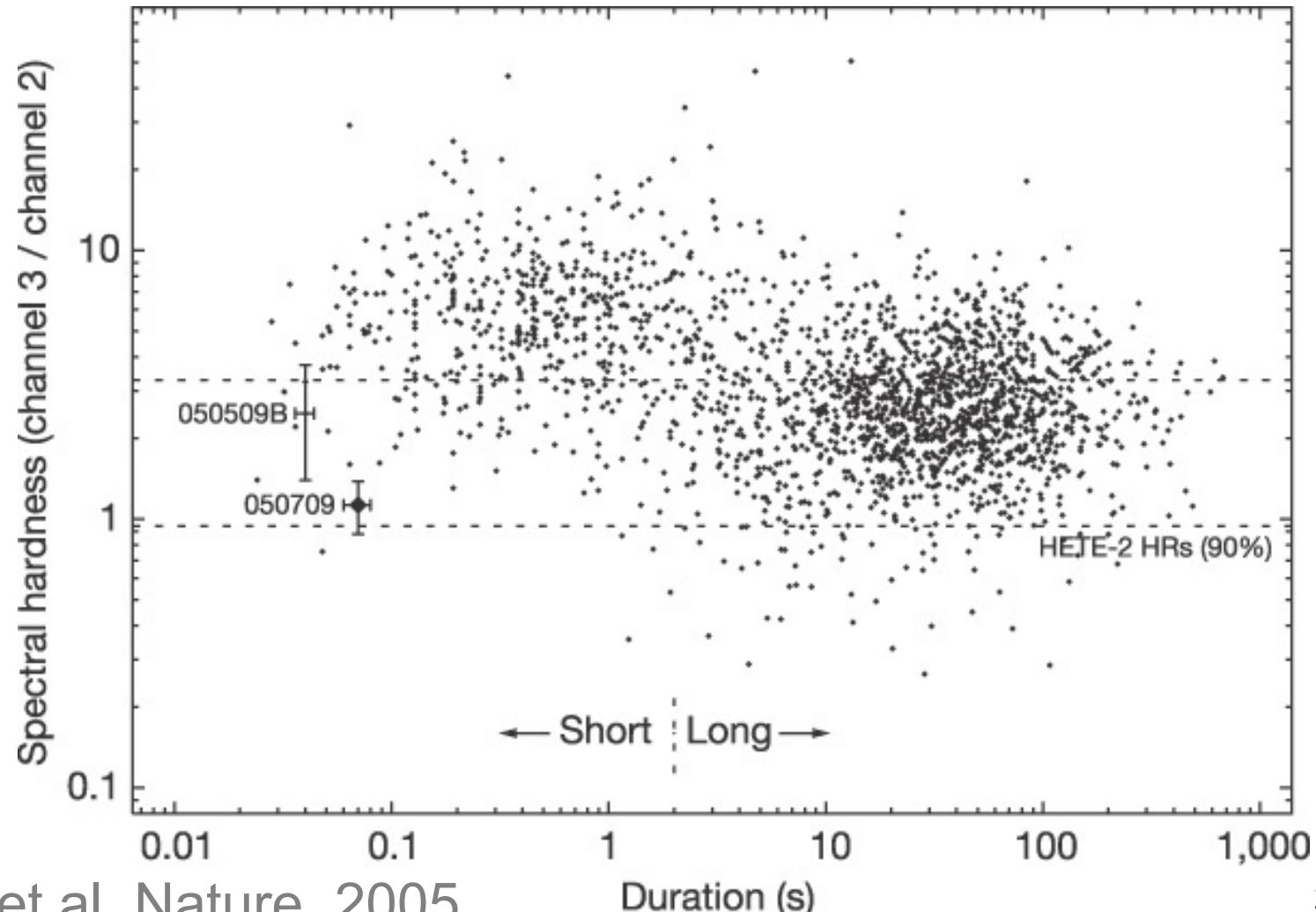
# Searches for Gravitational Wave Signals Associated to Gamma Ray Bursts

Stephen Fairhurst

Cardiff University  
for the LIGO Scientific Collaboration  
and the Virgo Collaboration

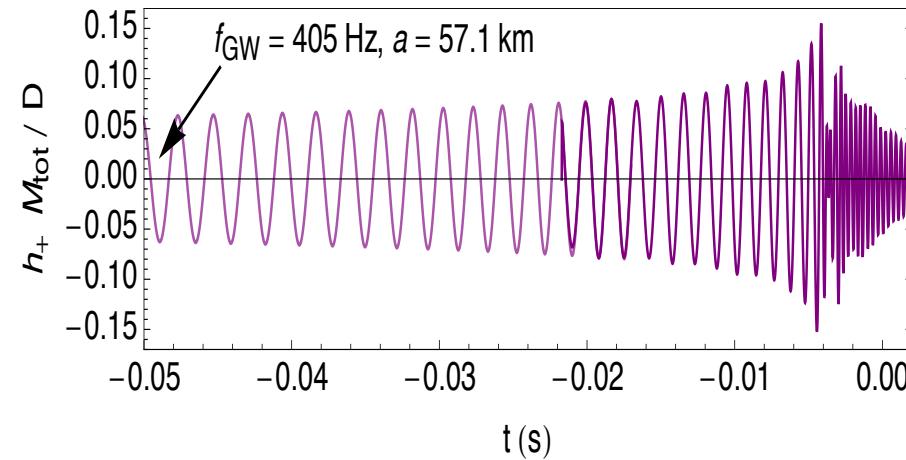


# Gamma Ray Burst Classification



# Gravitational Wave Emission

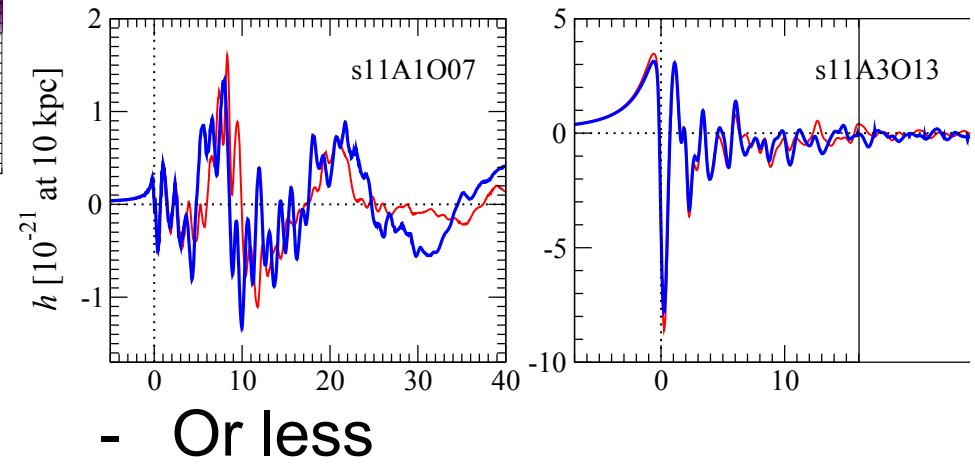
Short GRB: Merger of Neutron Star-Neutron Star or Neutron Star-Black Hole known GW emission



J Read/YITP (In Prep)

Long GRB: Collapsar uncertain GW emission

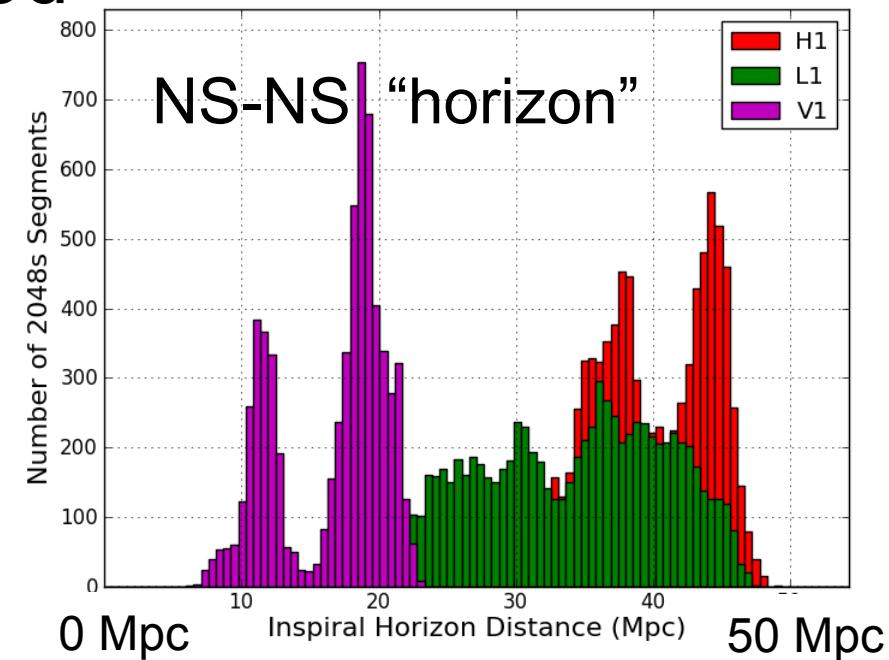
- Emission from  $10^{-2} M_{\odot} c^2$   
Piro and Pfahl 2007,  
Davies et al 2002.
- To  $10^{-8} M_{\odot} c^2$  Ott CGQ (2009)



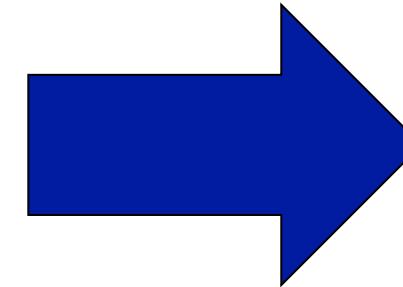
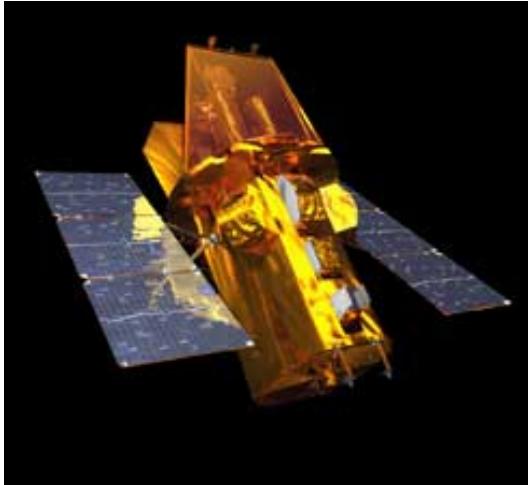
- Or less

# Gravitational Wave Detectors

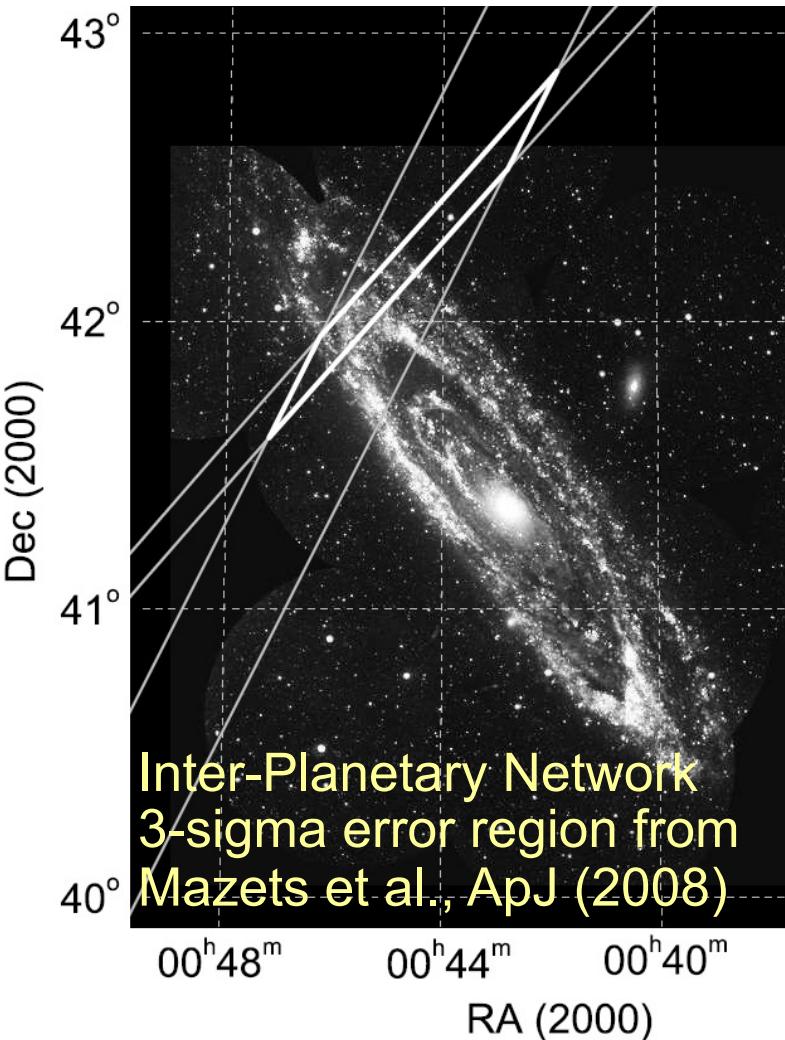
- Operated from Nov 2005 to Sept 2007 & July 2009 to Oct 2010
- Currently being upgraded
- From 2015, advanced detectors
  - 10x increase in (distance) sensitivity



# Triggered Search



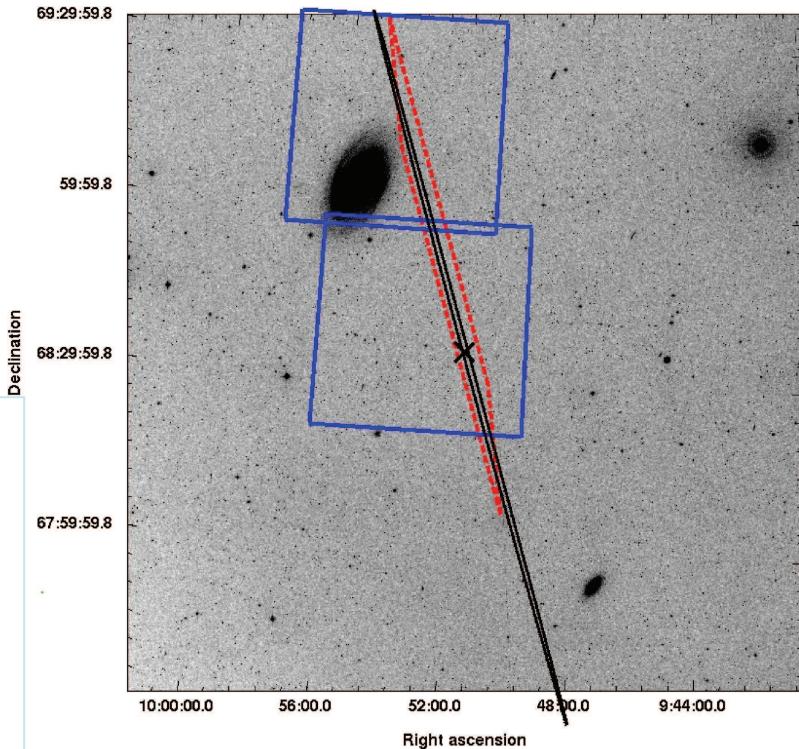
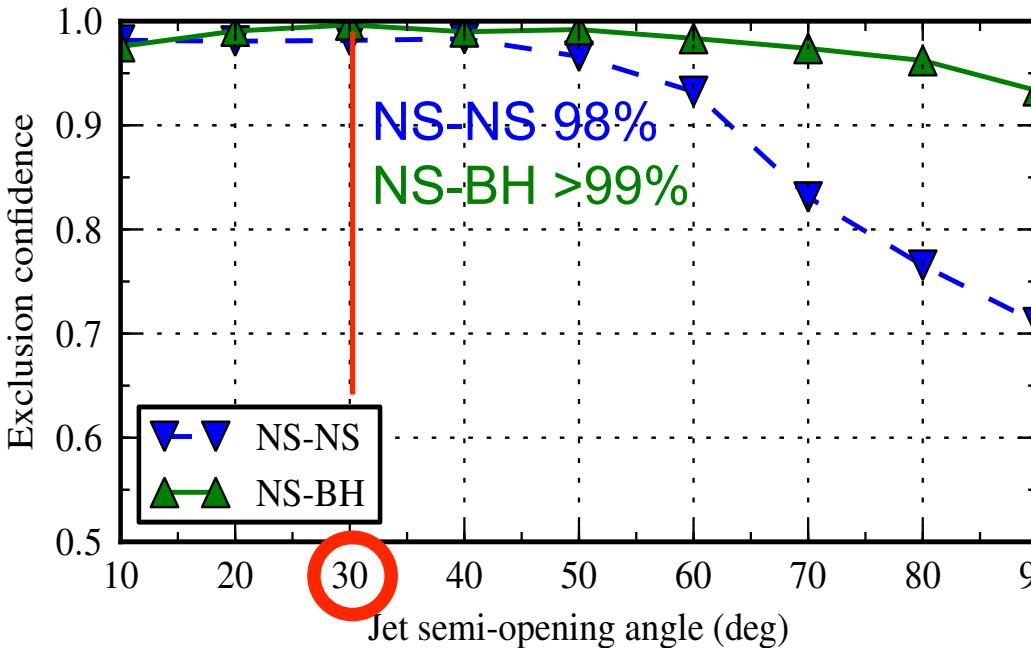
# GRB 070201



- Localization overlaps M31 (at 770 kpc)
  - No GW signal observed
  - Exclude NS-NS and NS-BH merger in M31 with 99% confidence
  - Indirect support for hypothesis of soft gamma repeater in M31
- Abbott et al. ApJ 2008

# GRB 051103

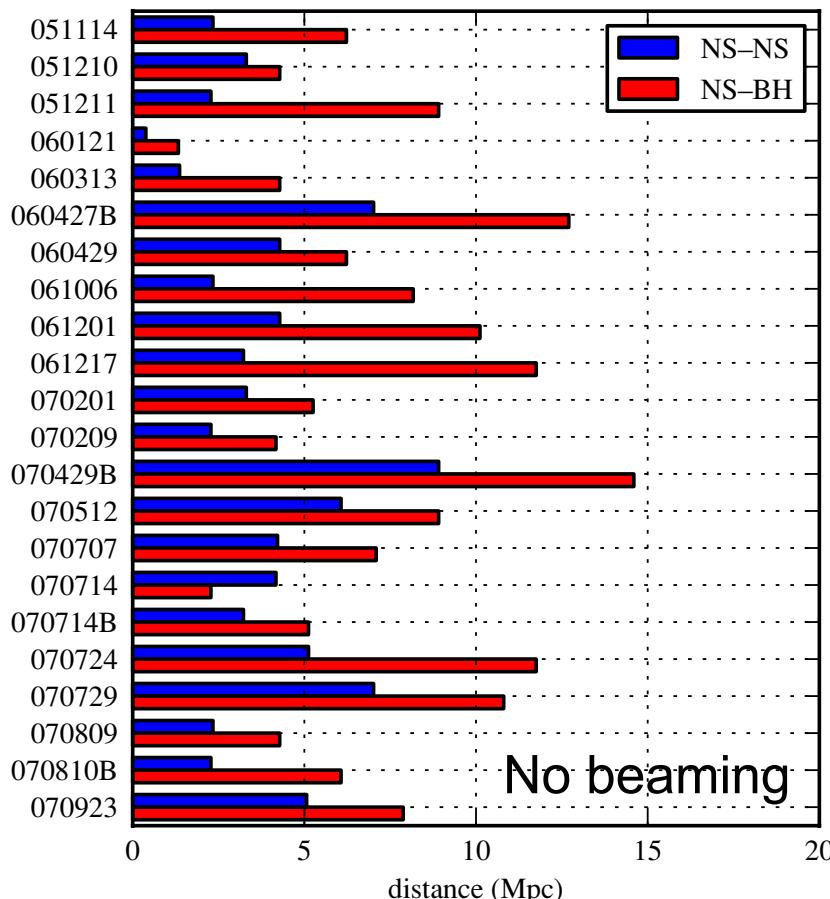
- Localization overlaps M81 (at 3.6 Mpc)
- No GW signal observed
- Exclude binary merger progenitor as function of opening angle



Abadie et al. 1201.1163

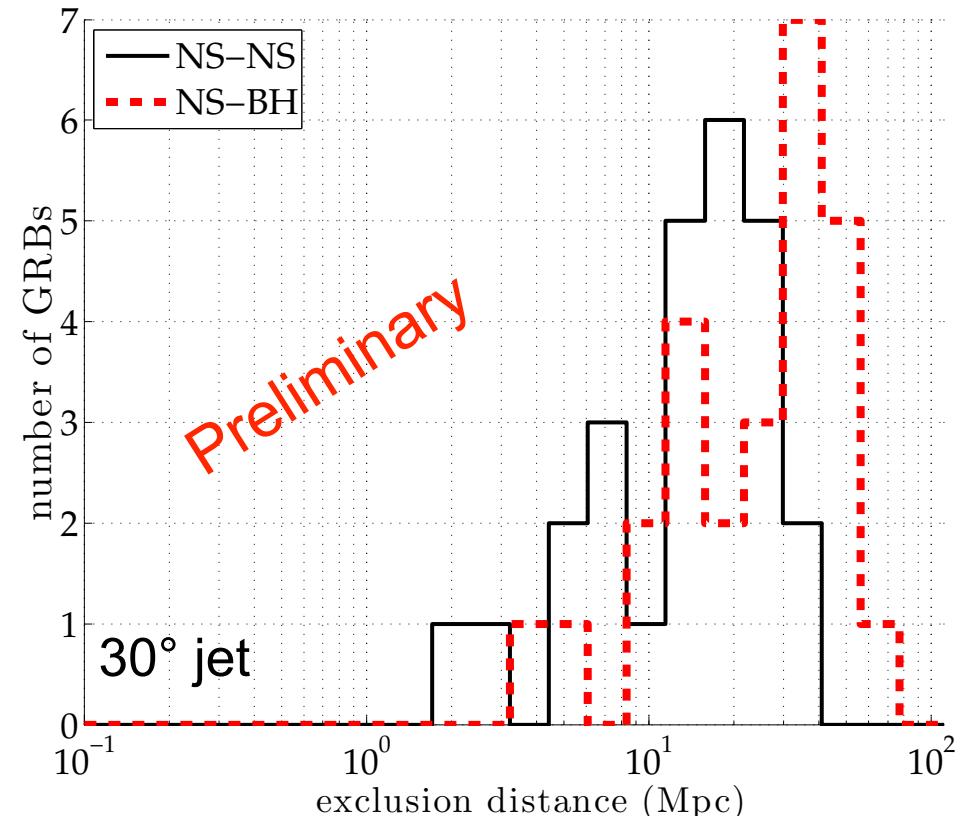
# Short GRB exclusion distances

2005-7 run



Binary Merger Model

2009-10 run



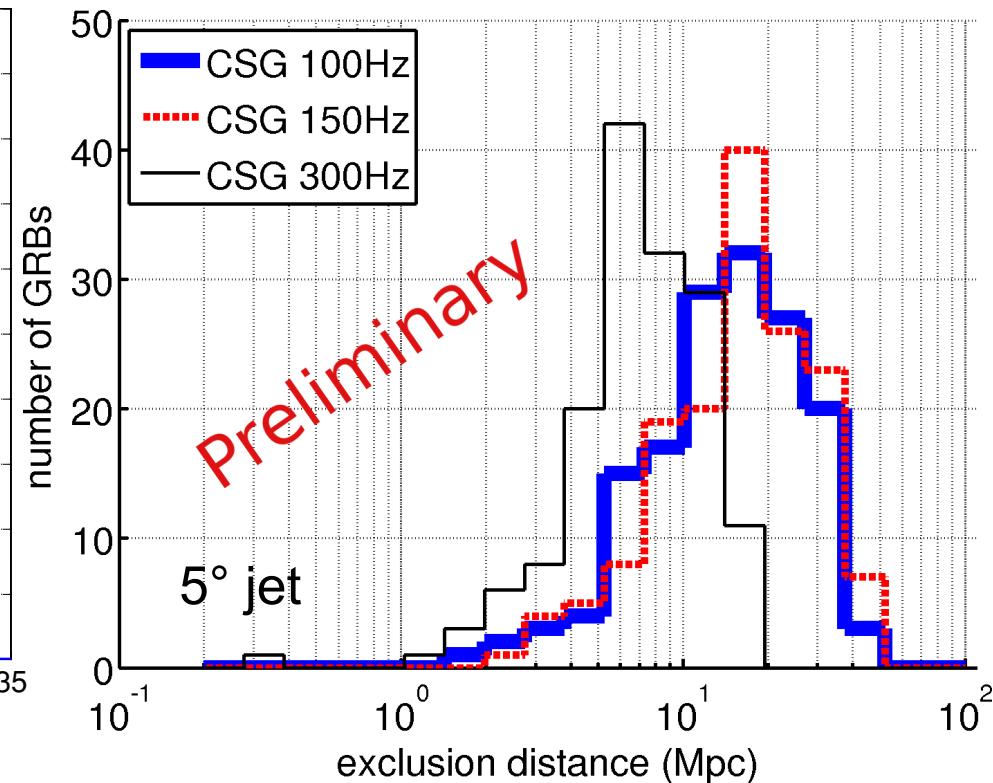
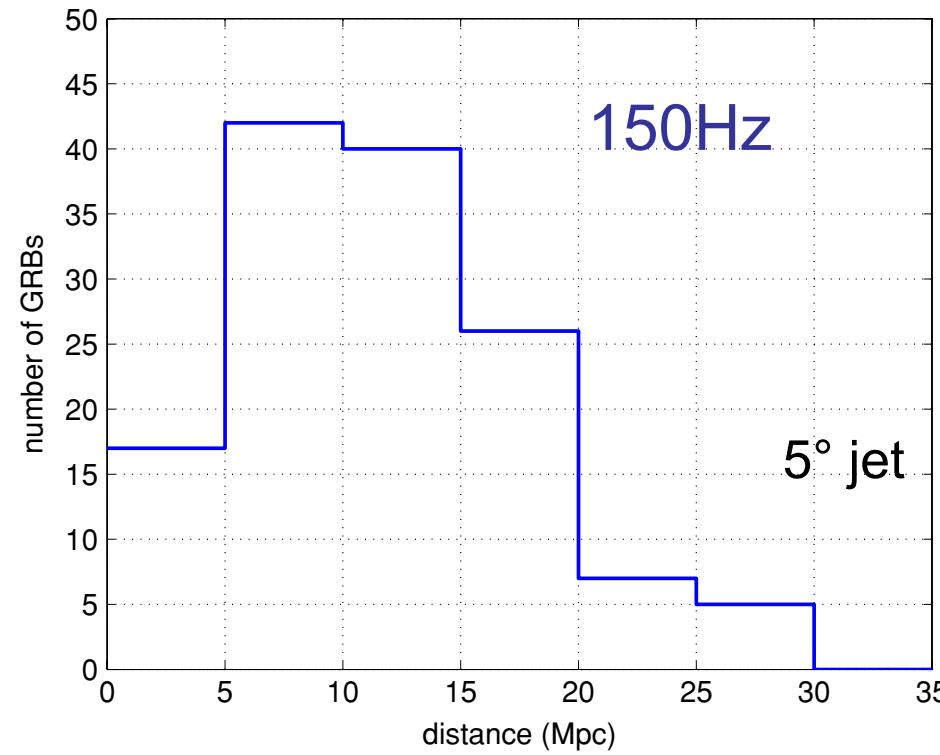
# Exclusion distances for all GRBs

## Unmodelled GW Burst

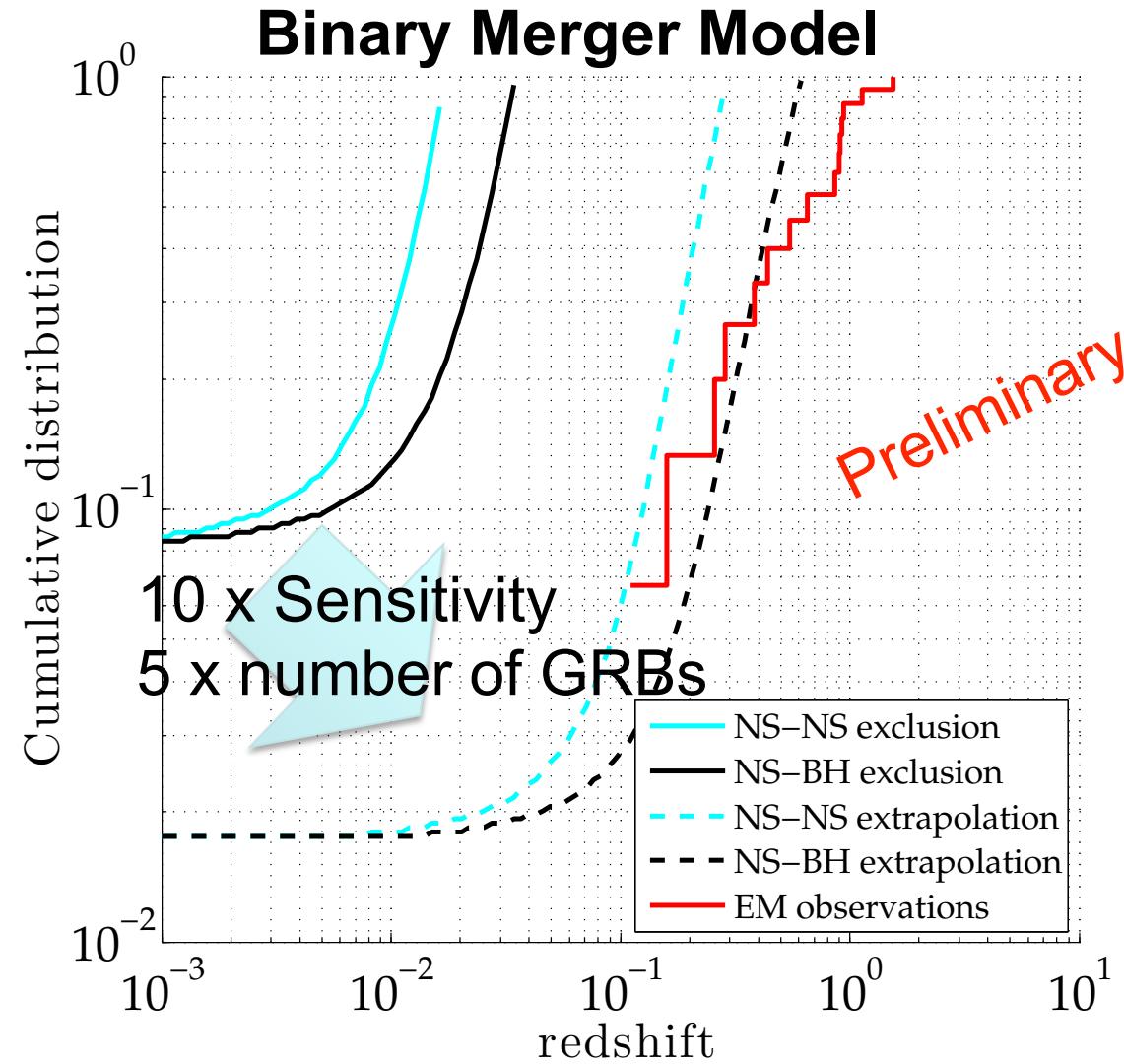
Assuming  $10^{-2} M_{\odot}c^2$  in GW;  
emitted in small frequency band

2005-7 run

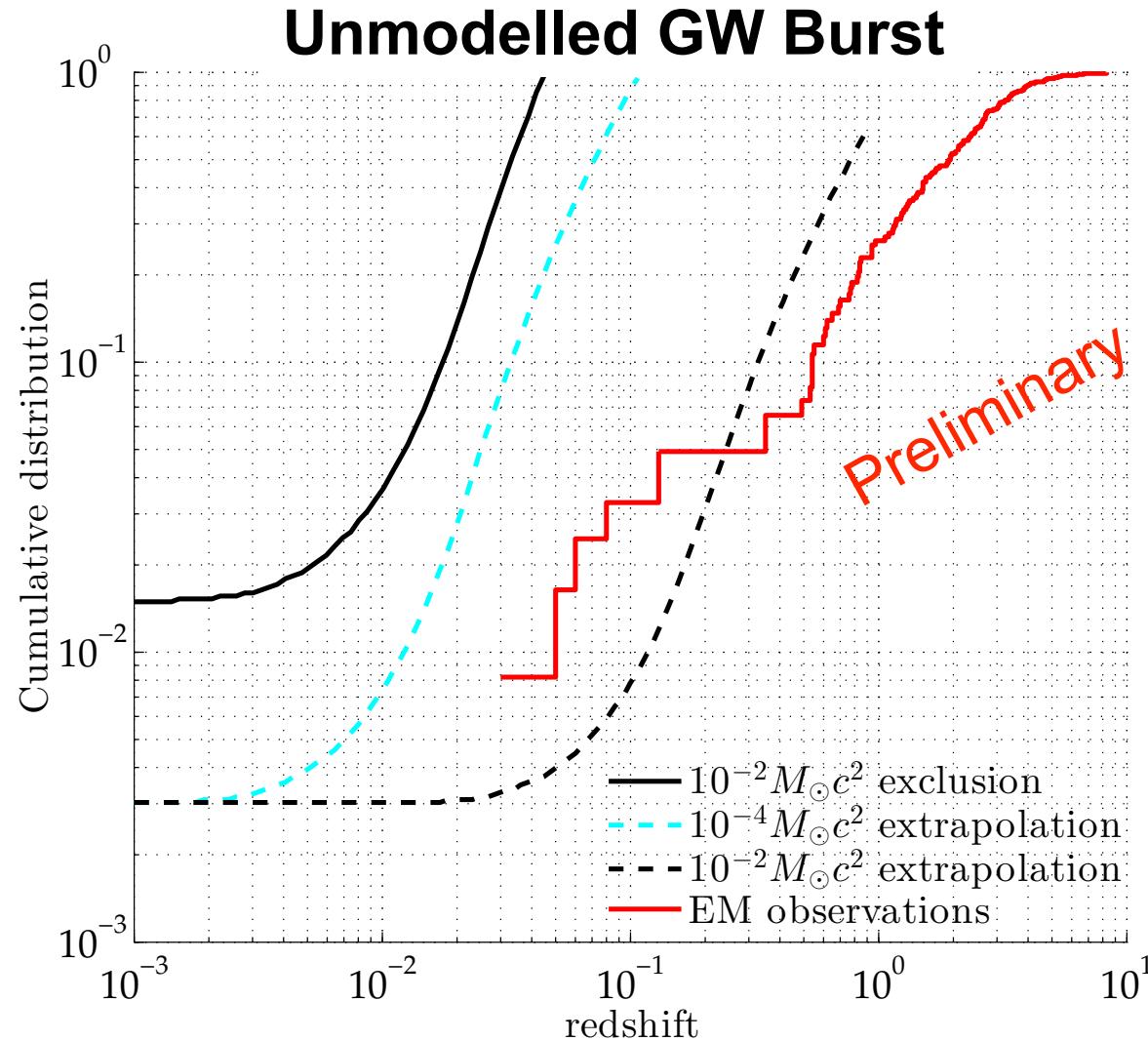
2009-10 run



# Future Prospects



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- Advanced detectors may detect GW associated to GRBs
  - Confirm (or rule out) progenitor models



- Prospects are strongly dependent on number of GRBs observed electromagnetically.