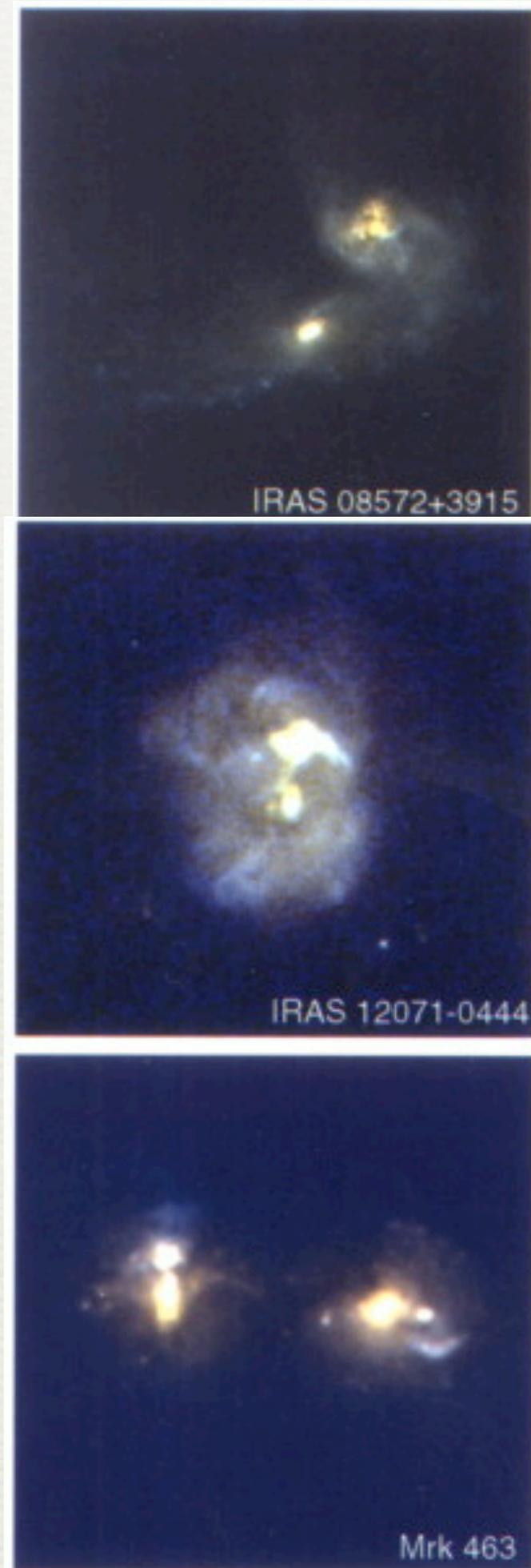
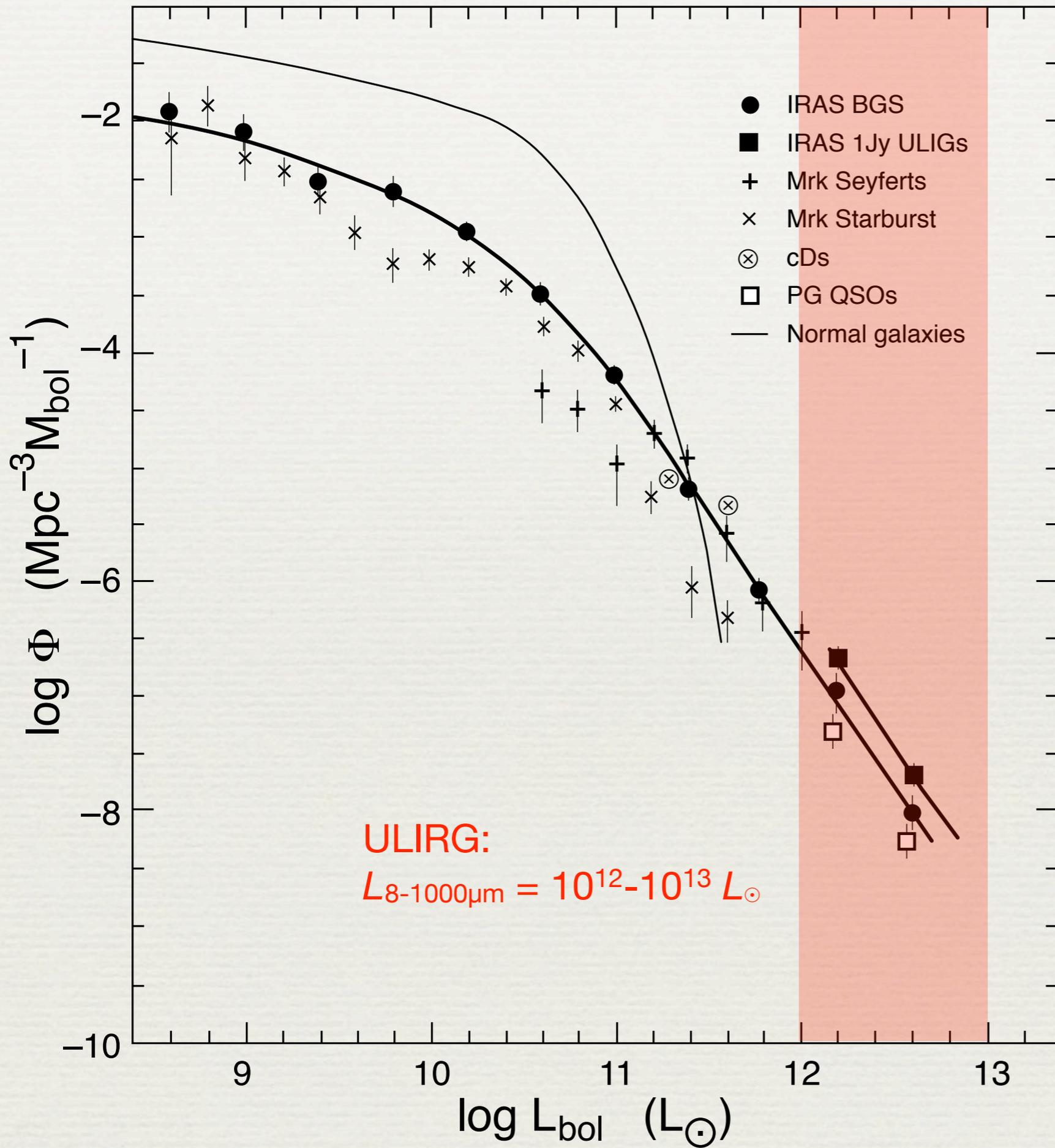




# Ionized vs. molecular gas feedback in local ULIRGs

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Mark Westmoquette  
ESO Garching



# VLT/VIMOS-IFU

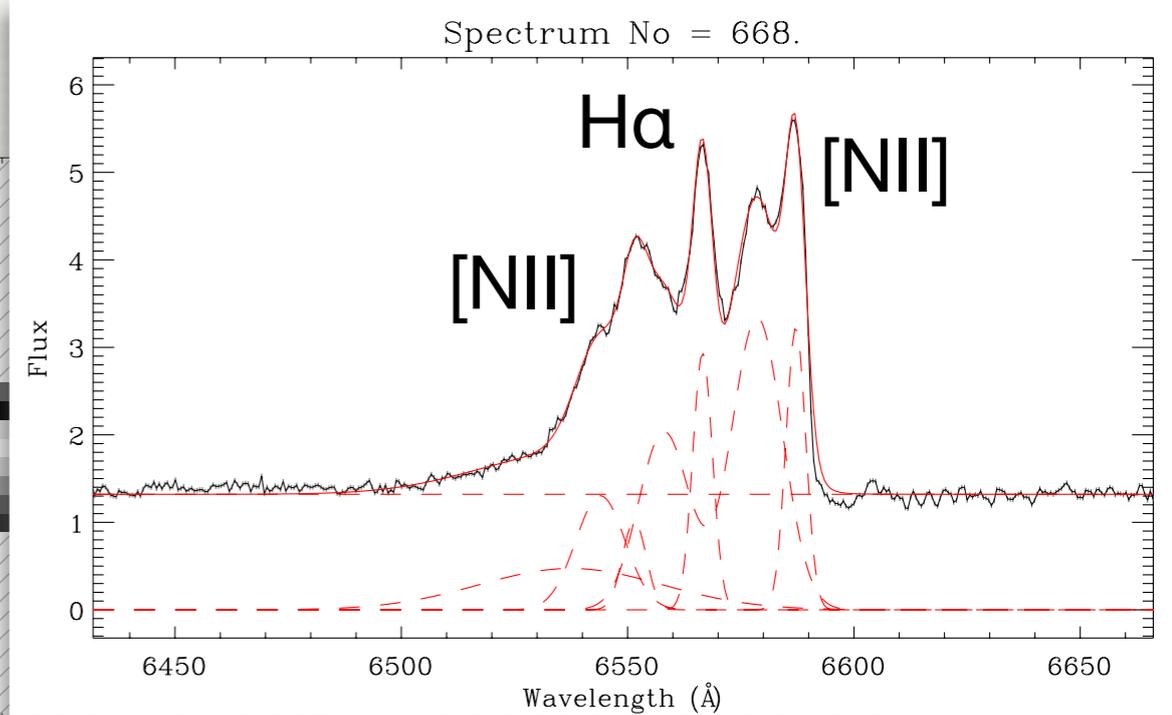
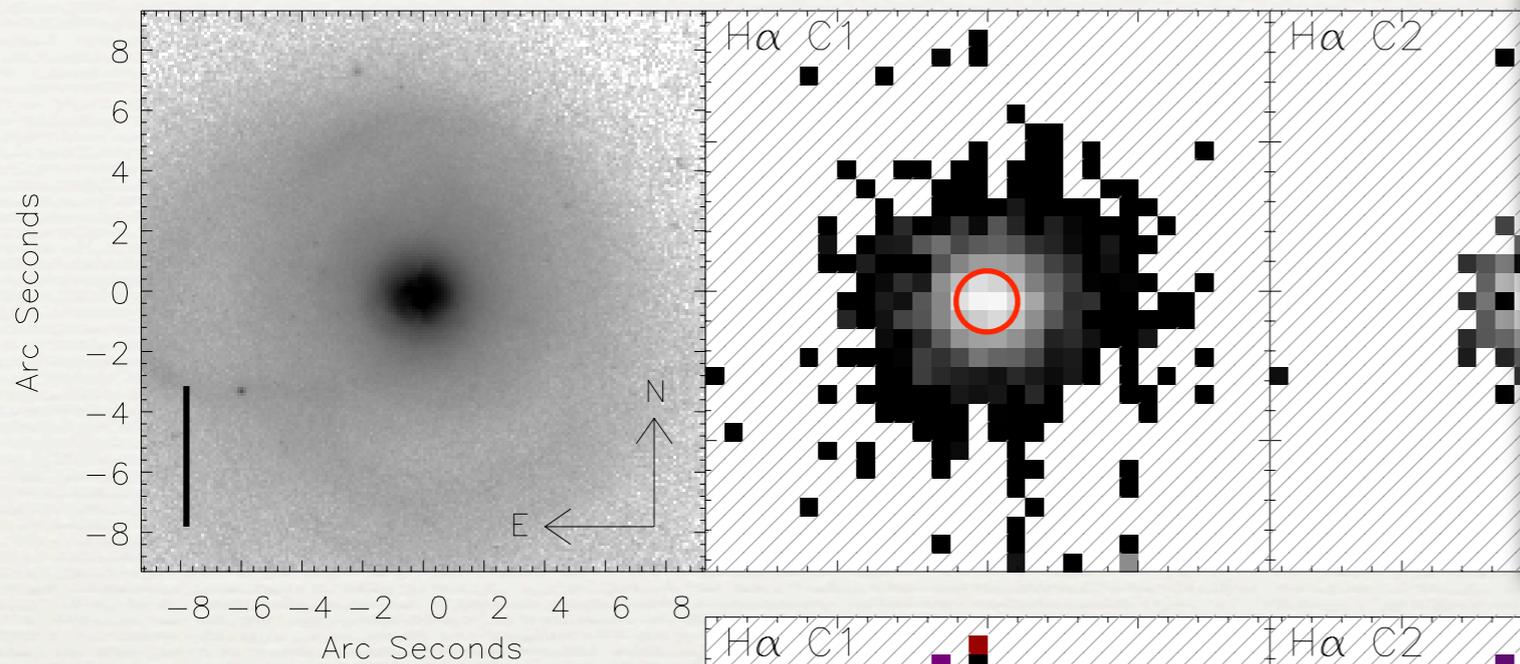
Galaxy	$\alpha$ (J2000), $\delta$ (J2000)	$z^a$	$D^a$ (Mpc)	$\log(L_{\text{IR}}/L_{\odot})^b$	Sp type <sup>c</sup>
IRAS 00198-7926	00 21 52.9, -79 10 08.0	0.0728	300	12.01	Sy2 (AGN in northern nucleus <sup>e</sup> )
IRAS 00335-2732	00 36 00.5, -27 15 34.5	0.0693	282	11.93	SB
IRAS 03068-5346	03 08 21.3, -53 35 12.0	0.0745	304	11.96	SB
IRAS 05189-2524	05 21 50.5, -25 21 46.3	0.0426	173	12.19	Sy2
IRAS 06035-7102	06 05 11.2, -71 03 05.5	0.0795	325	12.26	SB (outflow suggests AGN <sup>e</sup> )
IRAS 09111-1007W	09 13 38.8, -10 19 20.3	0.0541	220	11.98	SB+Sy2
IRAS 12112+0305	12 13 46.0, +02 48 38.0	0.0733	300	12.28	LINER
IRAS 14348-1447	14 37 38.4, -15 00 22.8	0.0827	345	12.30	LINER (SW nucleus Sy2 <sup>f</sup> )
IRAS 14378-3651	14 40 58.9, -37 04 33.0	0.0676	280	12.24	LINER+Sy2
IRAS 17208-0014	17 23 21.9, -00 17 00.4	0.0428	175	12.25	SB (LINER <sup>g</sup> )
IRAS 19254-7245	19 31 21.6, -72 39 21.7	0.0617	250	11.91	Sy2
IRAS 19297-0406	19 32 20.7, -04 00 06.0	0.0857	353	12.36	SB
IRAS 20046-0623	20 07 19.4, -06 14 26.0	0.0844	348	11.97	SB
IRAS 20414-1651	20 44 18.2, -44 33 37.7	0.0871	358	12.26	SB
IRAS 20551-4250	20 58 27.4, -42 39 03.1	0.0428	175	11.98	Sy2 <sup>h</sup>
IRAS 21504-0628	21 53 05.5, -06 14 49.9	0.0776	320	11.92	Sy2?
IRAS 22491-1808	22 51 49.3, -17 52 23.4	0.0778	320	12.09	SB
IRAS 23128-5919	23 15 47.0, -59 03 16.9	0.0446	182	11.96	SB+Sy2

# VLT/VIMOS-IFU

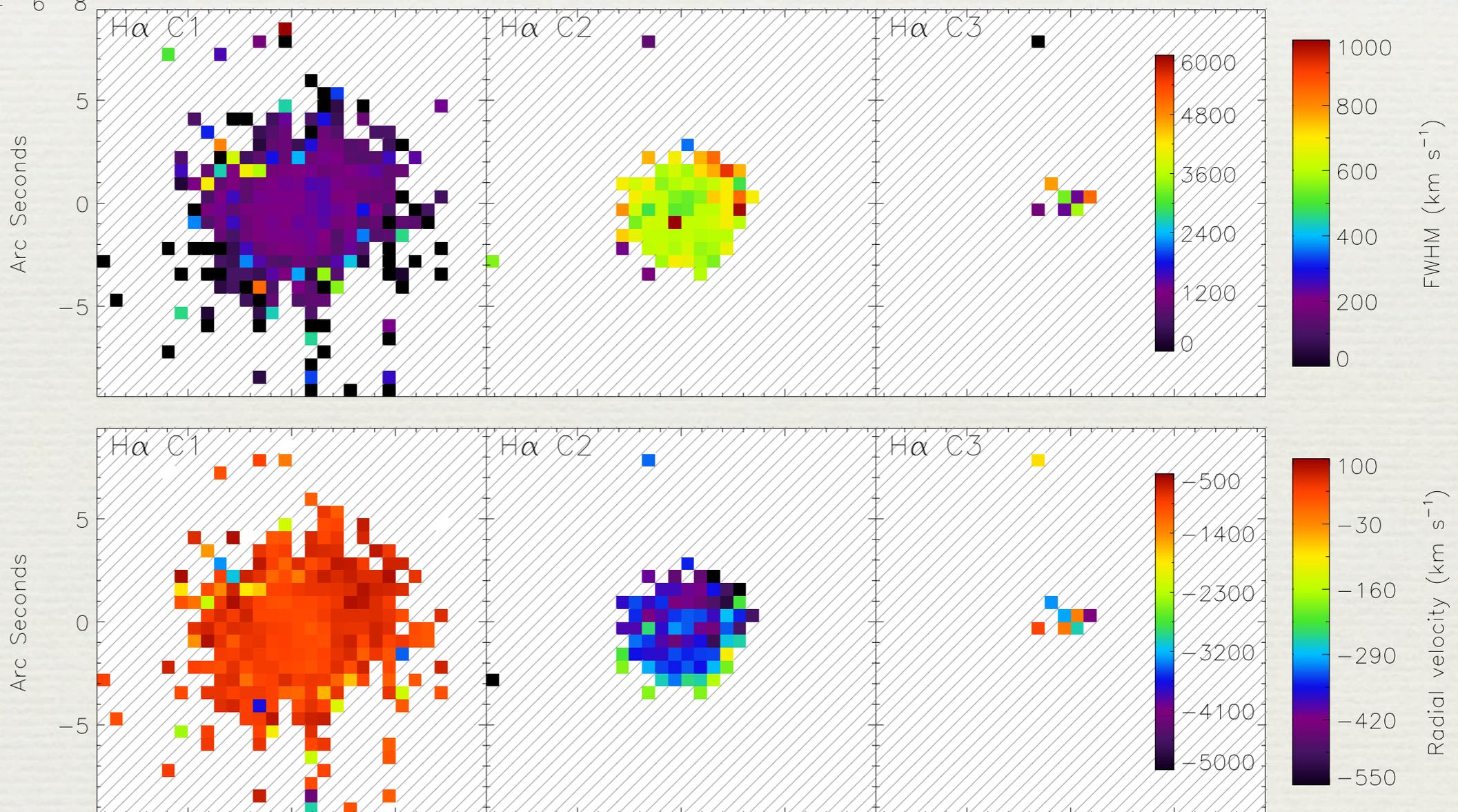
Galaxy	$\alpha$ (J2000), $\delta$ (J2000)	$z^a$	$D^a$ (Mpc)	$\log(L_{\text{IR}}/L_{\odot})^b$	Sp type <sup>c</sup>
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IRAS 03068-5346	03 08 21.3, -53 35 12.0	0.0745	304	11.96	SB
<b>IRAS 05189-2524</b>	<b>05 21 50.5, -25 21 46.3</b>	<b>0.0426</b>	<b>173</b>	<b>12.19</b>	<b>Sy2</b>
IRAS 06035-7102	06 05 11.2, -71 03 05.5	0.0795	325	12.26	SB (outflow suggests AGN <sup>e</sup> )
IRAS 09111-1007W	09 13 38.8, -10 19 20.3	0.0541	220	11.98	SB+Sy2
<b>IRAS 12112+0305</b>	<b>12 13 46.0, +02 48 38.0</b>	<b>0.0733</b>	<b>300</b>	<b>12.28</b>	<b>LINER</b>
IRAS 14348-1447	14 37 38.4, -15 00 22.8	0.0827	345	12.30	LINER (SW nucleus Sy2 <sup>f</sup> )
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IRAS 17208-0014	17 23 21.9, -00 17 00.4	0.0428	175	12.25	SB (LINER <sup>g</sup> )
<b>IRAS 19254-7245</b>	<b>19 31 21.6, -72 39 21.7</b>	<b>0.0617</b>	<b>250</b>	<b>11.91</b>	<b>Sy2</b>
IRAS 19297-0406	19 32 20.7, -04 00 06.0	0.0857	353	12.36	SB
IRAS 20046-0623	20 07 19.4, -06 14 26.0	0.0844	348	11.97	SB
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IRAS 20551-4250	20 58 27.4, -42 39 03.1	0.0428	175	11.98	Sy2 <sup>h</sup>
IRAS 21504-0628	21 53 05.5, -06 14 49.9	0.0776	320	11.92	Sy2?
IRAS 22491-1808	22 51 49.3, -17 52 23.4	0.0778	320	12.09	SB
IRAS 23128-5919	23 15 47.0, -59 03 16.9	0.0446	182	11.96	SB+Sy2



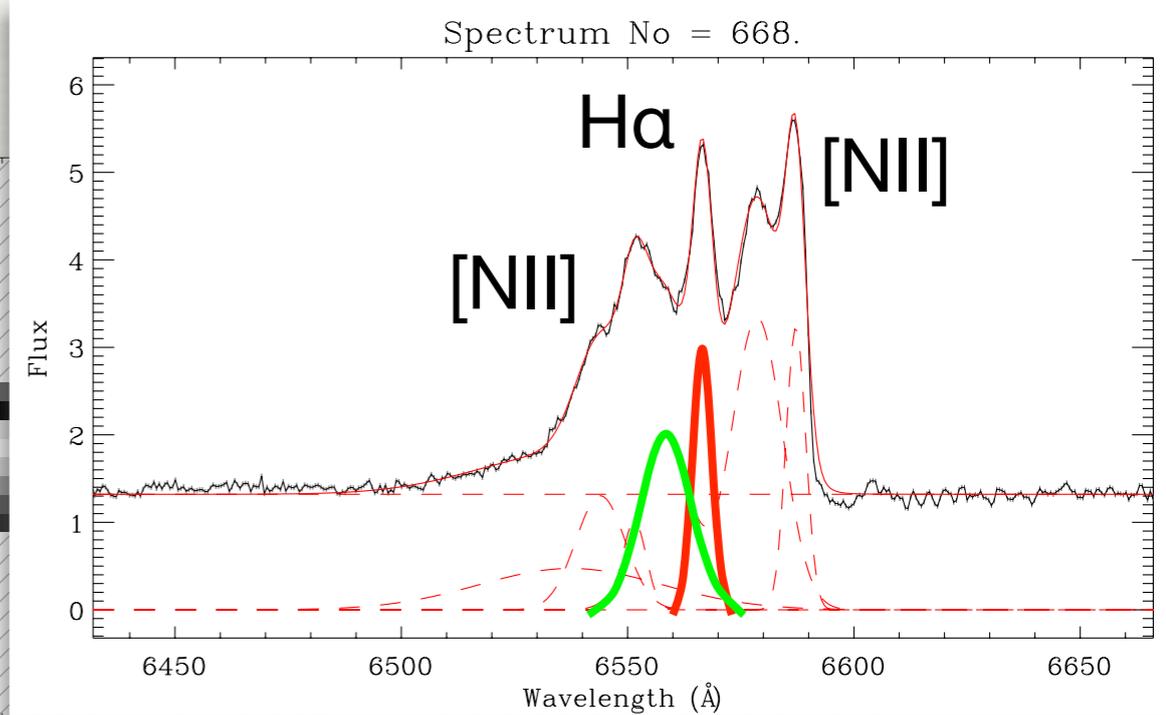
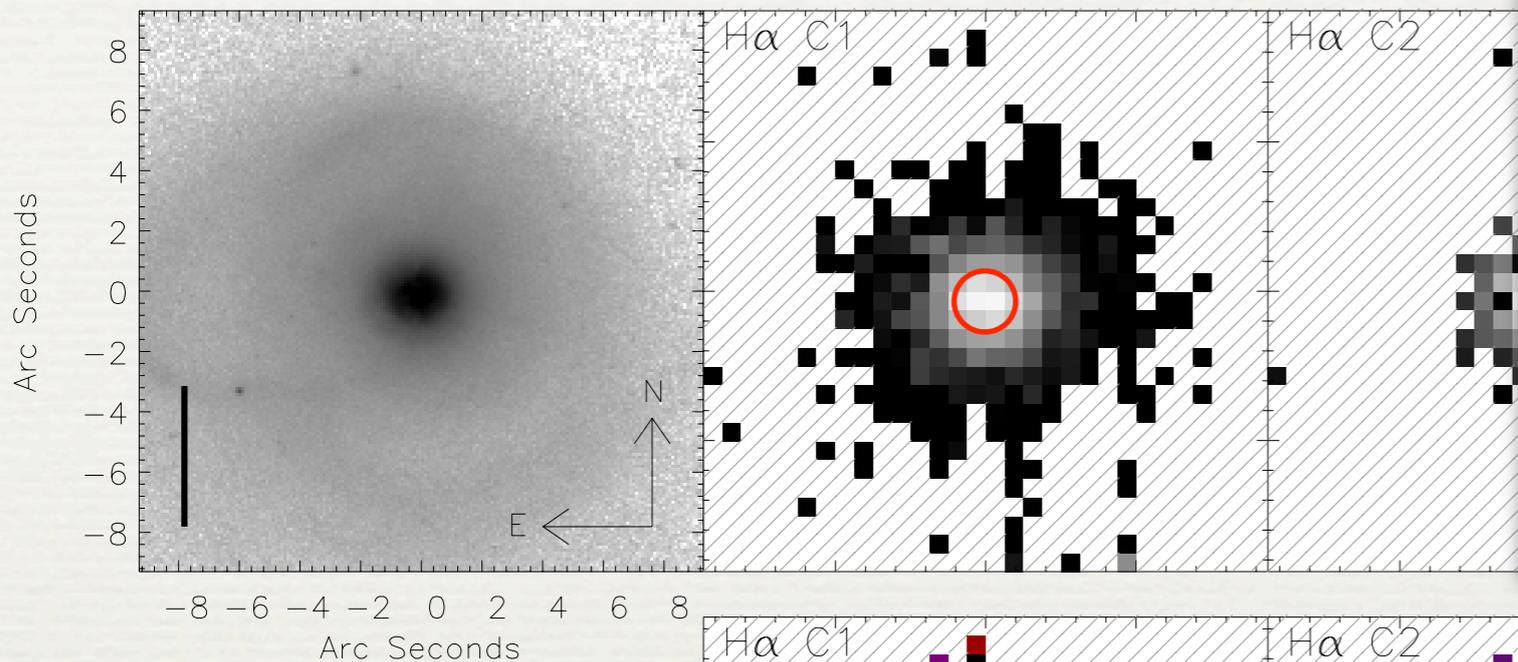
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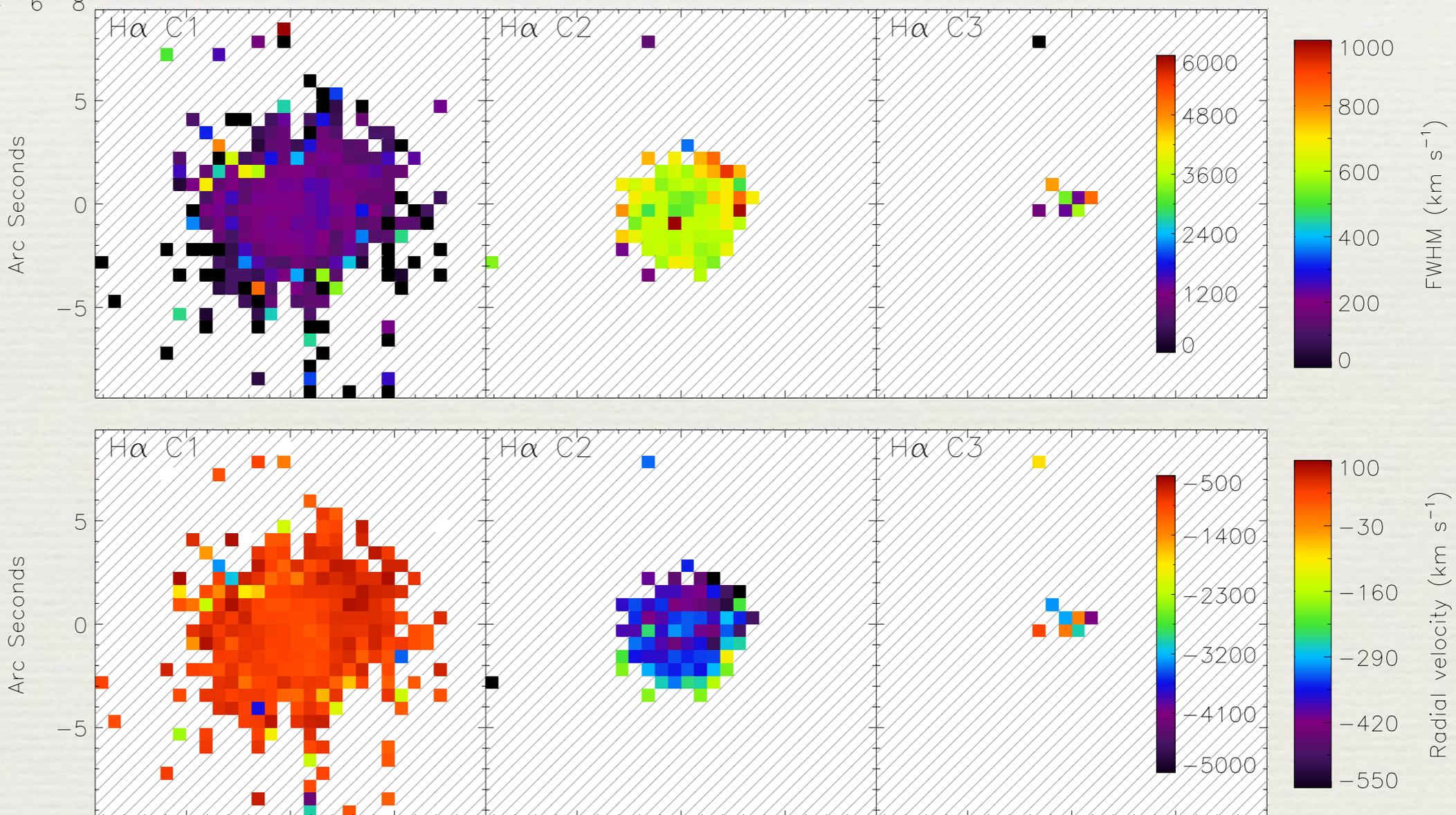
## H $\alpha$ Kinematics



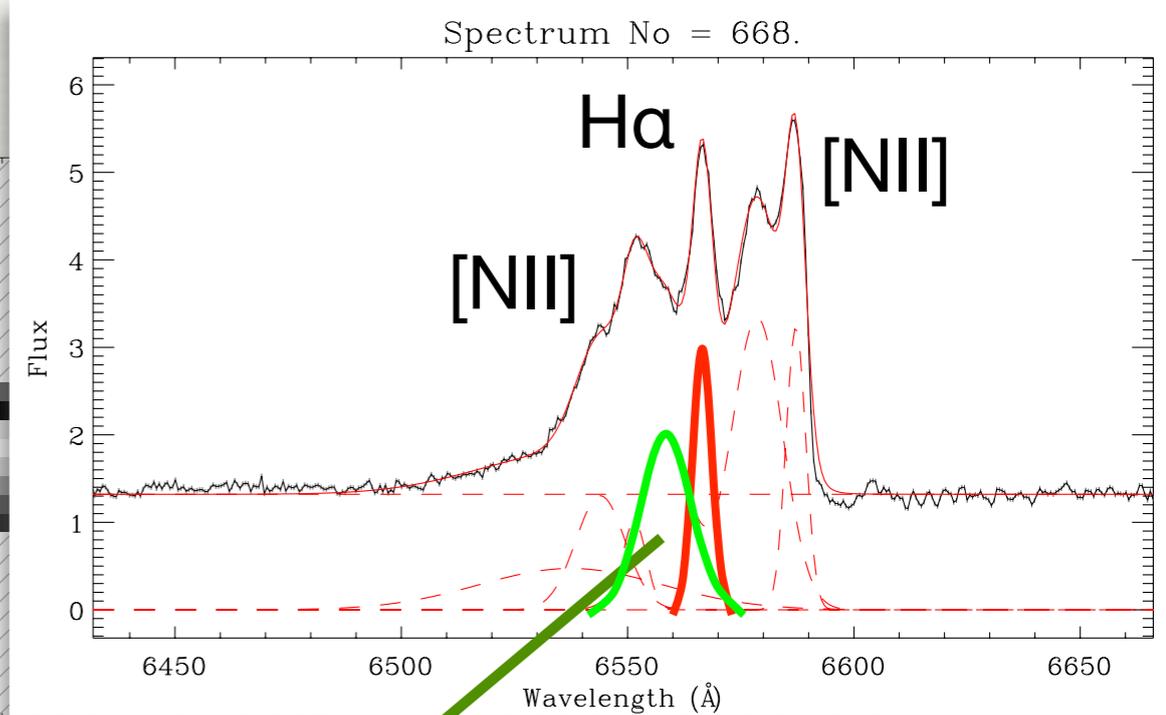
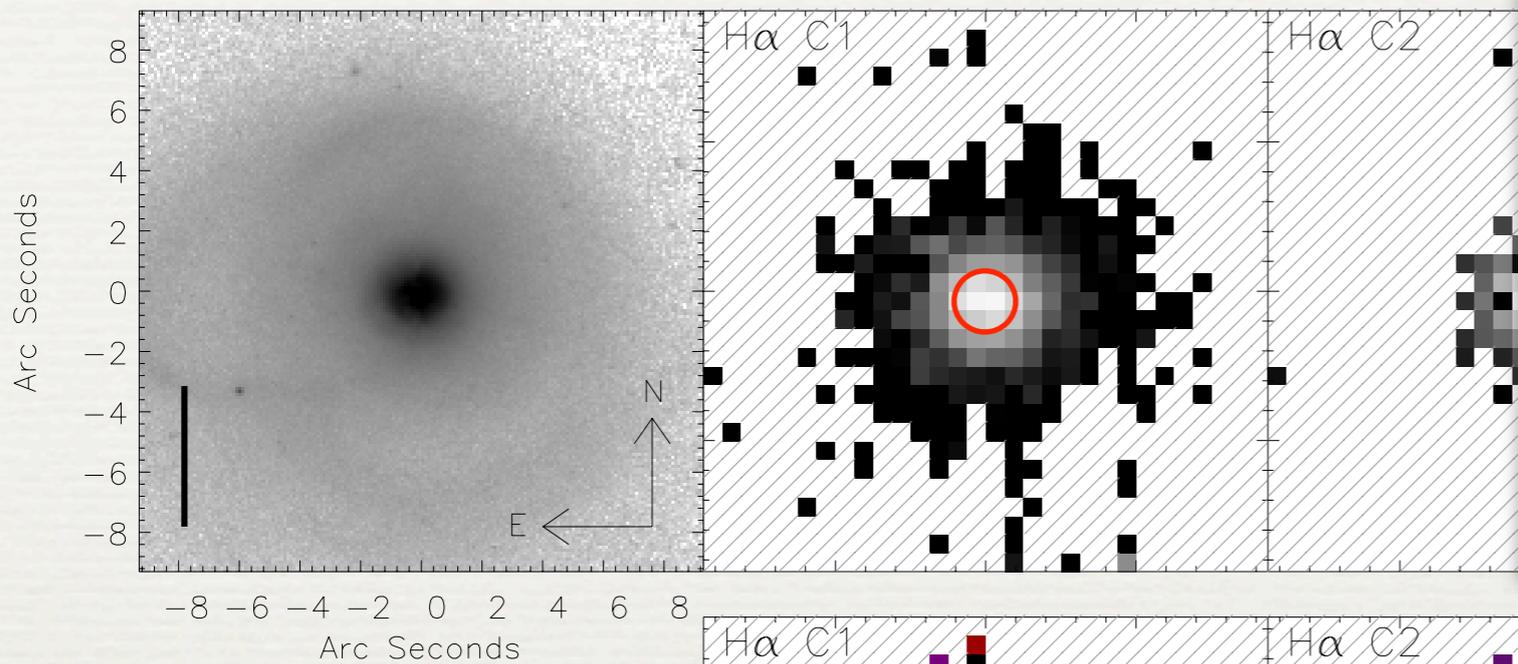
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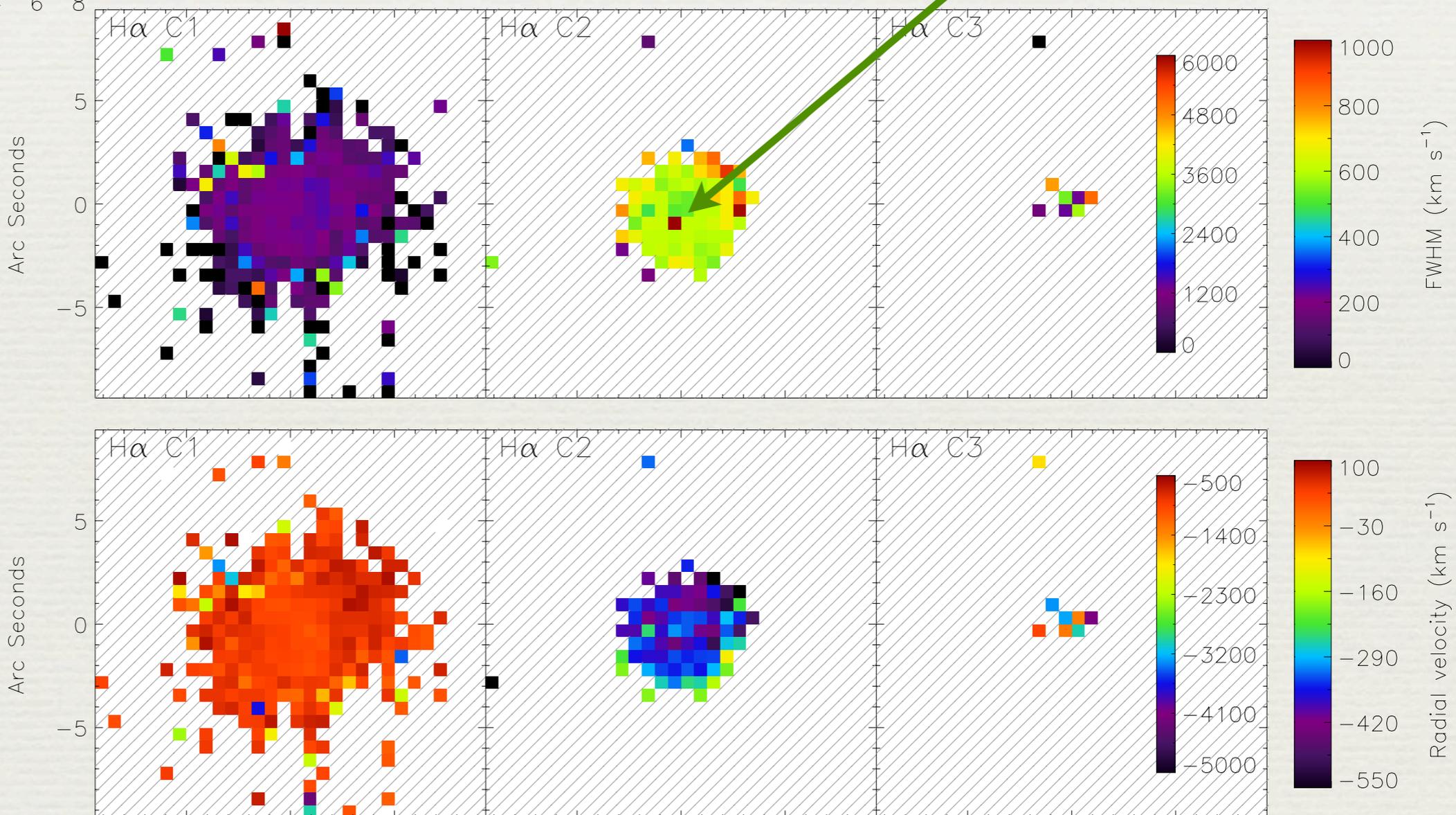
## H $\alpha$ Kinematics



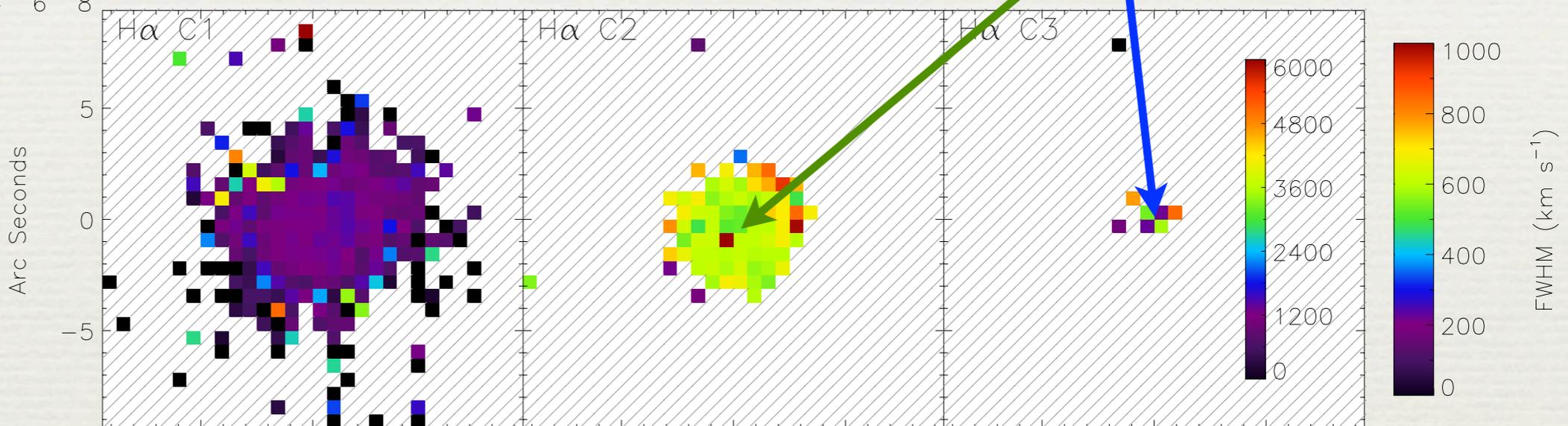
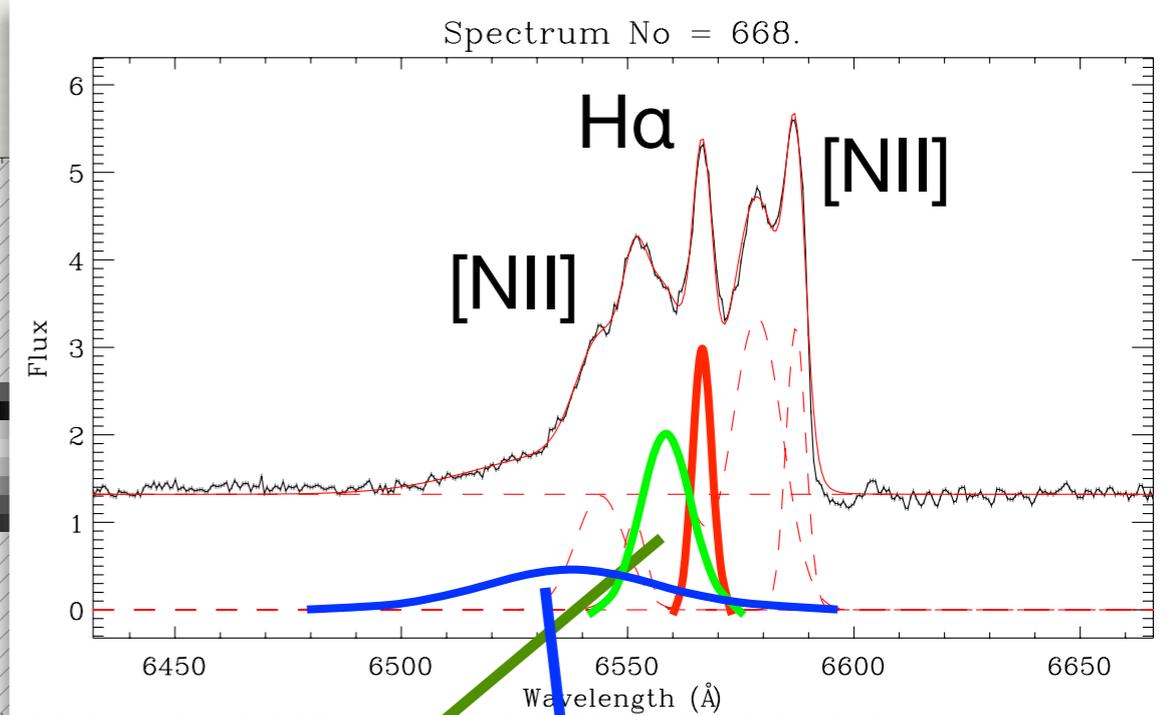
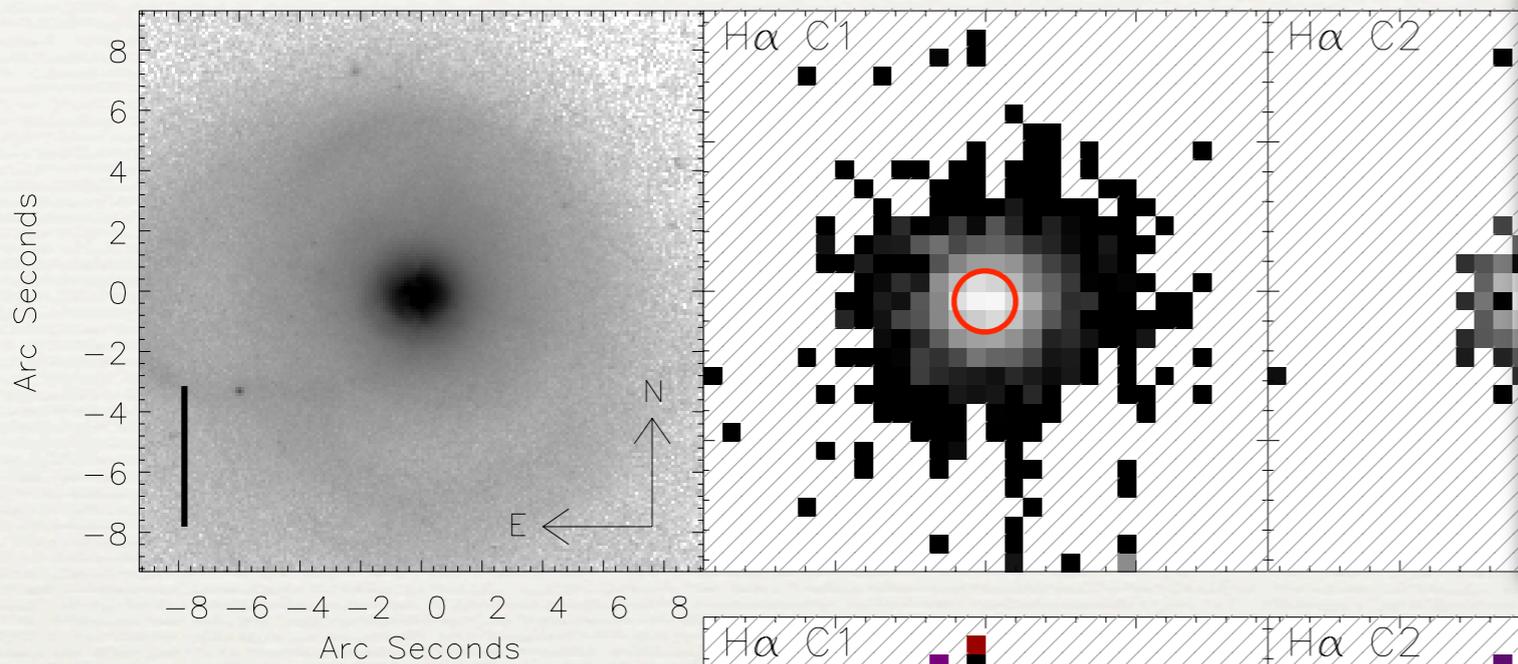
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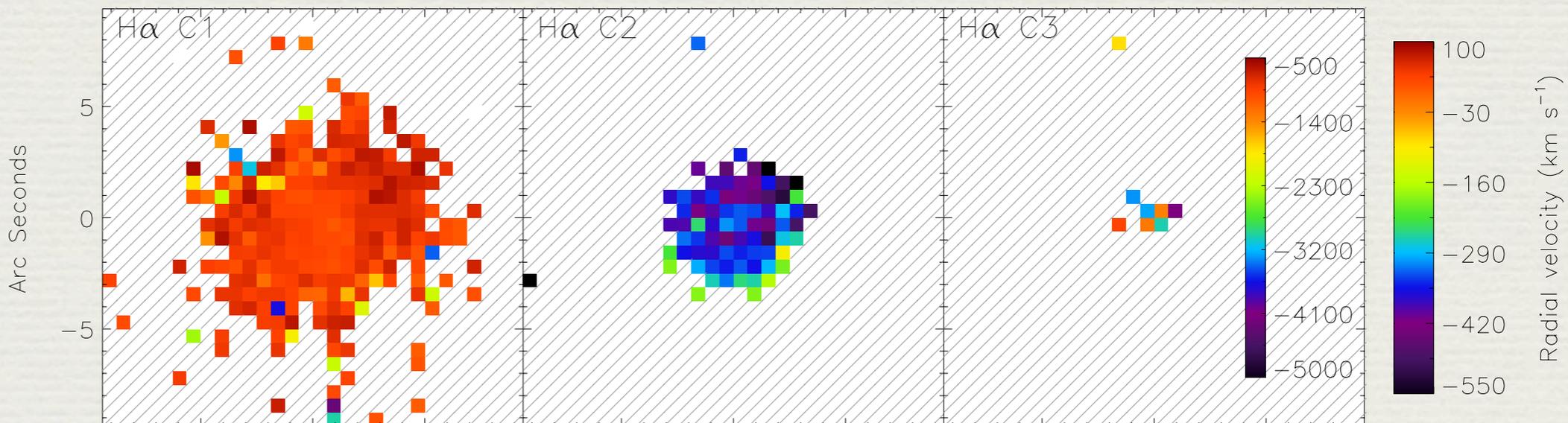
## H $\alpha$ Kinematics



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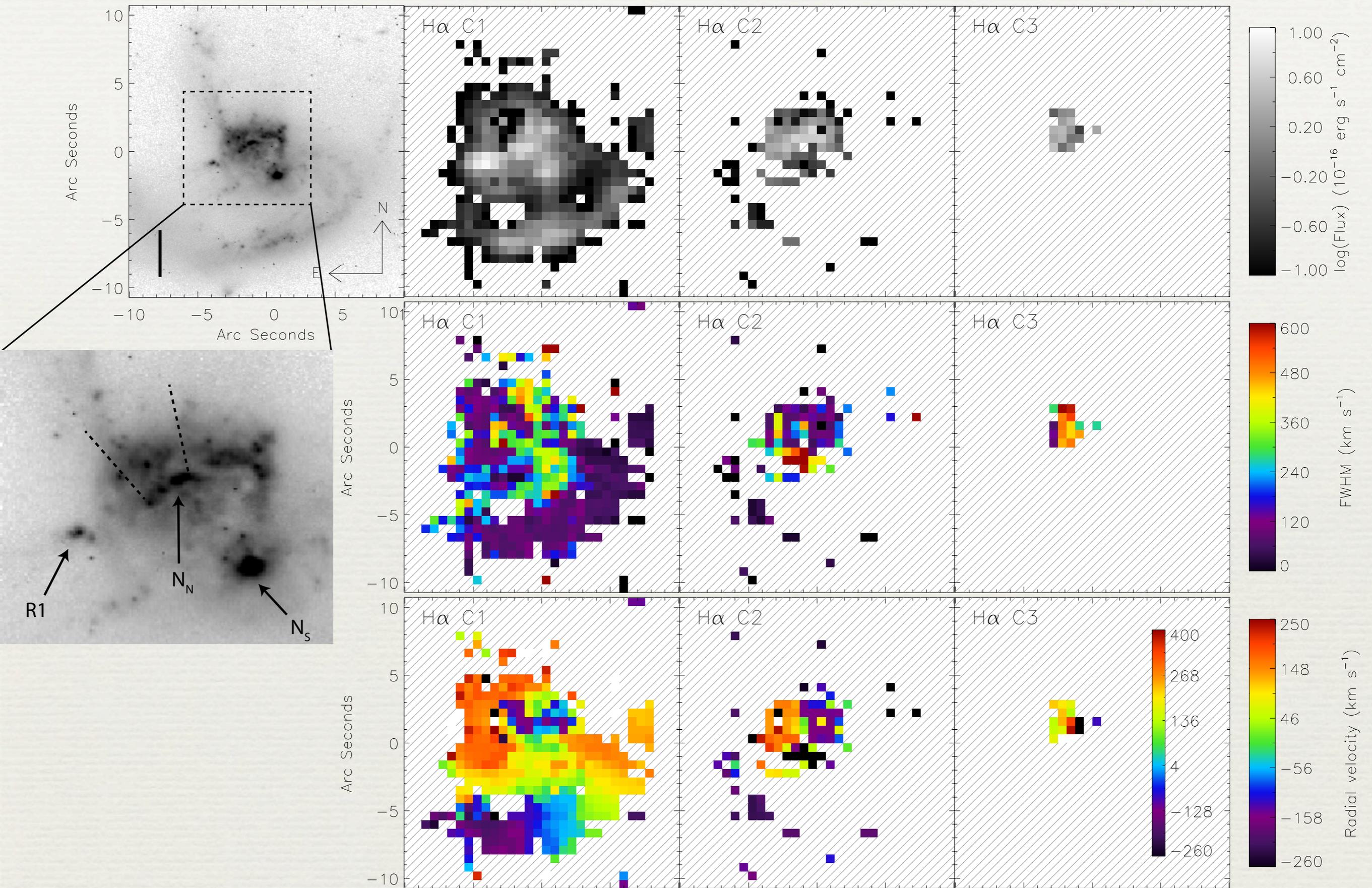


## H $\alpha$ Kinematics



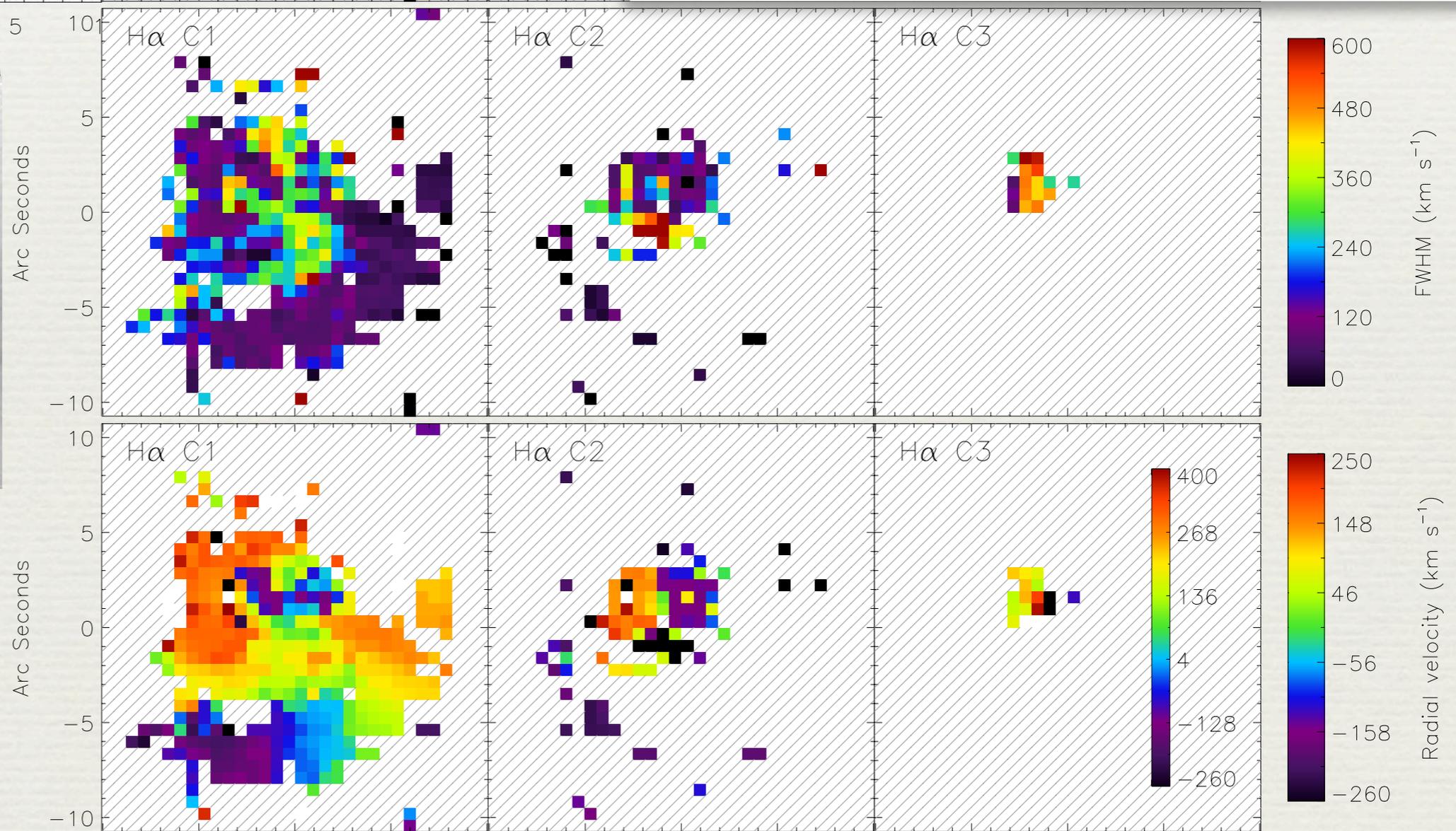
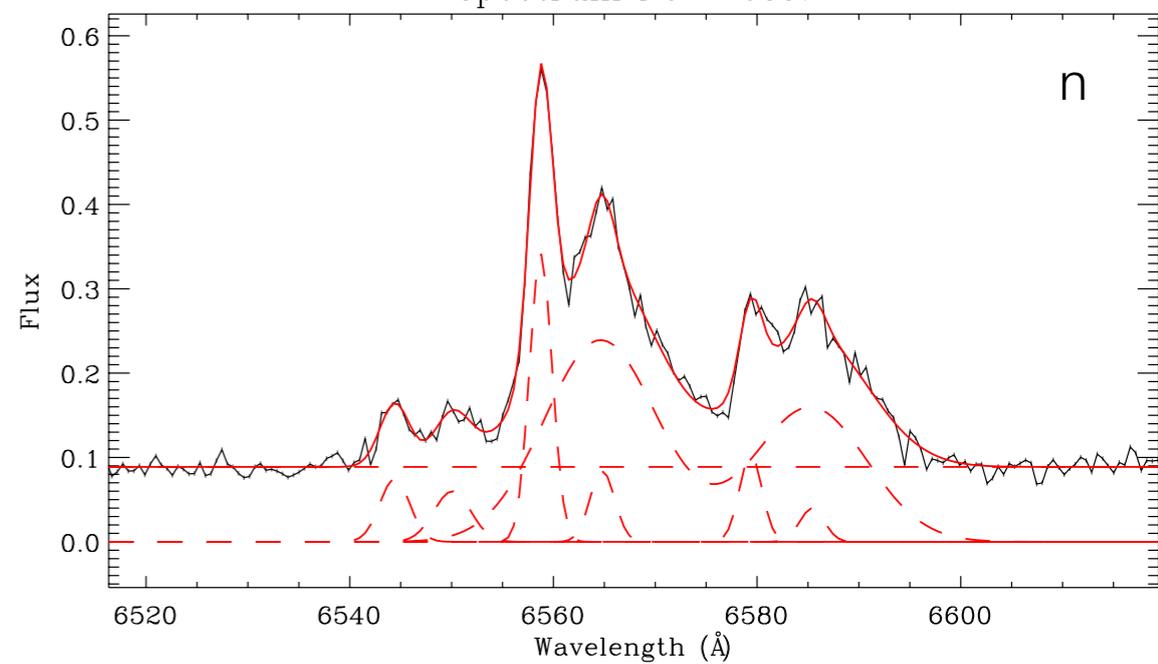
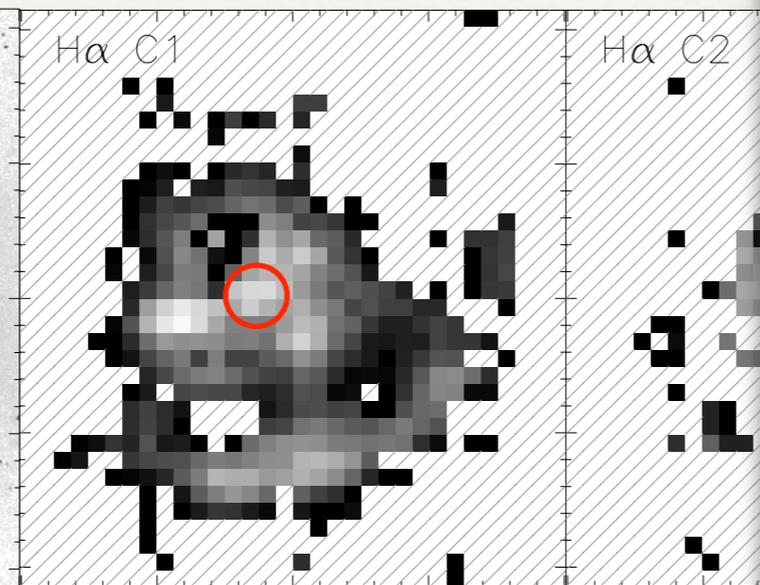
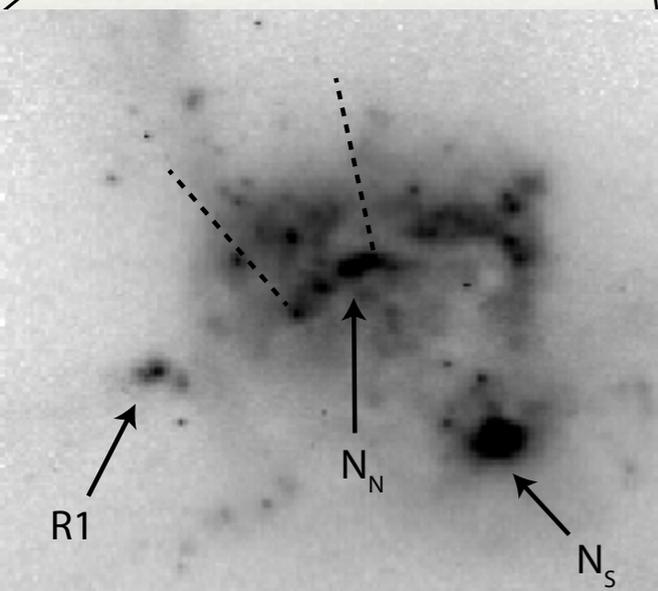
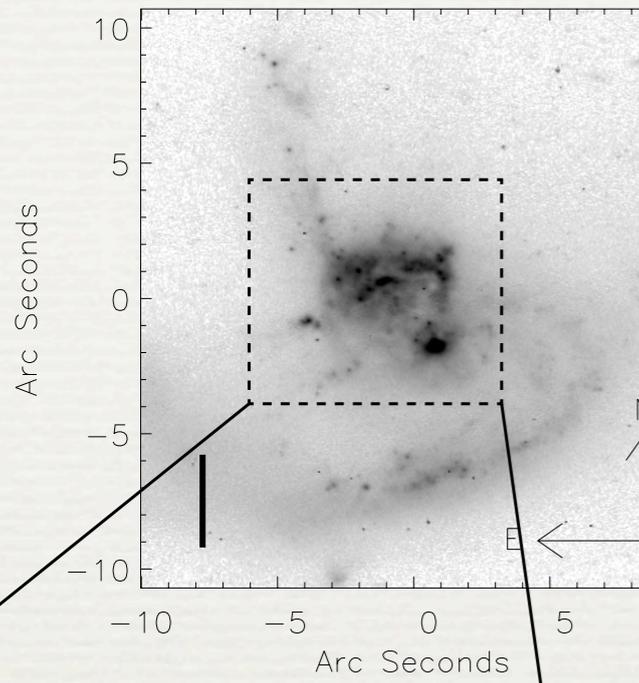
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IRAS 12112+0305



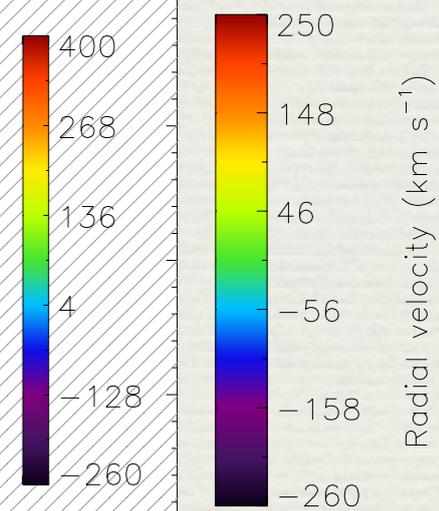
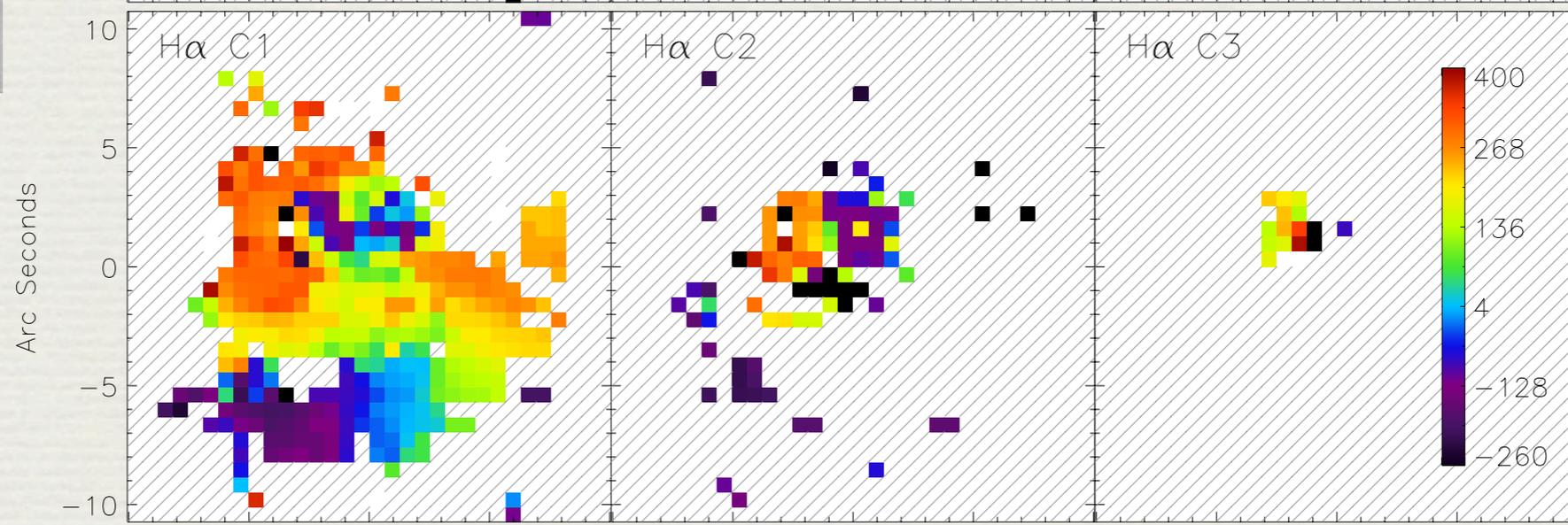
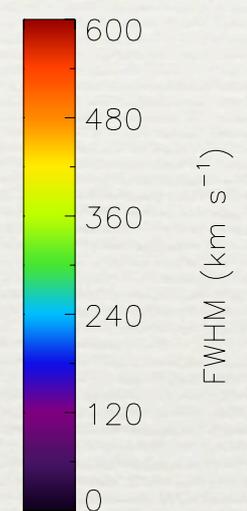
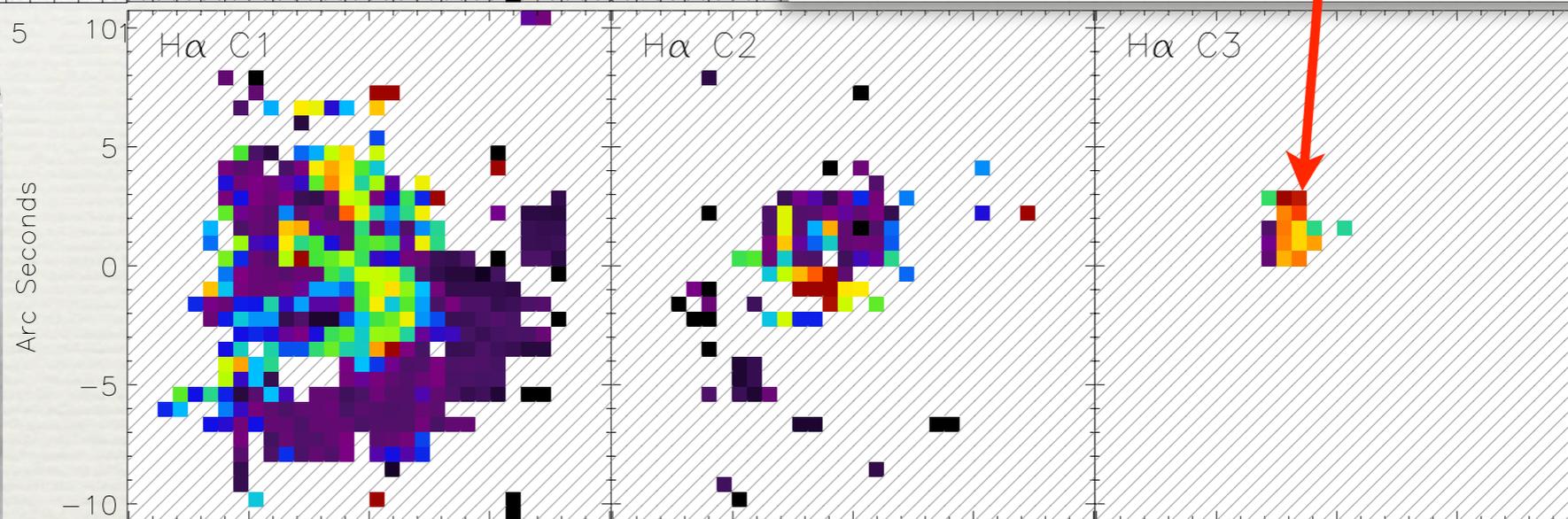
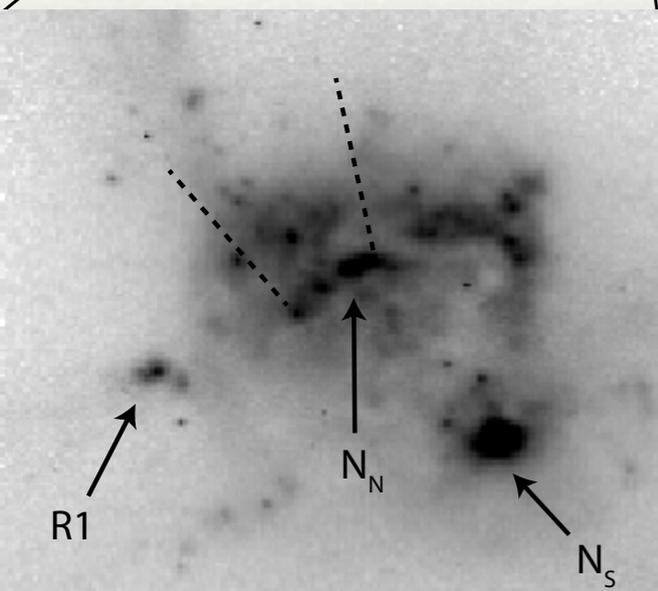
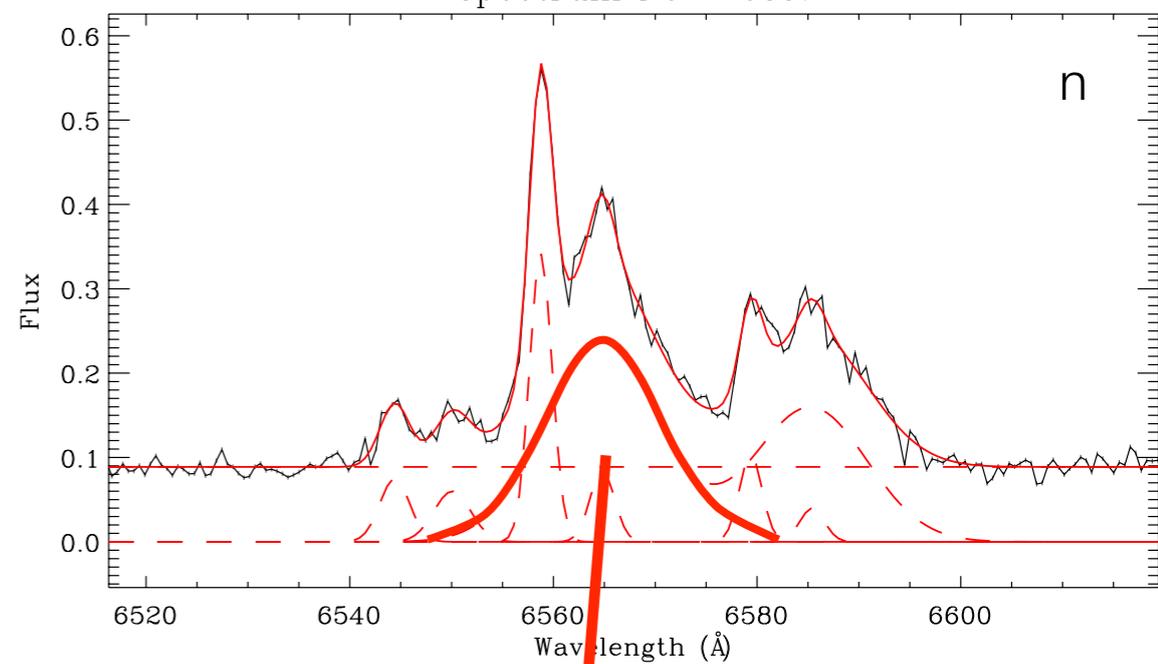
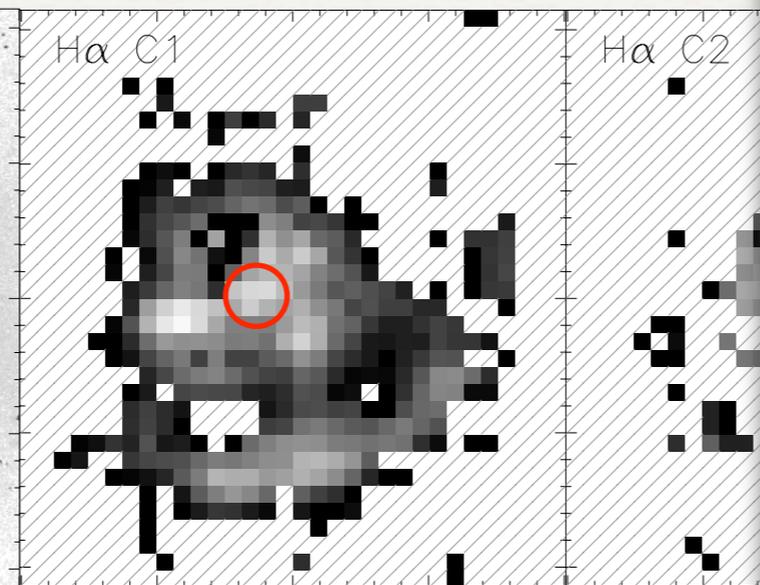
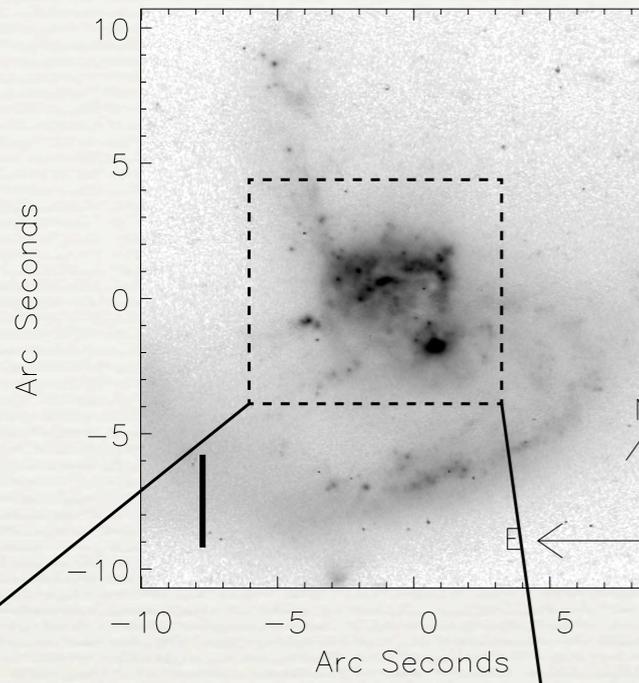
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IRAS 12112+0305

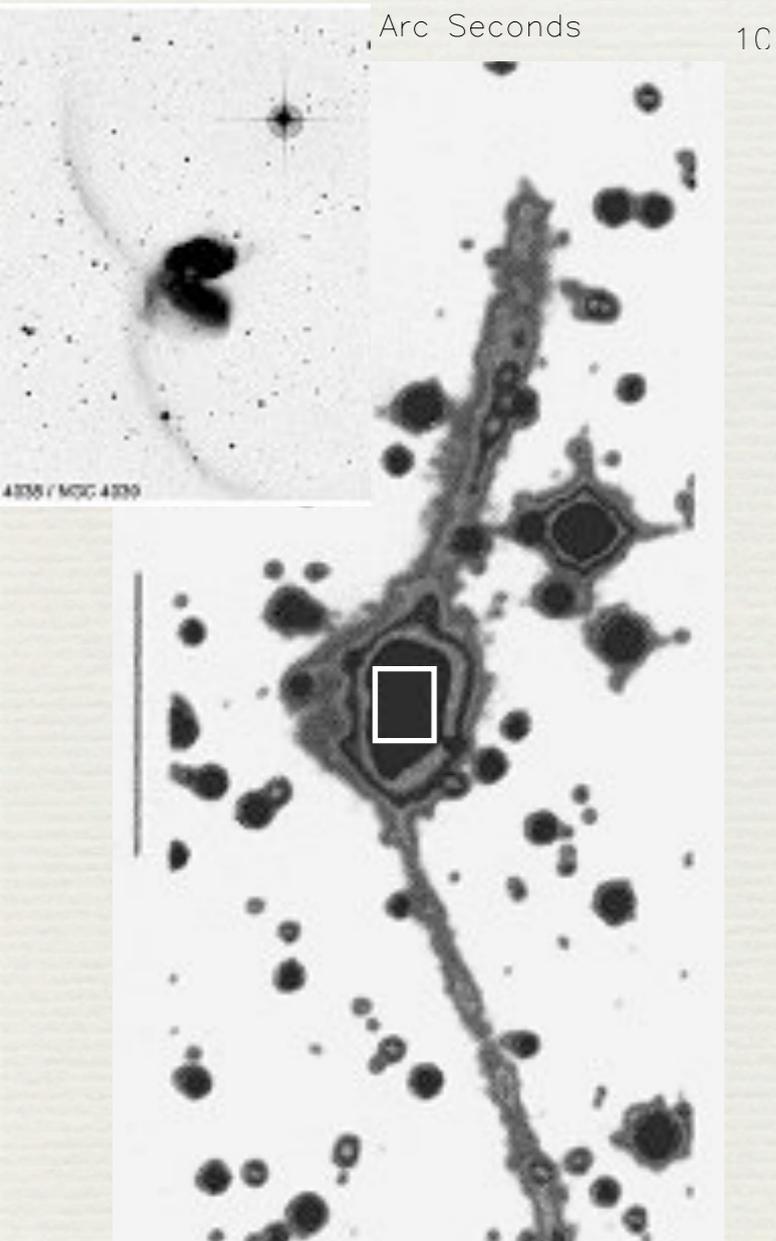
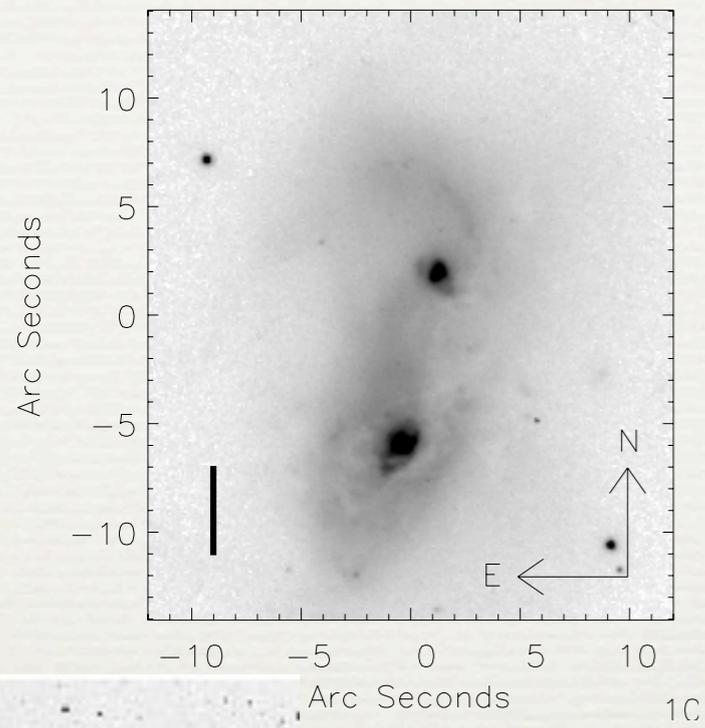


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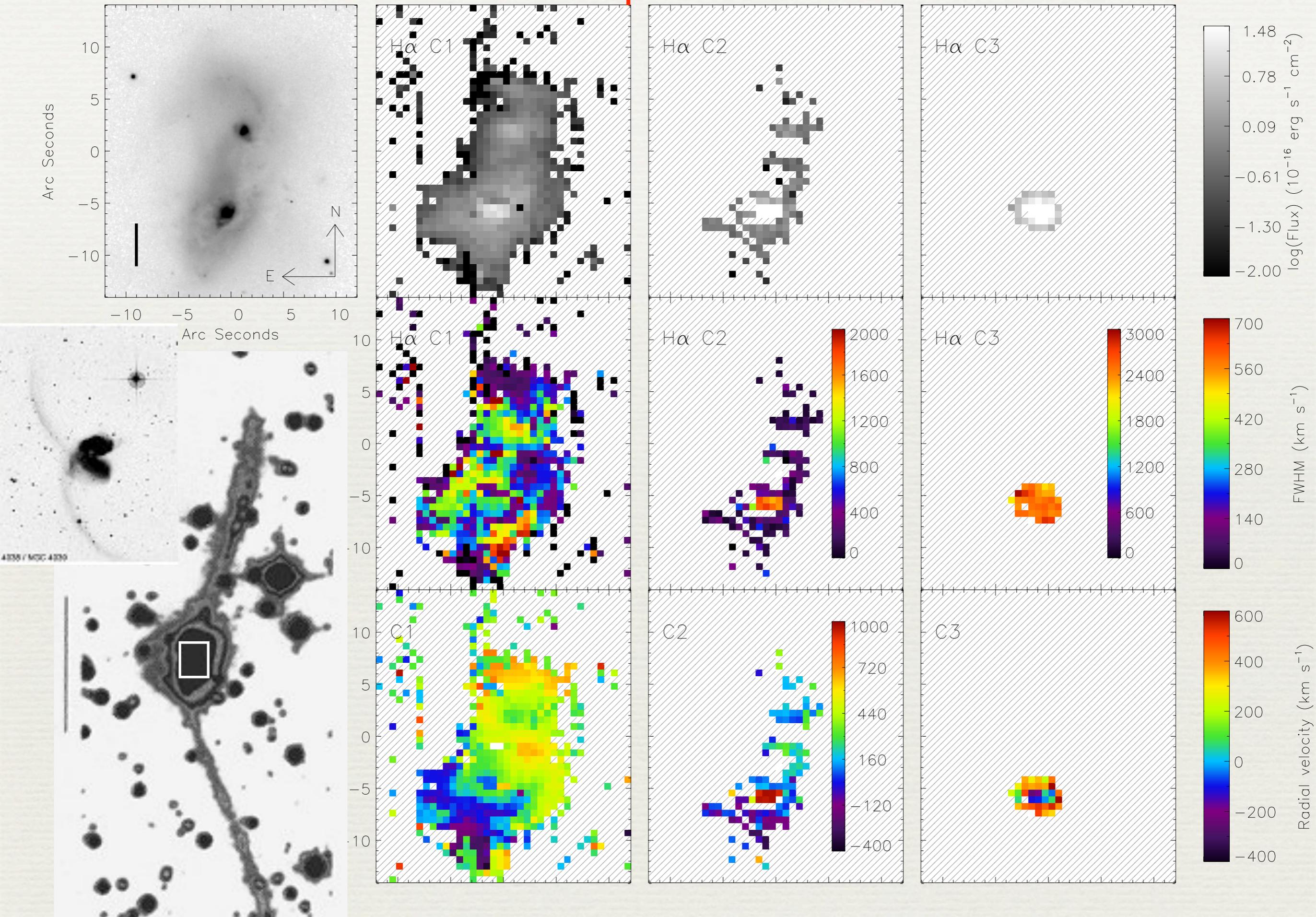
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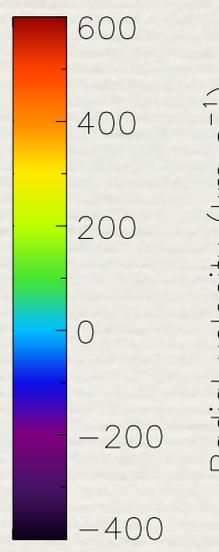
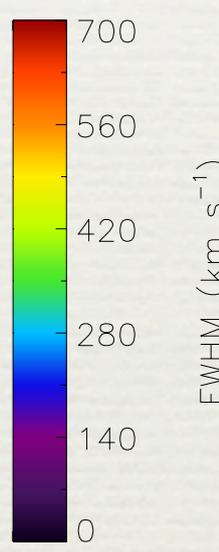
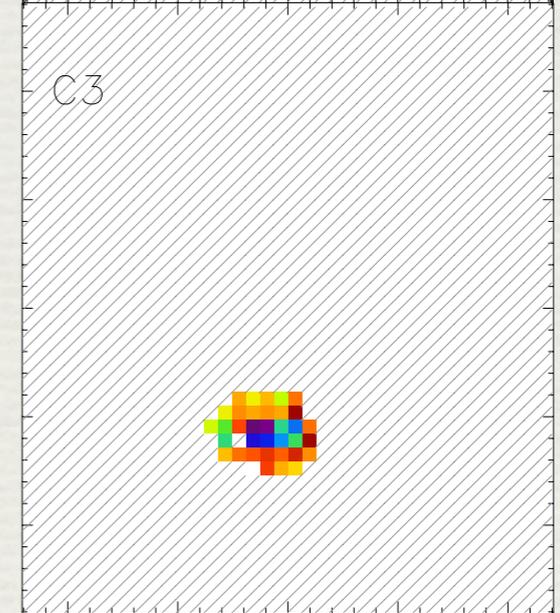
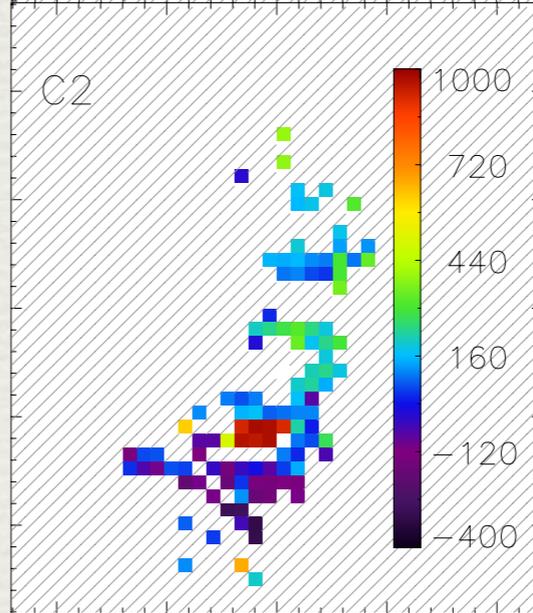
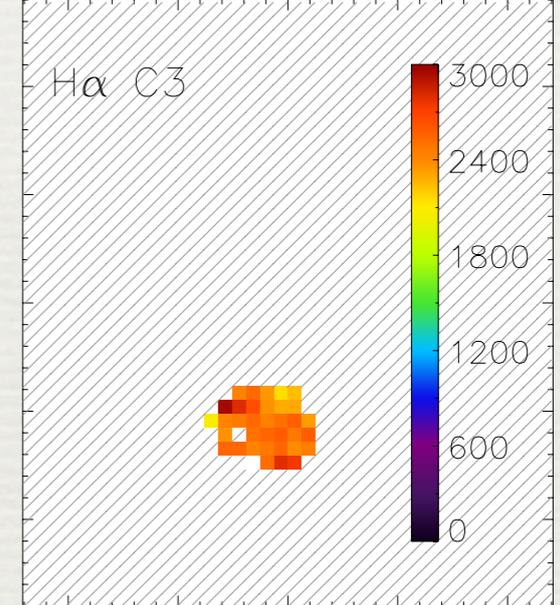
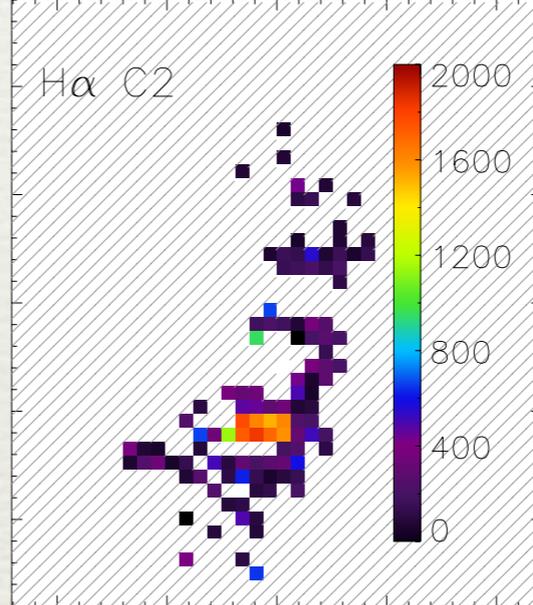
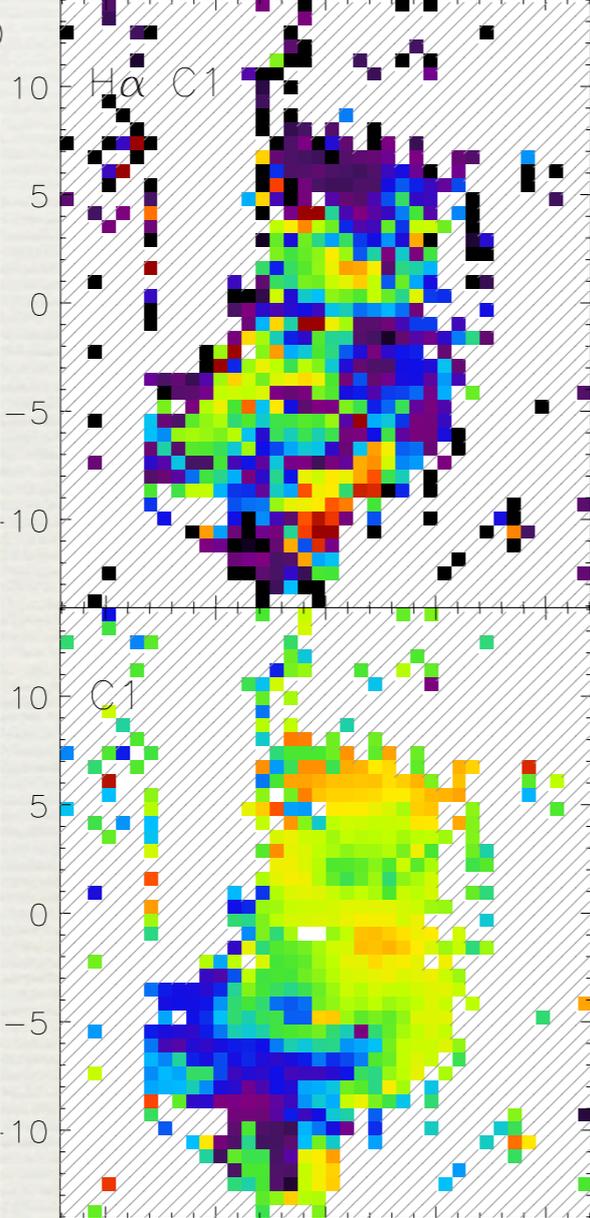
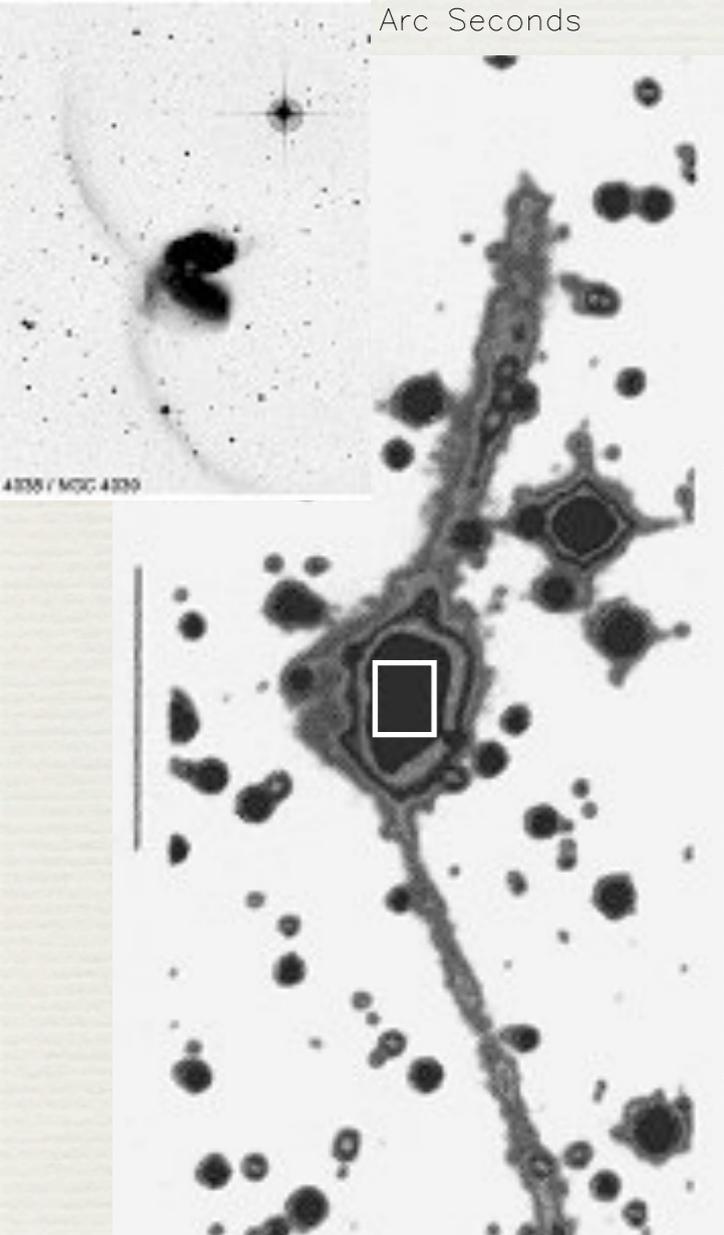
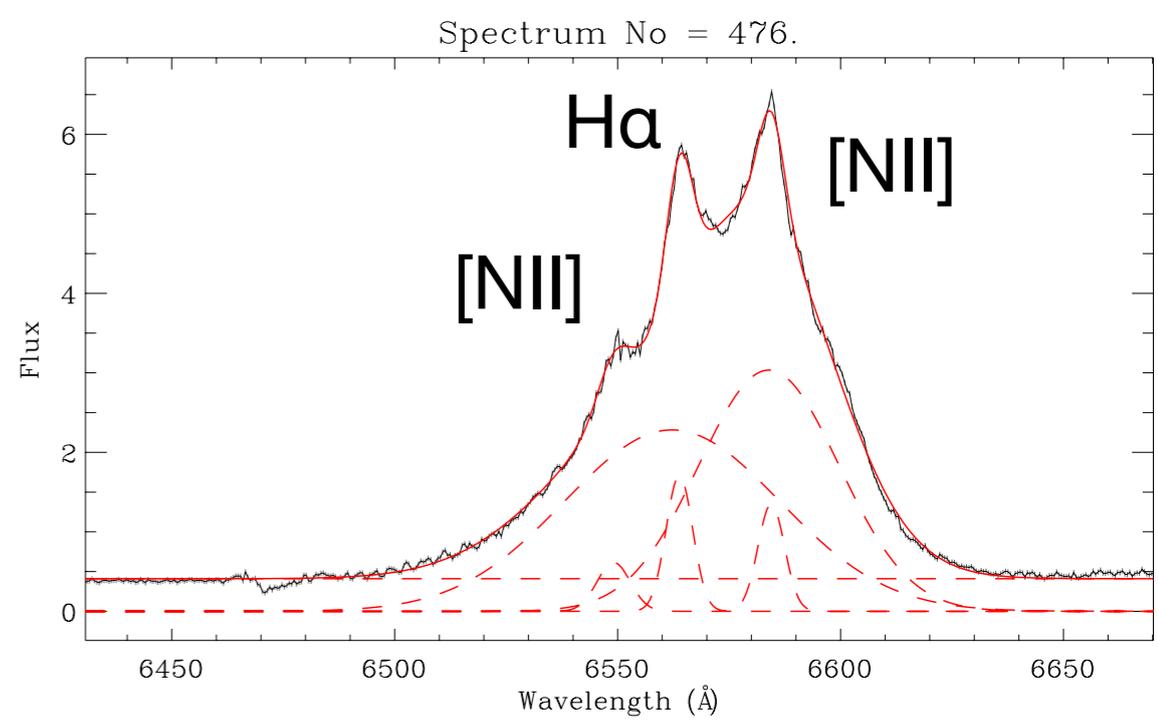
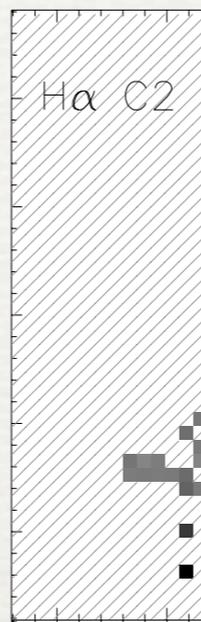
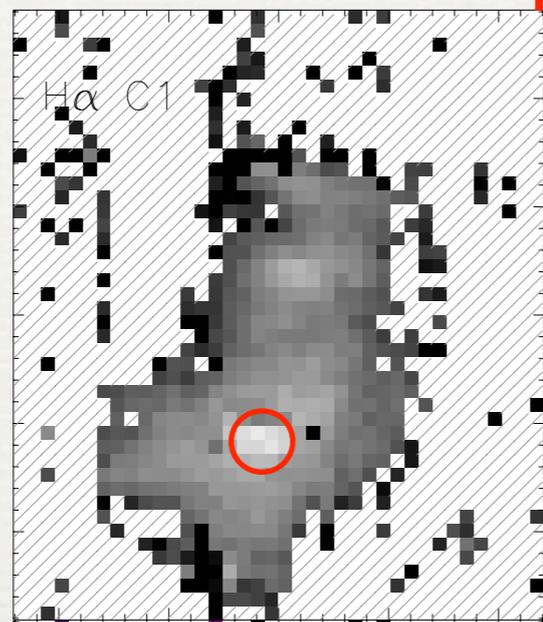
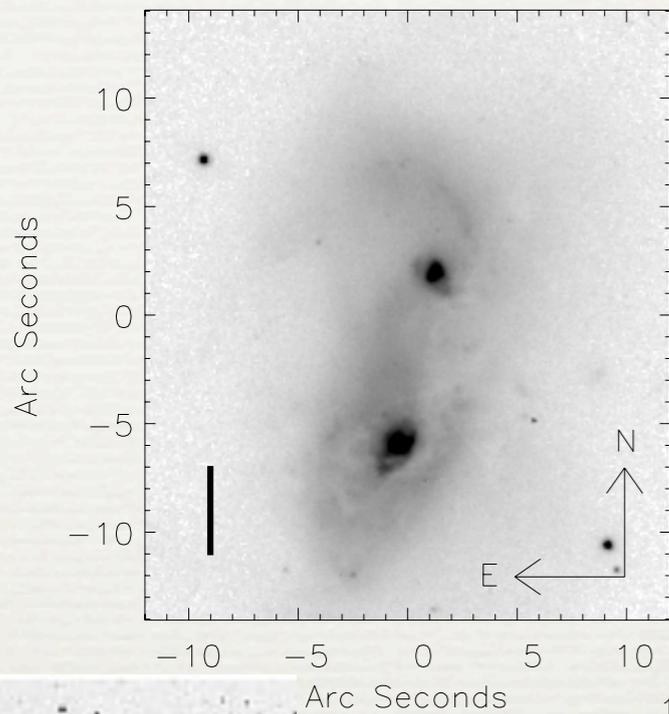
# IRAS 19254-7245: the **superantennae**



# IRAS 19254-7245: the **superantennae**

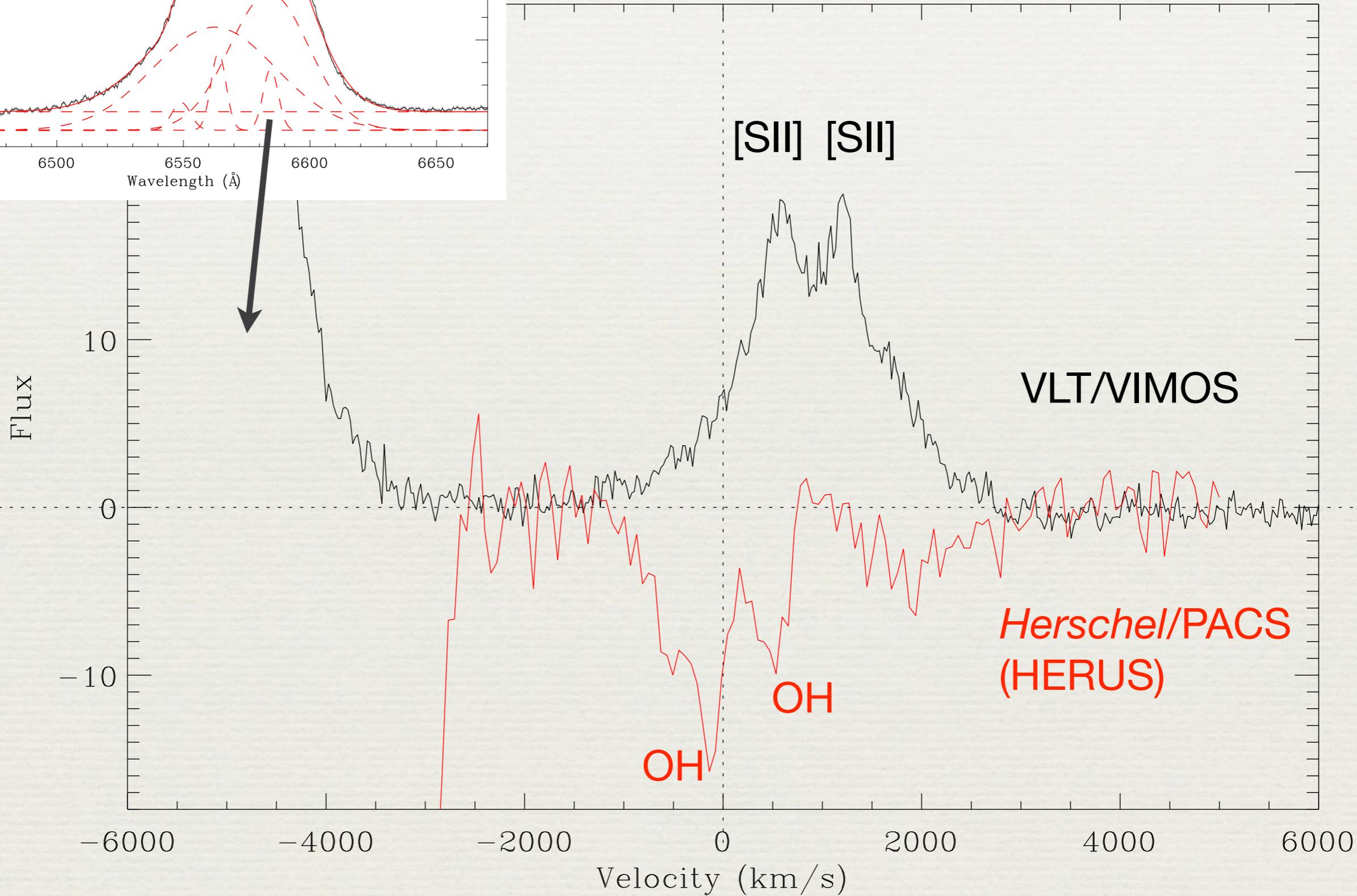
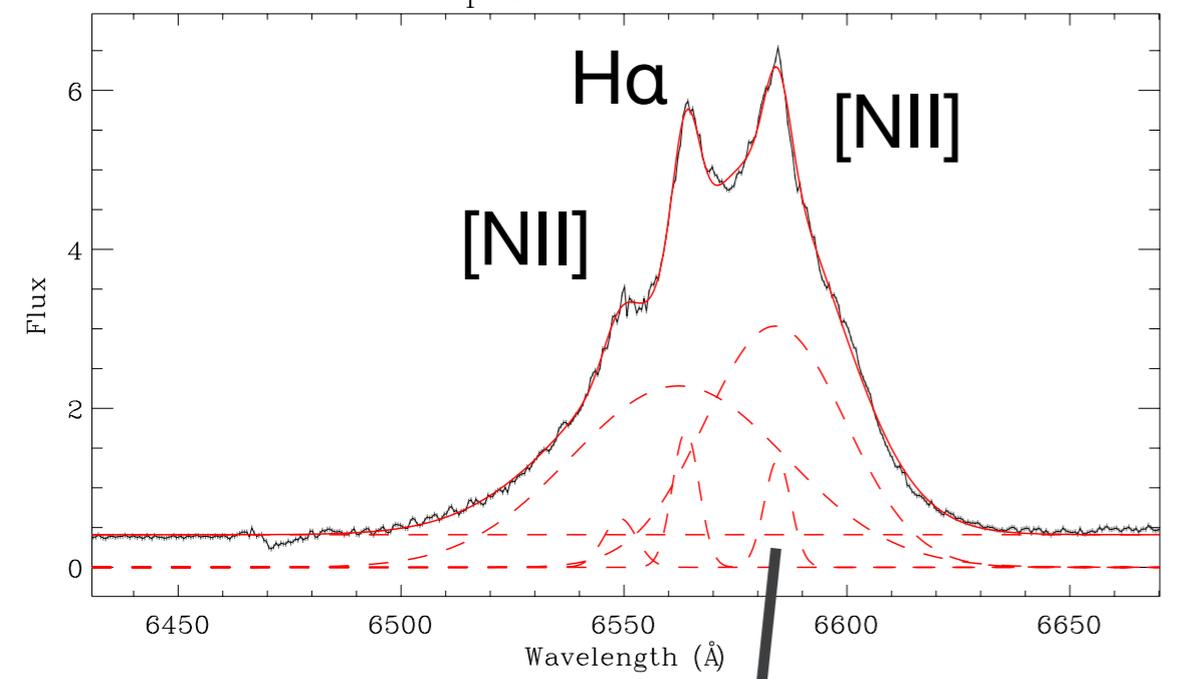


# IRAS 19254-7245: the **superan**



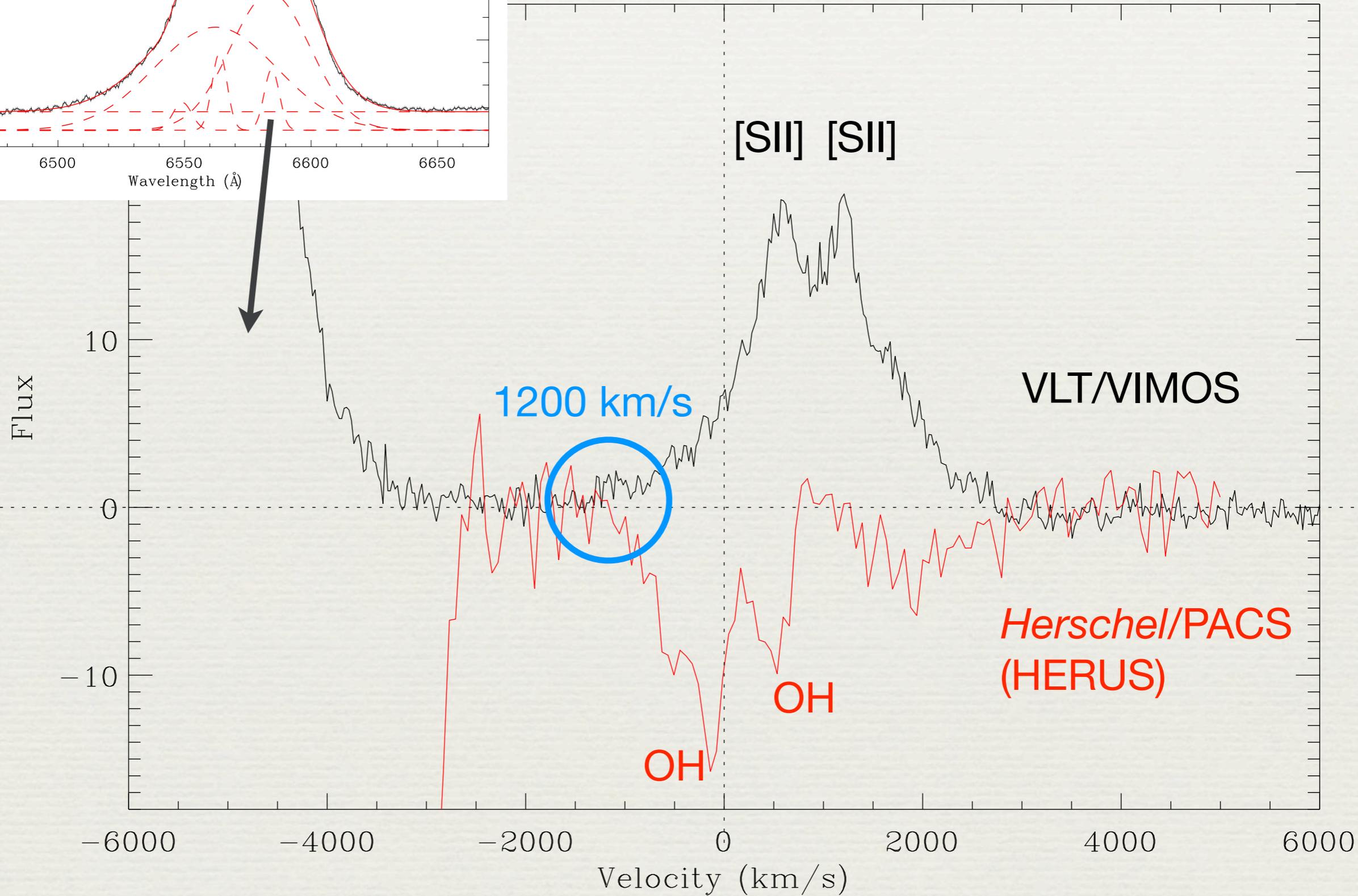
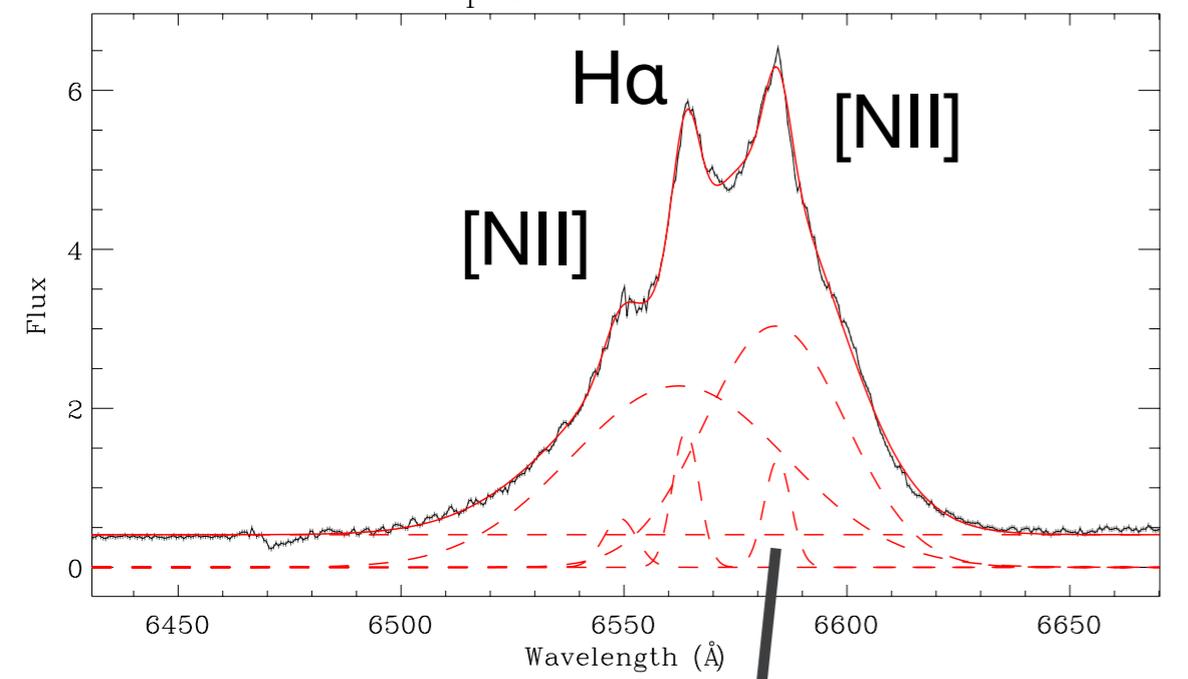
Spectrum No = 476.

IRAS 19254-7245  $N_s$

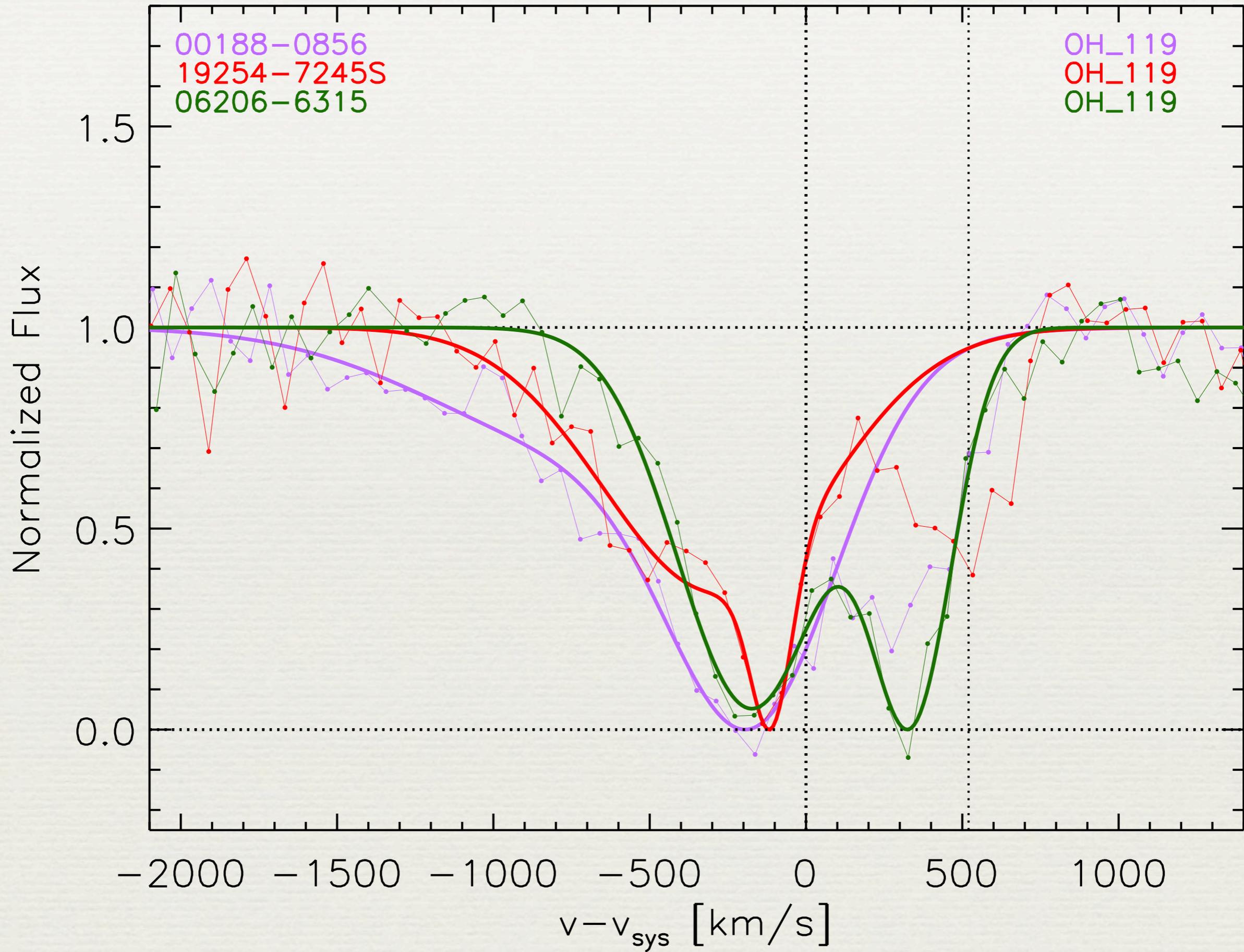


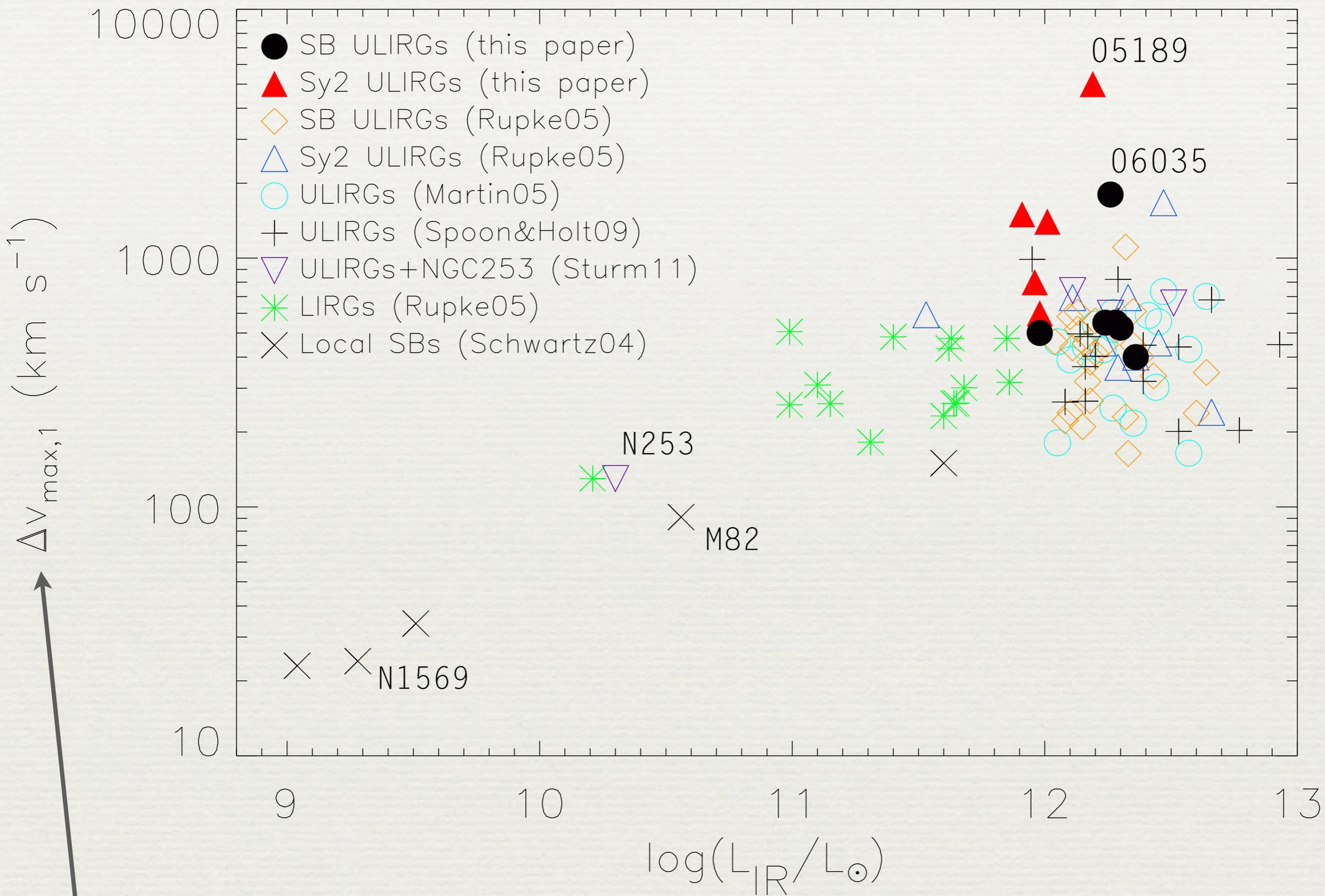
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# IRAS 19254-7245 $N_s$

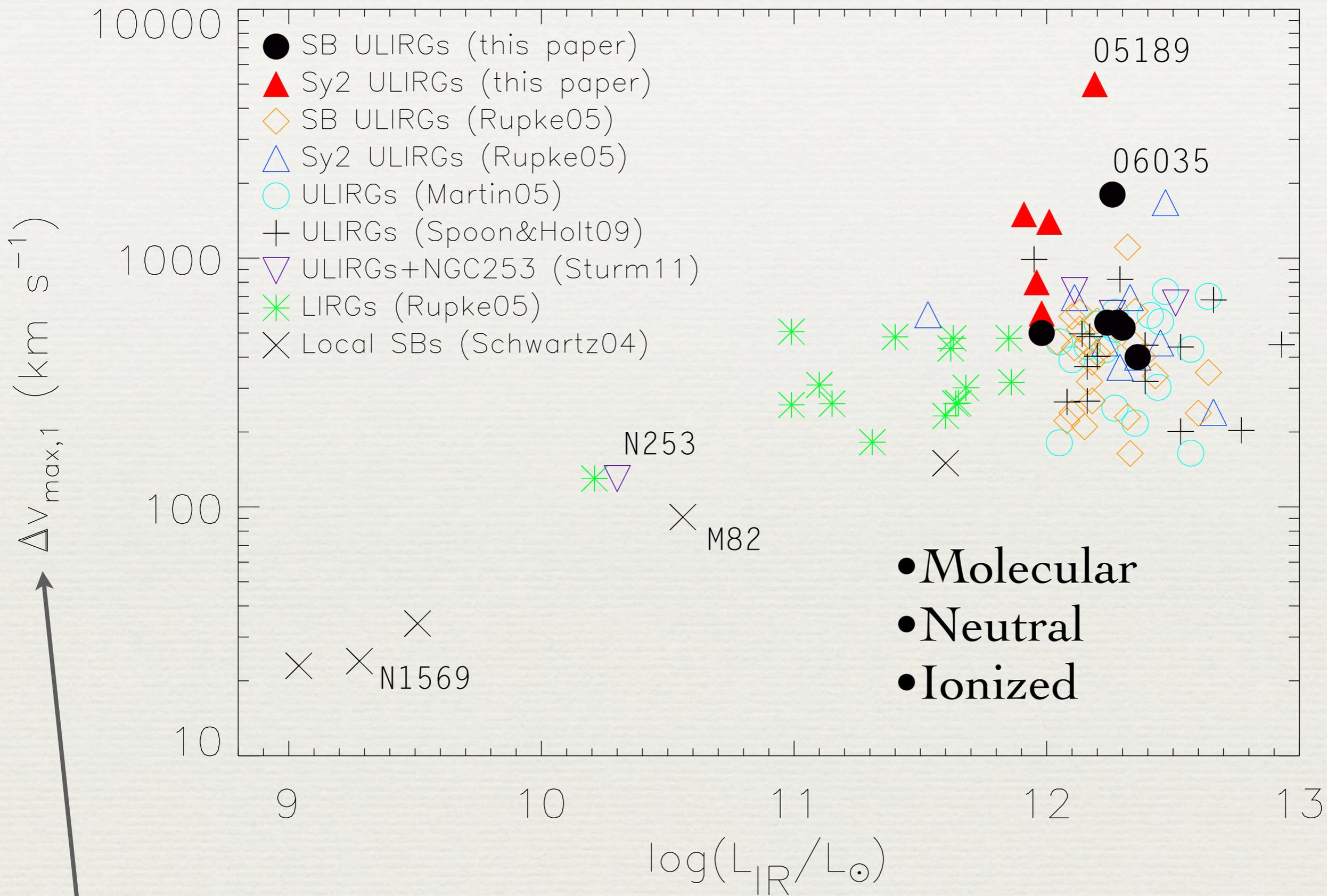


# Herschel/PACS OH 119 $\mu$ m lines





$$\Delta v_{\max,1} = \left| \langle v \rangle_{\text{broad}} - \Delta v_{\text{broad}}(\text{FWHM})/2 \right|$$



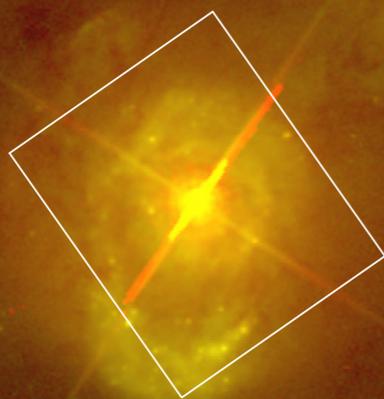
$$\Delta v_{\max,1} = \left| \langle v \rangle_{\text{broad}} - \Delta v_{\text{broad}}(\text{FWHM})/2 \right|$$

Mrk 231

HST 435W

HST 814W

5 kpc



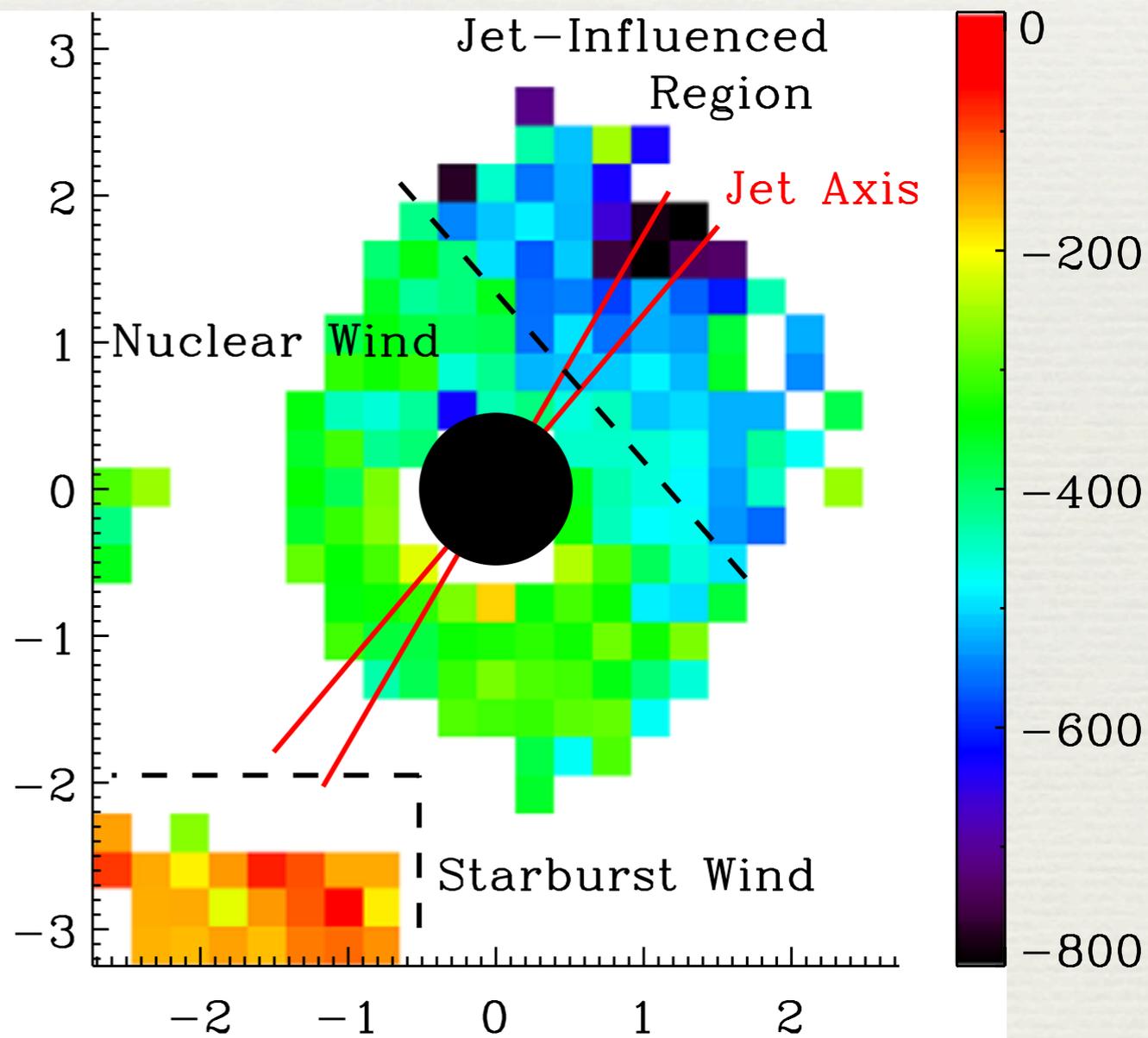
N

E

Gemini/GMOS-IFU

Rupke & Veilleux 2011

Neutral Gas  $v_{50\%}$



Thanks