Morphology of the Red Rectangle Proto-Planetary Nebula

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The morphology of the Red Rectangle (RR) proto-planetary nebula (PPN) exhibits several singular attributes. Most prominent are a series of linear features perpendicular to the symmetry axis which appear as "ladder rungs" across the nebula. At the edge of each "rung" gas seemingly flows in a parabolic shape towards the center of the nebula.

We present a new model of the RR which explains these features as a projection effect of the more-common spherically-symmetric outflows seen in other PPN (e.g. Egg Nebula). Using the 3D morpho-kinematic modeling software *shape*, we have created a model of the RR that consists of spherical shells evacuated by a biconal outflow. When the symmetry axis is oriented perpendicular to the line of sight, the spherical shells become linear thereby reproducing the "rungs" seen in the RR. When oriented at different inclinations, the linear features become spherical as observed in the Egg Nebula.

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Introduction

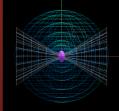




Conclusion

Model

A morphological model of the RR was developed using the SHAPE (Steffen et al 2010) software which can model radiaphotoluminescence from an organic solid and is not the result of simple dust scattering (Wada et al 2009, Cohen et al ceses are similar. The model consists of 4 parts (figure 3):



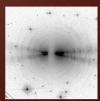
1) A central source supplying the photons which are scat-

walls of the bi-cone are dense and represent material that has been swept up by the evacuating process. The walls also scatter light









Results

The bi-cones cut off the "tips" of the spherical shells and thus produce a sequence of evacuated cones. Photons luminate the ring around the cone base. When the nebula is perpendicular to the line of sight, the base of the

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