

$$
v_{C M}=40 \mathrm{~km} \mathrm{~s}^{-1}
$$

$P=17.5$ days

- If you assume $i=90^{\circ}$ the masses are:

$$
\begin{array}{ll}
M_{A}+M_{B}=\frac{P}{2 \pi G}\left(\frac{v_{r A}+v_{r B}}{\sin i}\right)^{3} & \frac{M_{B}}{M_{A}}=\frac{v_{r A}}{v_{r B}}=\frac{25}{70} \\
=\frac{17 \times 24 \times 3600}{2 \pi 6.7 .10^{-11}}\left(\frac{(25+70) \times 10^{3}}{\sin 90^{\circ}}\right)^{3} & M_{A}=2.8 M_{B} \\
3.8 M_{B}=1.4 \mathrm{M}_{S u} \\
=3.0 \times 10^{30} \mathrm{~kg} & M_{B}=0.37 \mathrm{M}_{\text {Sun }} \\
=1.5 \mathrm{M}_{\text {Sun }} & M_{A}=1.0 \mathrm{M}_{\text {Sun }}
\end{array}
$$

## Galaxy Types

- Definition of a galaxy
- Galaxy Types
- Spiral
- Elliptical
- Irregular
- Galaxy Classification
- Hubble's Tuning Fork Diagram


## Galaxies

- A galaxy is a gravitationally bound collection of stars, gas and dust
- e.g. our Galaxy contains of order $10^{11}$ stars
- Usually isolated in space, although can interact with near neighbours
- The main visible component of the Universe


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## Galaxy Types

- Galaxies are seen in three major types
- Spirals
- Ellipticals
- Irregulars


## Spiral Galaxies

- Rotating disc dominated by spiral arms
- Spiral Arms are
- rich in young, hot, blue stars, i.e. Population I
- rich in gas and dust
- where formation of new stars takes place


Credit: Gemini Observatory, GMOS Team

- An elliptical concentration of stars at the centre is called the bulge
- Bulge is rich in red stars - Population II and old Population I

- Also come in barred form where the two arms originate from the ends of a central linear feature of bulge-like stars



## Elliptical Galaxies

- Elliptical collections of red stars - Population II and old Population I
- Smooth variation in intensity
- Very little gas \& dust
- Little organized rotation
- Come in both giant and dwarf forms



Leo I
Dwarf Elliptical
From Universe textbook

## Irregular Galaxies

- No regular structure
- Contain plenty of gas and dust and blue stars
- Mixture of Population I and II
- Usually relatively small



Irregular galaxy NGC 4214. Credit NASA HST

## Class Example

- What is the typical angular separation of stars that are about 1 pc apart in a galaxy that is at a distance of 1 Mpc ?

$$
\begin{aligned}
& \theta=\frac{l}{d}=\frac{1}{1 \times 10^{6}}=1 \times 10^{-6} \text { radians } \\
& =206265 \times 1 \times 10^{-6}=0.2^{\prime \prime}
\end{aligned}
$$

## Galaxy Classification

- Galaxies are classified according to Hubble's tuning fork diagram


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## Classification of Ellipticals

- Ellipticals (E) are classified according to their apparent ellipticity by the nearest integer to

$$
10 \frac{(a-b)}{a}
$$

where $a$ is the major axis and $b$ the minor axis of the ellipse

- This is the apparent ellipticity, i.e. it is affected by the viewing angle

(b) E3 (NGC 4365)

(c) E6 (NGC 3377)

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## Classification of Spirals

- Spirals (S) are classified by
- the relative size of the bulge,
- the amount of dust and gas,
- the tightness of the spiral arms
- Sa galaxies have a large bulge, small disc and tight spiral arms and little gas \& dust
- Sc galaxies have a small bulge, large disc and loose spiral arms and lots gas \& dust
- Sb are intermediate

(b) Sb (M81)
(c) Sc (NGC 4321)

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- Barred spirals (SB) have an equivalent classification going from SBa to SBc with the bulge and bar being dominant in class 'a' and progressively less so towards class 'c'


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## S0 Galaxies

- S0 galaxies are intermediate between the most elliptical ellipticals (E7) and the spirals Sa or SBa.
- Has a disc and a bulge as well as an elliptical component
- A little gas but no obvious spiral arms
- Also called lenticular galaxies


## Galaxy NGC 2787



## Summary

- Galaxies can be classified as either spirals, ellipticals or irregulars
- Hubble's tuning fork diagram is a convenient memory aid but is not an evolutionary sequence
- Spirals and irregulars contain gas, dust and blue stars whilst ellipticals contain only red stars


## Class Example

- What is the physical diameter of the Andromeda galaxy that has an angular diameter of $3^{\circ}$ and a distance of 0.78 Mpc ?


